

Shear viscosity and resonance lifetimes in the hadron gas

presented by Jean-Bernard Rose

with D. Oliinychenko, J. Torres-Rincon, A. Schäfer, H. Petersen

based on arXiv:1709.03826 and arXiv:1709.00369



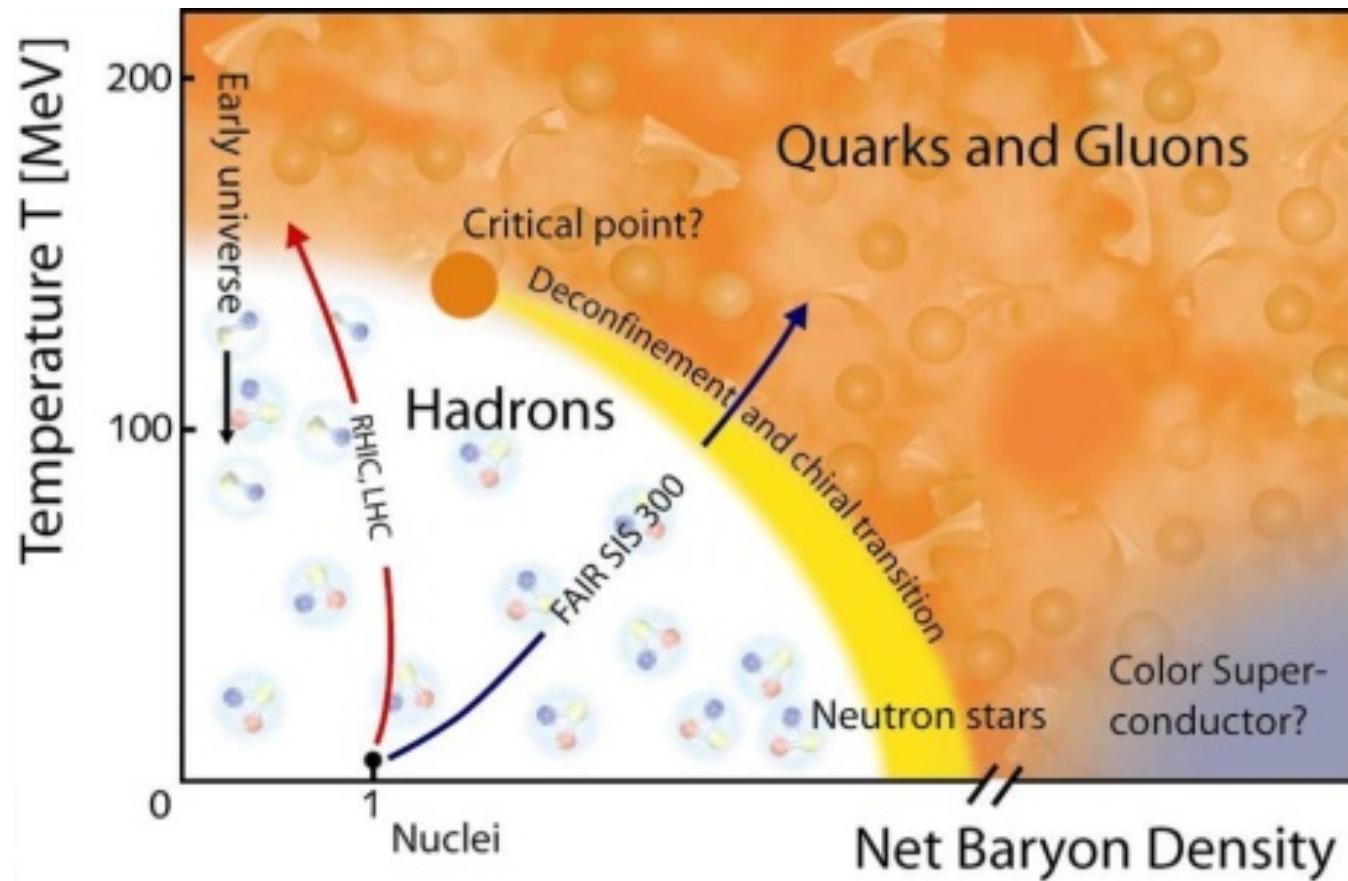
Outline

1. Introduction: Viscosity of the hadron gas
2. Transport
 - SMASH
3. Methodology
 - Viscosity considerations
 - Green-Kubo formalism
 - Test case #1: Constant isotropic cross-section
 - Test case #2: Energy-dependent cross-section
 - Entropy considerations
4. Results
 - Full hadron gas viscosity
 - Comparison & discussion
5. Conclusion

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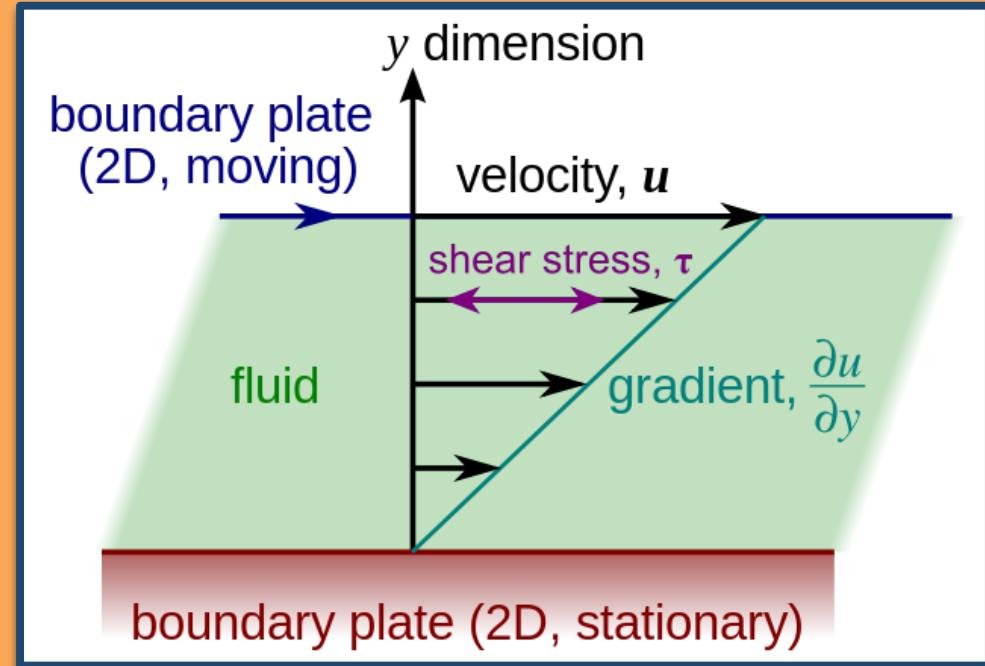
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What is the hadron gas?



What is viscosity?

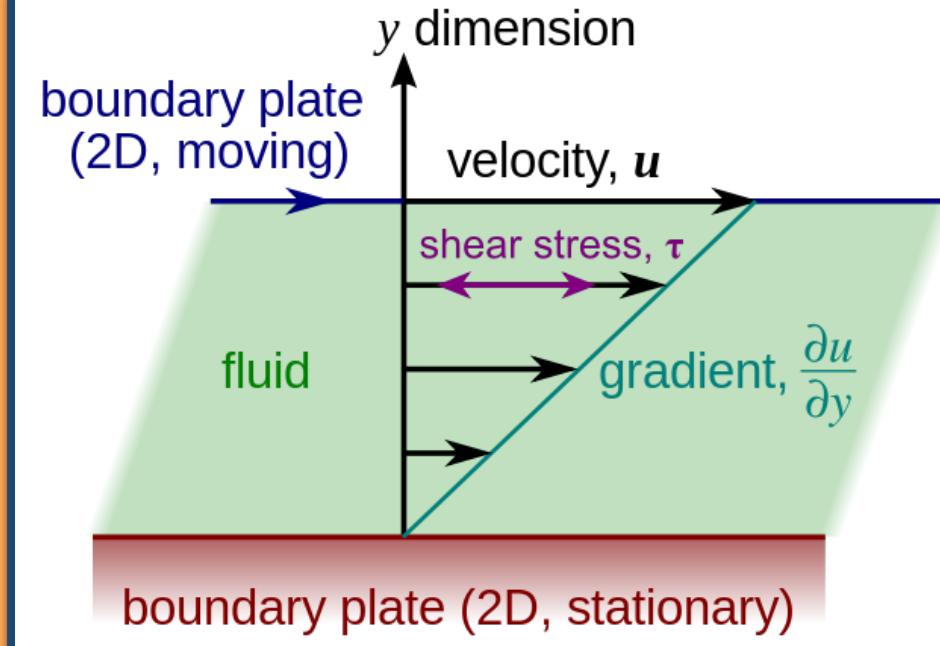
Viscosity is a measure of the friction between layers of a fluid



Wikipedia-Viscosity

...and why do we need it?

Viscosity is a measure of the friction between layers of a fluid



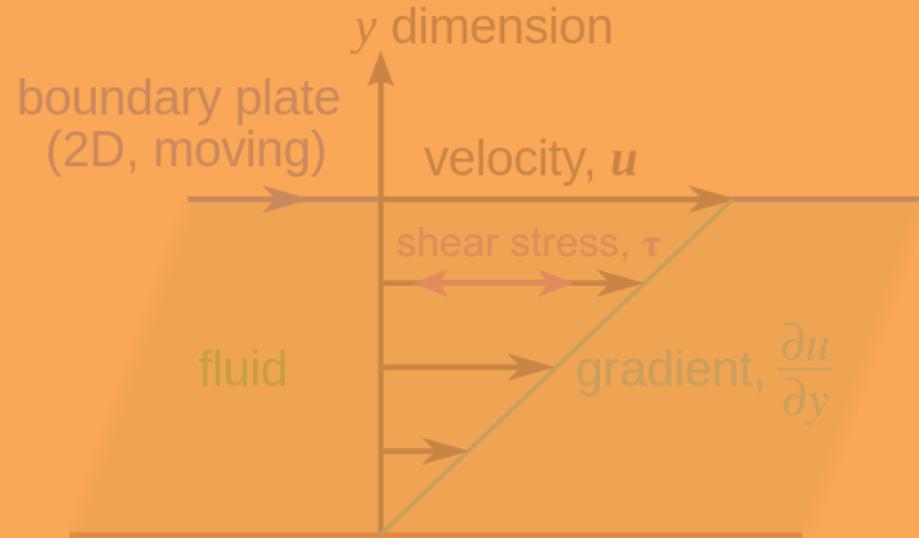
Wikipedia-Viscosity

- Hydrodynamics is conservation laws:
$$\partial_\mu T^{\mu\nu} = 0, \quad \partial_\mu N^\mu = 0$$
- With 1st order dissipative corrections (Navier-Stokes):

$$T^{\mu\nu} = e u^\mu u^\nu - (p + \zeta \theta) \Delta^{\mu\nu} + 2\eta \sigma^{\mu\nu}, \quad N^\mu = n u^\mu + \kappa \partial^\mu \frac{u}{T}$$

...and why do we need it?

Viscosity is a measure of the friction between layers of a fluid



Bulk

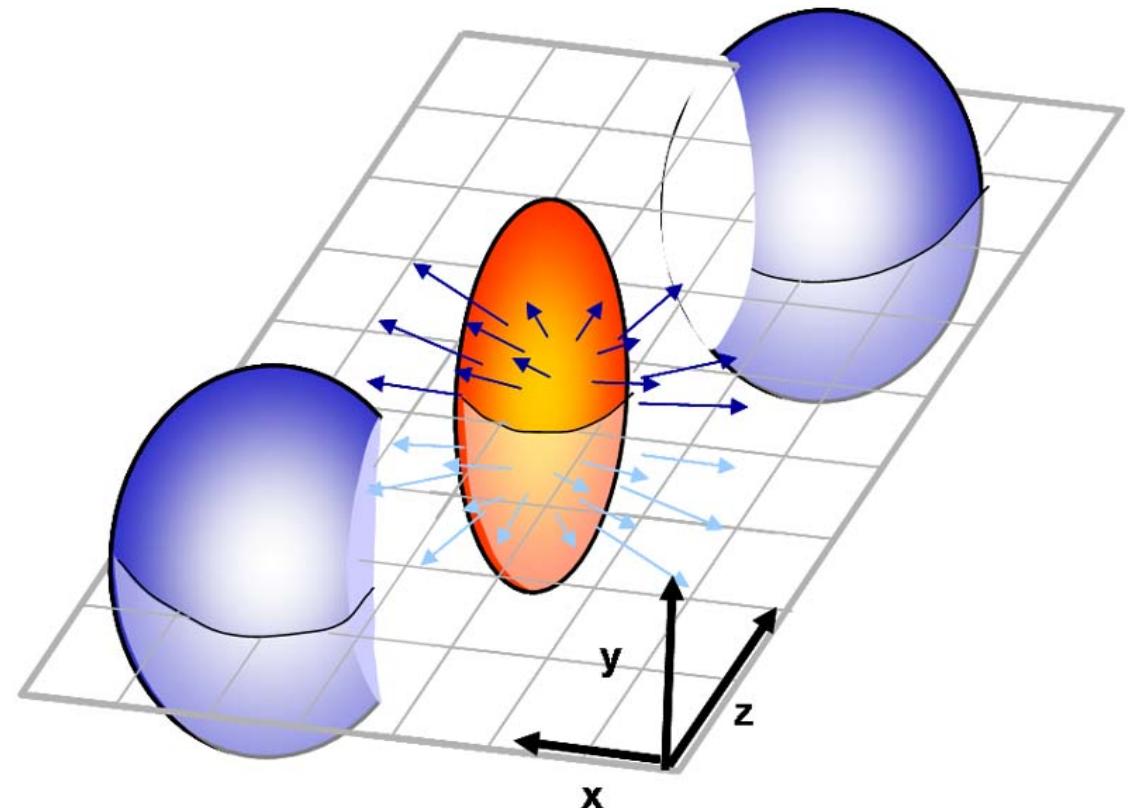
Shear

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Viscosity in heavy ion collisions

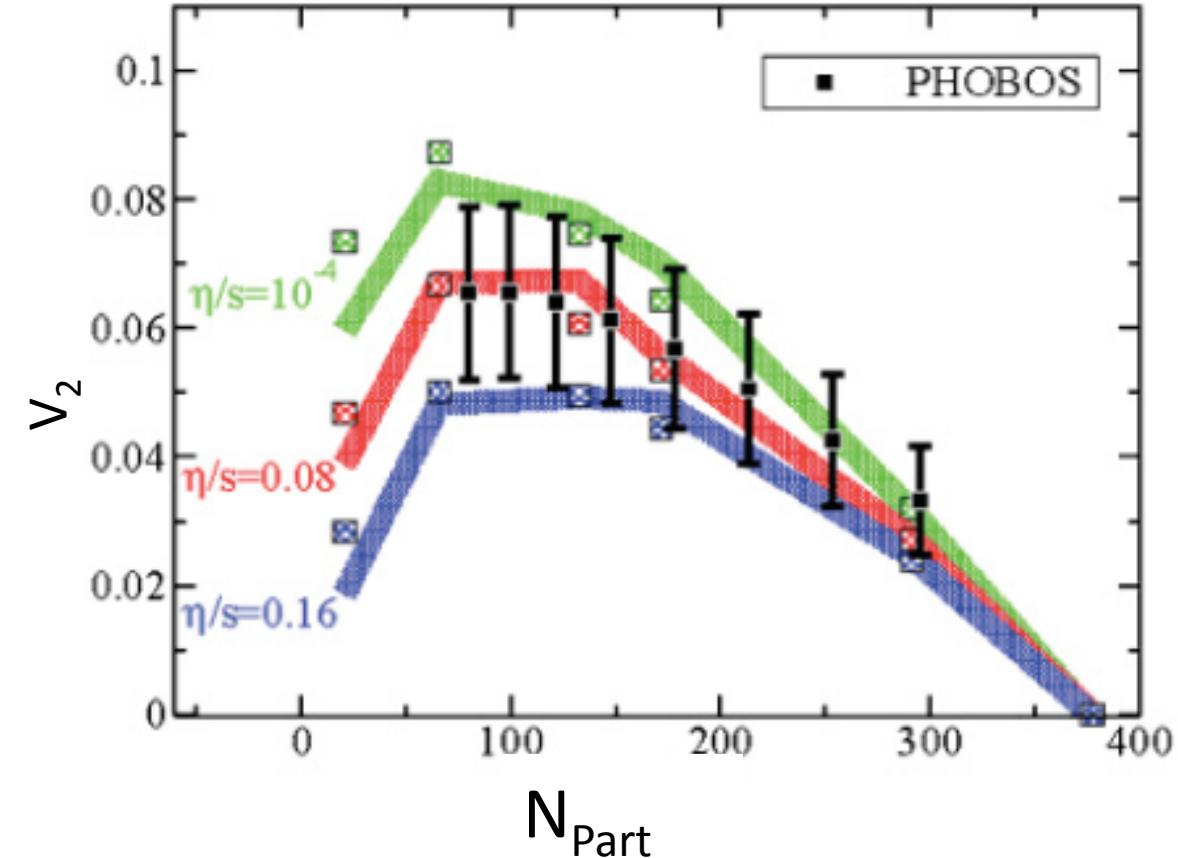
- RHIC and LHC measured large elliptic flow at the high energies corresponding to what is thought to be QGP



<http://www.quantumdiaries.org/wp-content/uploads/2011/02/FlowPr.jpg>

Viscosity in heavy ion collisions

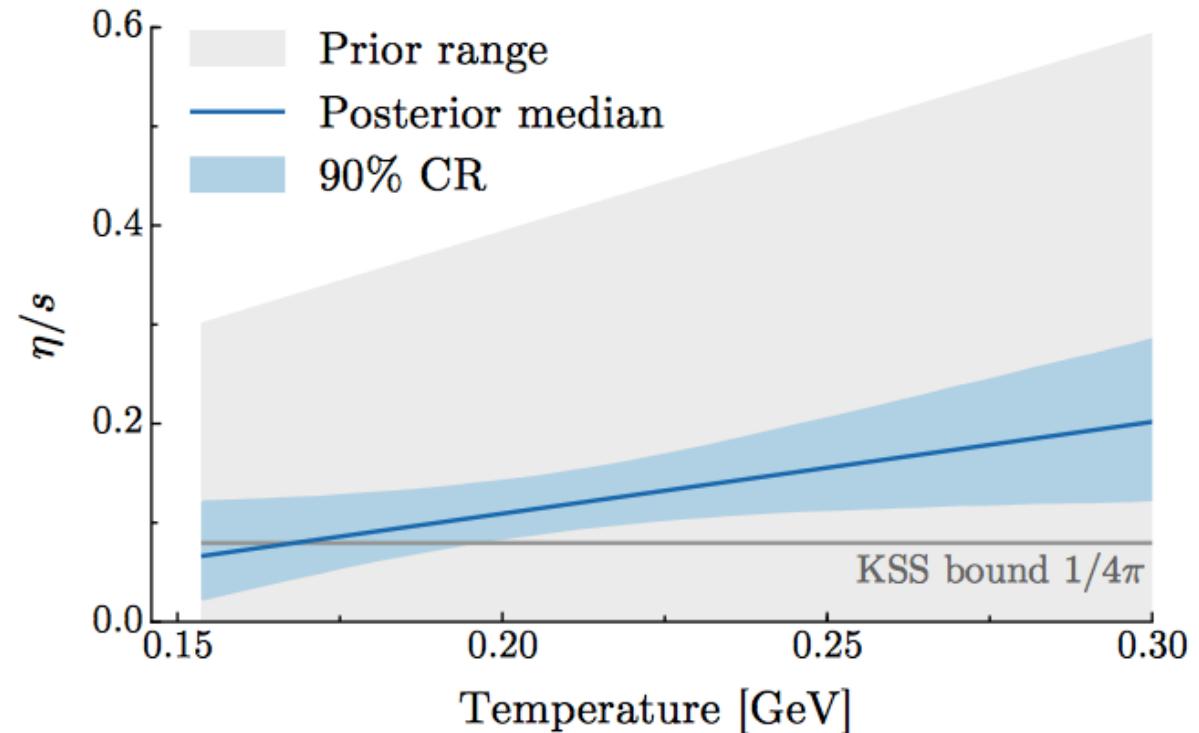
- RHIC and LHC measured large elliptic flow at the high energies corresponding to what is thought to be QGP
- Hydrodynamics relatively successful at explaining this with small η/s



Luzum & Romatschke 10.1103/Phys. Rev. C 78.034915

Viscosity in heavy ion collisions

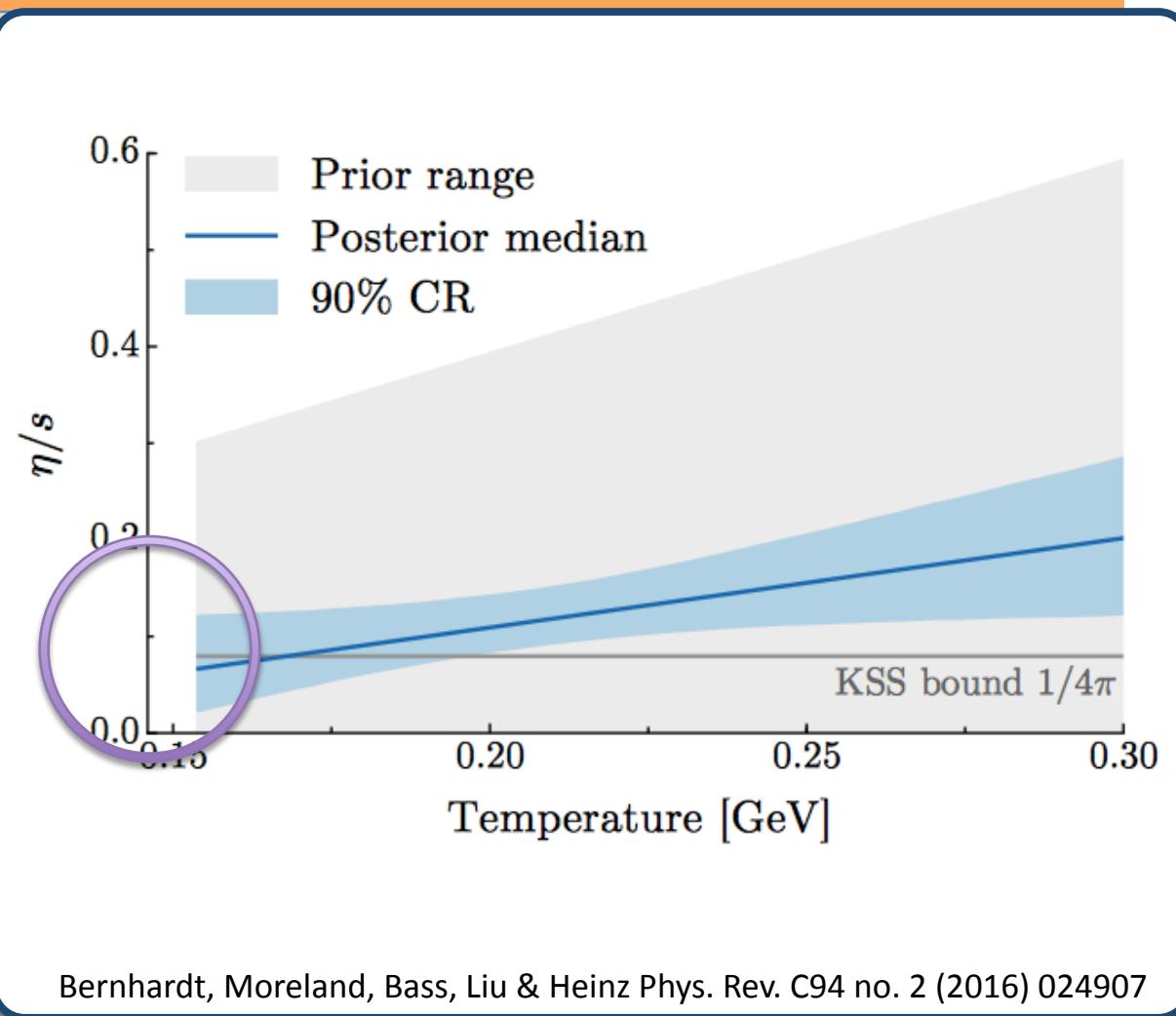
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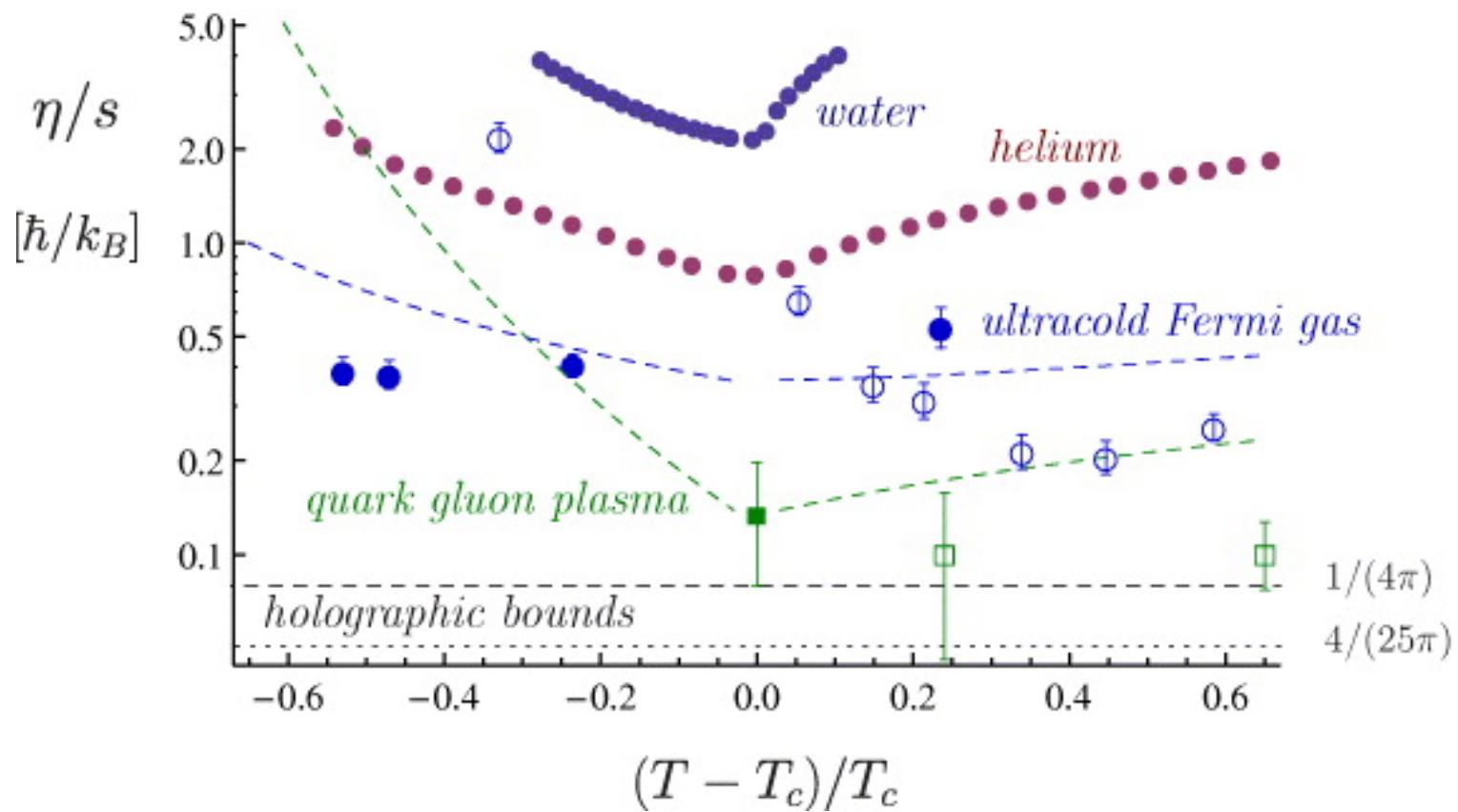
Bernhardt, Moreland, Bass, Liu & Heinz Phys. Rev. C94 no. 2 (2016) 024907

Viscosity in heavy ion collisions

- RHIC and LHC measured large elliptic flow at the high energies corresponding to what is thought to be QGP
- Hydrodynamics relatively successful at explaining this with small η/s
- Still not clear what the behavior of η/s is at low energies (FAIR, late stage RHIC/LHC)

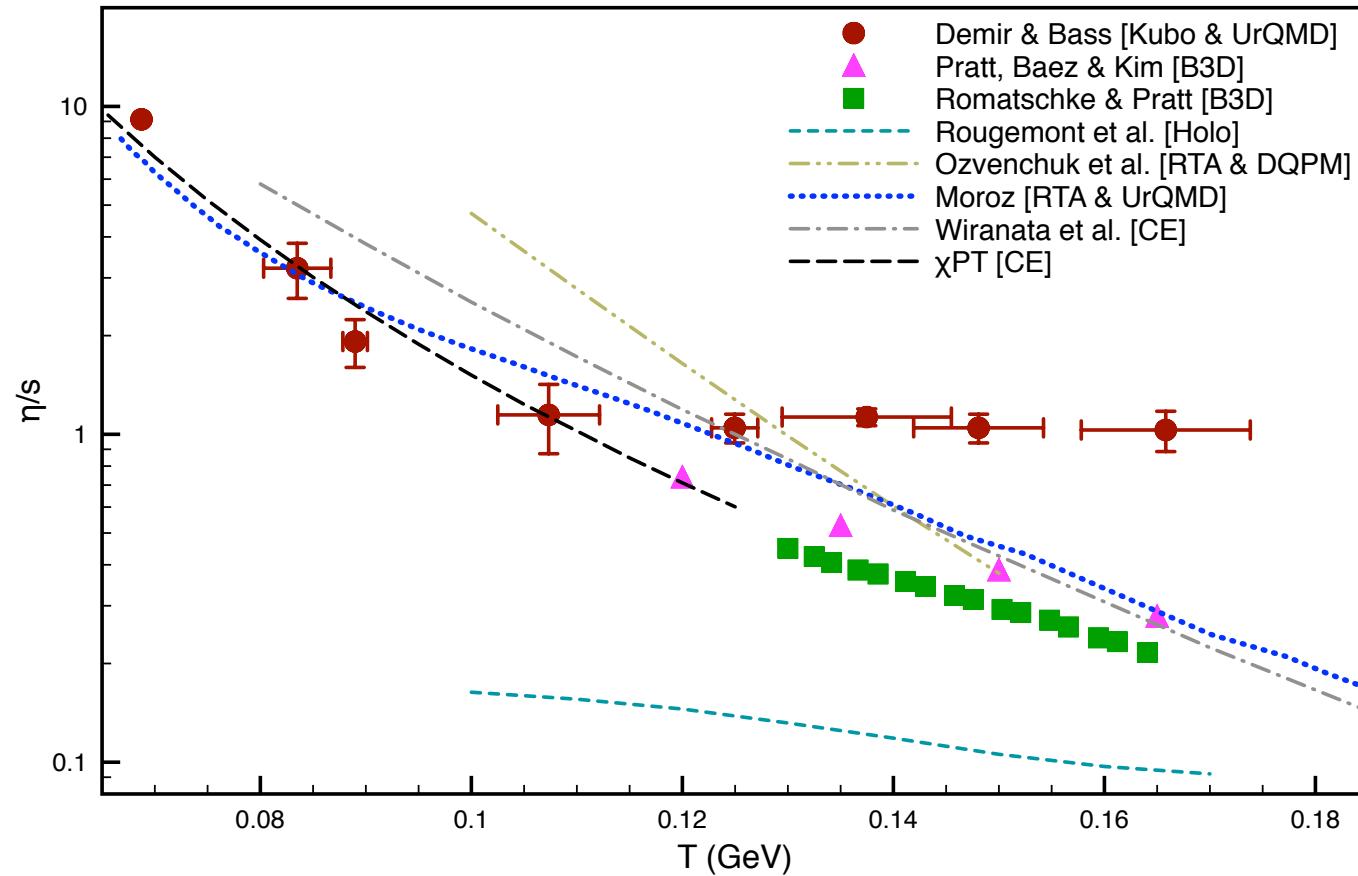


How low is low?



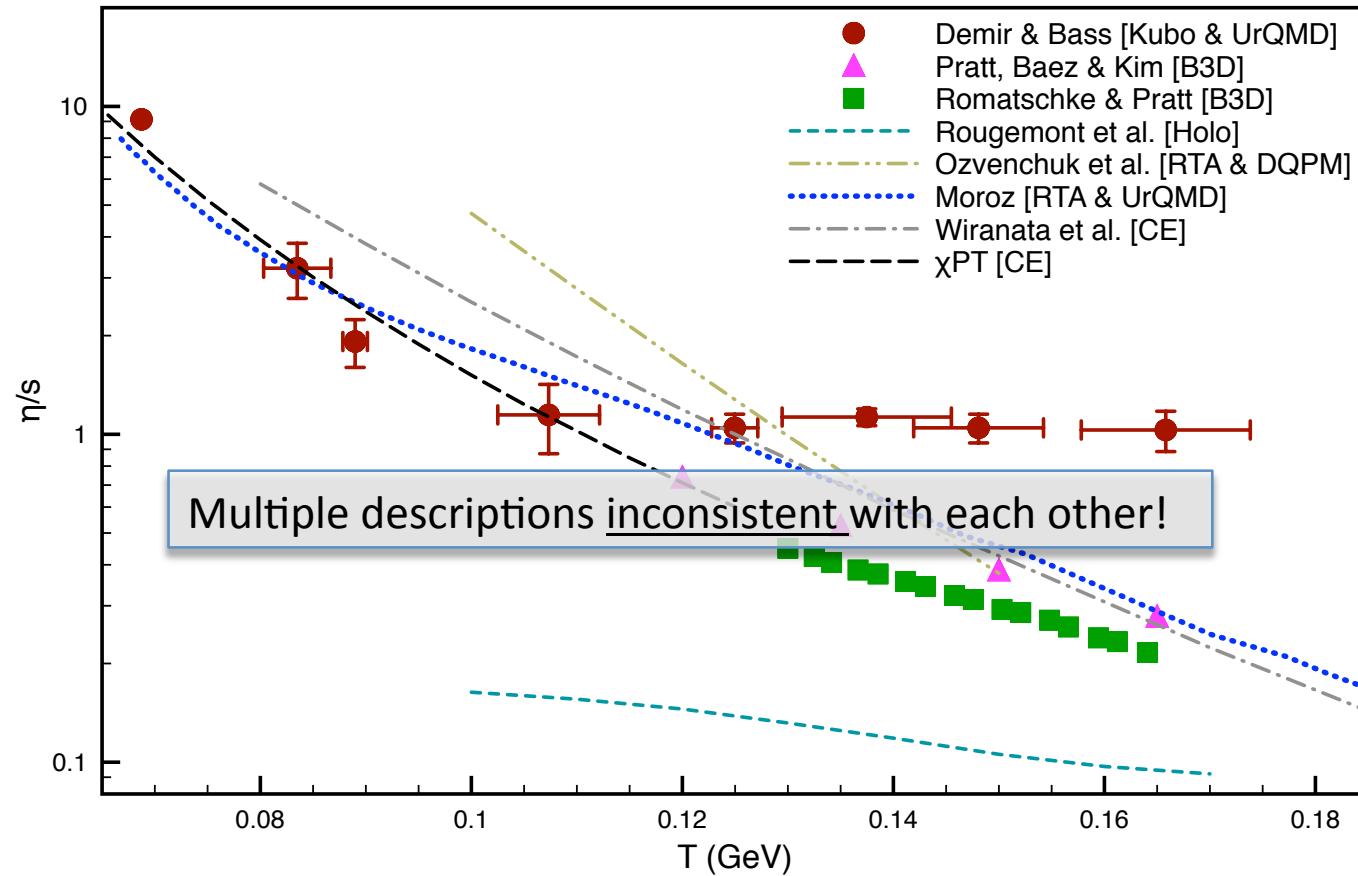
A. Adams, L. D. Carr, T. Schäfer, P. Steinberg, J E Thomas, New J. Phys. 15 (2013) 045022

Previous HG viscosity calculations



- Demir & Bass, Phys. Rev. Lett. 102 (2009) 172302
- Pratt, Baez & Kim, Phys. Rev. C95 (2017) 024901
- Romatschke & Pratt, arXiv:1409.0010v1
- Rougemont et al., arXiv:1704.05558
- Ozvenchuk et al., Phys. Rev. C87 (2013) 064903
- Moroz, arXiv:1301.6670
- Wiranata et al., Phys. Rev. C88 (2013) 044917
- Torres-Rincon, arXiv:1205.0782

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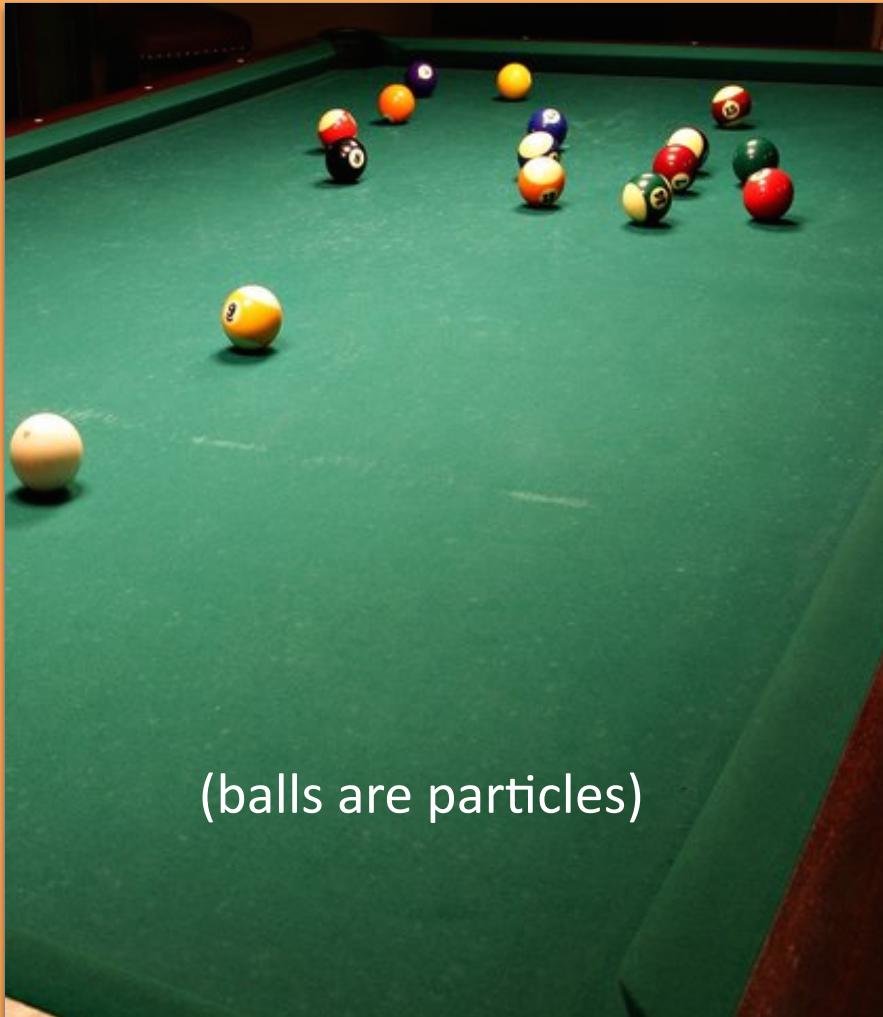


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Transport approaches



(balls are particles)

- **Transport models are 3D billiard tables**
- **But...**
 - Balls do not see each other as being the same size
 - Balls can annihilate
 - Balls can decay
 - Balls can become other balls on collision

Transport approaches



- More fundamentally, transport *effectively solves the Boltzmann equation:*

$$p^\mu \partial_\mu f_i(x, p) + m_i F^\alpha \partial_\alpha^p f_i(x, p) = C_{coll}^i$$

where $f_i(x, p)$ is the one-particle distribution function, F^α the force experienced by particles and C_{coll}^i is the collision term.

The transport code: SMASH



*Not the official logo. Definitely never will be.

- Transport *effectively* solves the Boltzmann equation:

$$p^\mu \partial_\mu f_i(x, p) + m_i F^\alpha \partial_\alpha^p f_i(x, p) = C_{coll}^i$$

- Particles represented by gaussian wave packets
- Geometric collision criterion:

$$d_{trans} < d_{int} = \sqrt{\frac{\sigma_{tot}}{\pi}}$$

- Each particles species is represented with point-like test particles

$$\begin{aligned}\sigma &\rightarrow \sigma \cdot N_{test}^{-1} \\ N &\rightarrow N \cdot N_{test}\end{aligned}$$

SMASH: General setup

- Timestepless evolution
 - All future possible actions stored
 - Propagate to next action
 - Update list of possible actions
 - And so on...
- Mean-field potentials
$$U = a \left(\frac{\varrho}{\varrho_0} \right) + b \left(\frac{\varrho}{\varrho_0} \right)^{\tau} \pm 2 S_{pot} \frac{\varrho I 3}{\varrho_0}$$
- Modii:
 - Collider, Sphere, Infinite matter, Afterburner

SMASH: Degrees of freedom

N	Δ	Λ	Σ	Ξ	Ω	Unflavored					Strange
N_{938}	Δ_{1232}	Λ_{1116}	Σ_{1189}	Ξ_{1321}	Ω^{-}_{1672}	π_{138}	$f_0\ 980$	$f_2\ 1275$	$\pi_2\ 1670$	K_{494}	
N_{1462}	Δ_{1620}	Λ_{1405}	Σ_{1385}	Ξ_{1532}	Ω^{-}_{2252}	π_{1300}	$f_0\ 1370$	$f_2'\ 1525$		K^*_{892}	
N_{1515}	Δ_{1700}	Λ_{1520}	Σ_{1660}	Ξ_{1690}		π_{1800}	$f_0\ 1500$	$f_2\ 1950$	$\rho_3\ 1690$	$K_1\ 1270$	
N_{1535}	Δ_{1905}	Λ_{1600}	Σ_{1670}	Ξ_{1820}			$f_0\ 1710$	$f_2\ 2010$		$K_1\ 1400$	
N_{1655}	Δ_{1910}	Λ_{1670}	Σ_{1750}	Ξ_{1950}		η_{548}		$f_2\ 2300$	$\phi_3\ 1850$	K^*_{1410}	
N_{1675}	Δ_{1920}	Λ_{1690}	Σ_{1775}	Ξ_{2030}		η'_{958}	$a_0\ 980$	$f_2\ 2340$		$K_0^*\ 1430$	
N_{1685}	Δ_{1930}	Λ_{1800}	Σ_{1915}			η_{1295}	$a_0\ 1450$		$a_4\ 2040$	$K_2^*\ 1430$	
N_{1700}	Δ_{1950}	Λ_{1810}	Σ_{1940}			η_{1405}		$f_1\ 1285$		K^*_{1680}	
N_{1710}		Λ_{1820}	Σ_{2030}			η_{1475}	ϕ_{1019}	$f_1\ 1420$	$f_4\ 2050$	$K_2\ 1770$	
N_{1720}		Λ_{1830}	Σ_{2250}				ϕ_{1680}			$K_3^*\ 1780$	
N_{1875}		Λ_{1890}				σ_{800}		$a_2\ 1320$		$K_2\ 1820$	
N_{1900}		Λ_{2100}					$h_1\ 1170$			$K_4^*\ 2045$	
N_{1990}		Λ_{2110}				ρ_{776}		$\pi_1\ 1400$			
N_{2000}		Λ_{2350}				ρ_{1450}	$b_1\ 1235$	$\pi_1\ 1600$			
N_{2190}						ρ_{1700}		$a_1\ 1260$	$\eta_2\ 1645$		
N_{2220}						ω_{783}					
N_{2250}						ω_{1420}			$\omega_3\ 1670$		
						ω_{1650}					
<ul style="list-style-type: none"> • Isospin symmetry • Perturbative treatment of non-hadronic particles (photons, dileptons) 											

SMASH: Resonances

- Particles stable if decay width smaller than 10 keV
- Spectral functions of resonances are described by relativistic Breit-Wigner functions

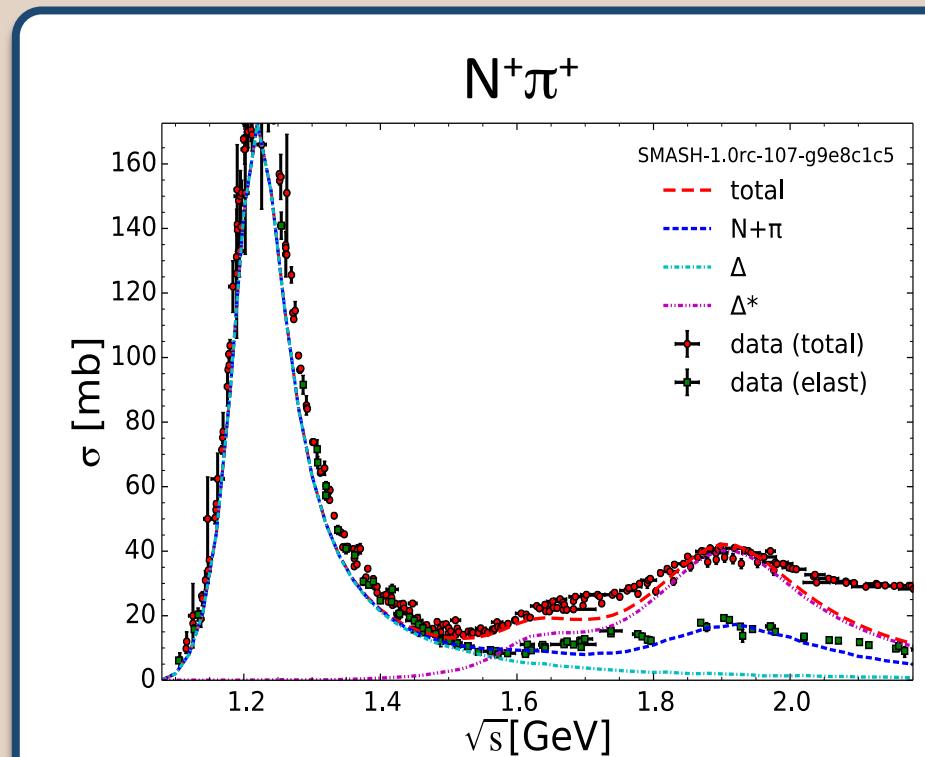
$$A(m) = \frac{2N}{\pi} \frac{m^2 \Gamma(m)}{\left(m^2 - M_0^2\right)^2 + m^2 \Gamma(m)^2}$$

- Manley et al. treatment for partial widths $\Gamma_{R \rightarrow ab}$
- Resonance lifetime for $1 \rightarrow 2$

$$\tau_{\text{res}} = \frac{1}{\Gamma(m)}$$

- Cross-sections for $2 \rightarrow 1$:

$$\sigma_{ab \rightarrow R} = \frac{2J_R + 1}{(2J_a + 1)(2J_b + 1)} S_{ab} \frac{2\pi^2}{p_i^2} \Gamma_{ab \rightarrow R}(s) A_R(\sqrt{s})$$



Weil et al., Phys. Rev. C94 (2016) no.5, 054905

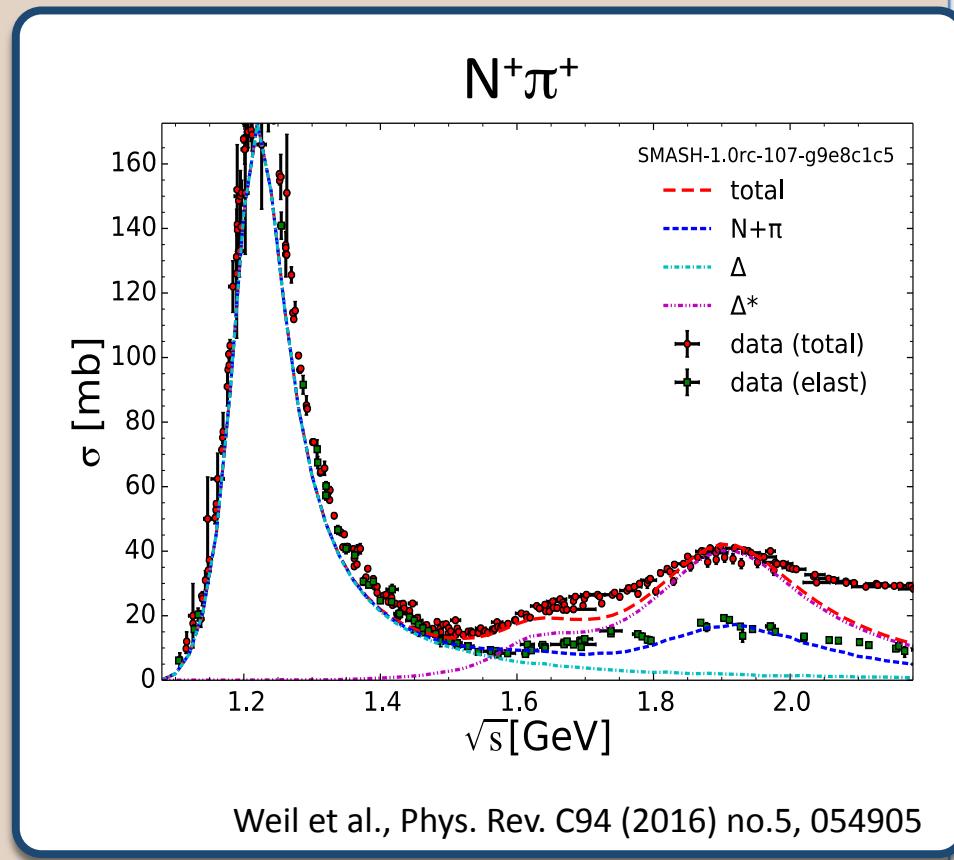
Manley et al., Phys. Rev. D45 (1992), 4002

SMASH: Scatterings

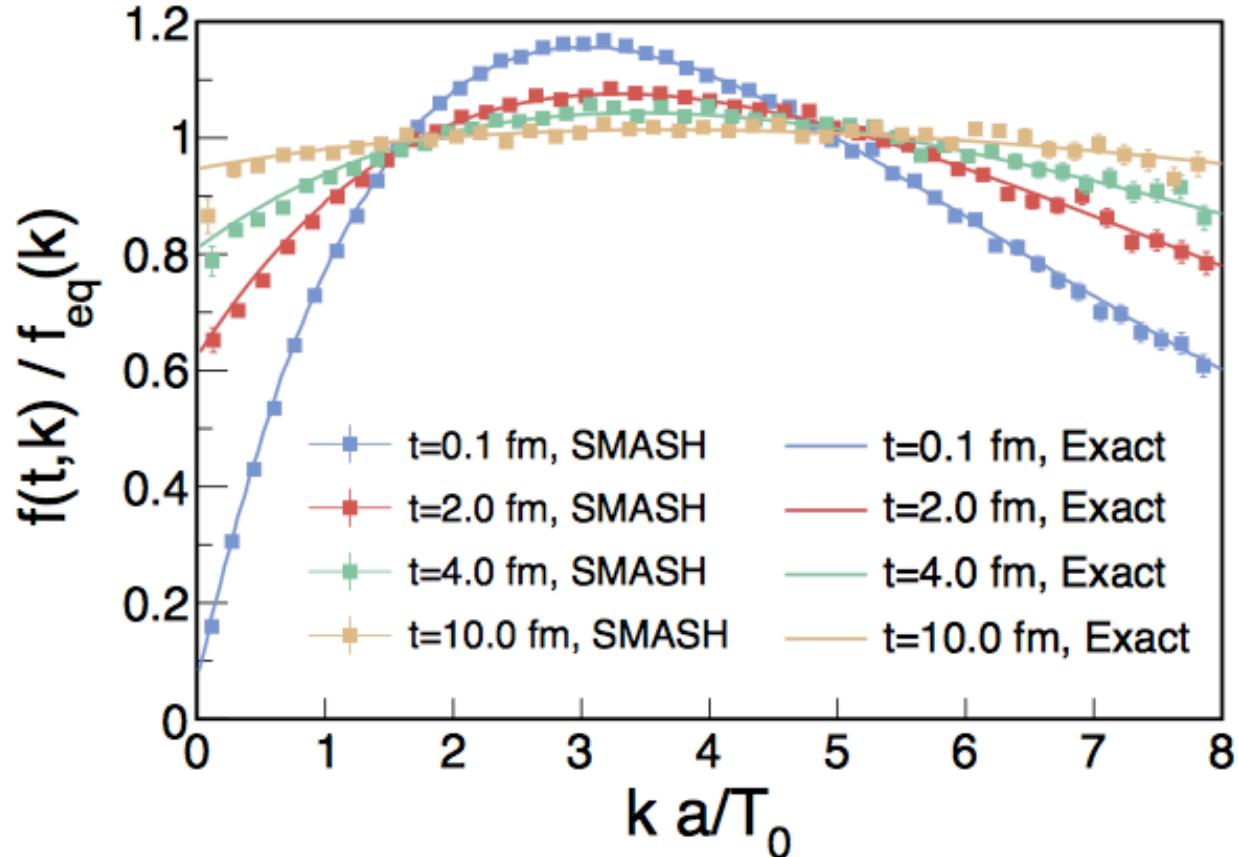
- Elastic scatterings parameterized for NN reactions; all other elastic scatterings assumed to go through resonances
- Inelastic scatterings implemented:

$$\sigma_{ab \rightarrow R_c}(s) = \frac{(2J_R + 1)(2J_c + 1)}{s |p_i|} \cdot \sum (C_{ab}^I C_{Rc}^I)^2 \frac{|M|_{ab \leftrightarrow R_c}^2(s, I)}{16\pi} \\ \cdot \sqrt{s - m_c} \cdot \int_{m_R^{\min}}^{m_R^{\max}} dm A_R(m) \cdot |p_f|(\sqrt{s}, m, m_c)$$

- Currently include
 - NN \leftrightarrow NR, NN \leftrightarrow ΔR
 - KN \leftrightarrow KN, KN \leftrightarrow πH
 - +antiparticles
- Strings (turned off)



SMASH: Does it all work?

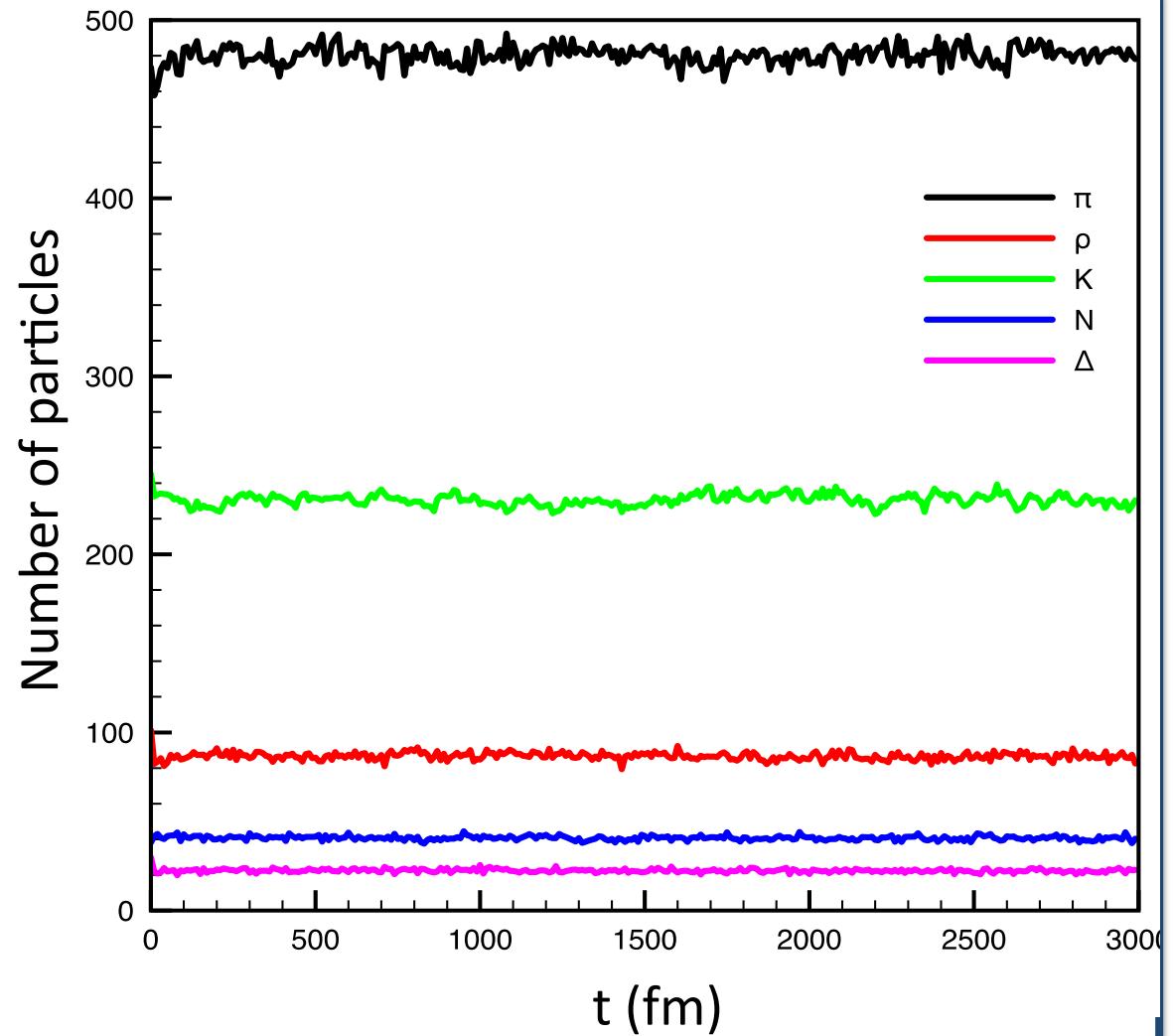


Tindall et al., Phys.Lett. B770 (2017) 532-538

Exact solution of Boltzmann equation in expanding universe

Viscosity in SMASH

- Box calculations simulating infinite matter to apply the Green-Kubo procedure
- MUST have thermal & chemical equilibrium
 - Baryon/antibaryon annihilation implemented to conserve detailed balance via an average decay to 5π



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Green-Kubo Formalism

The shear viscosity
is calculated from

$$\eta = \frac{V}{T} \int_0^\infty C^{xy}(t) dt$$

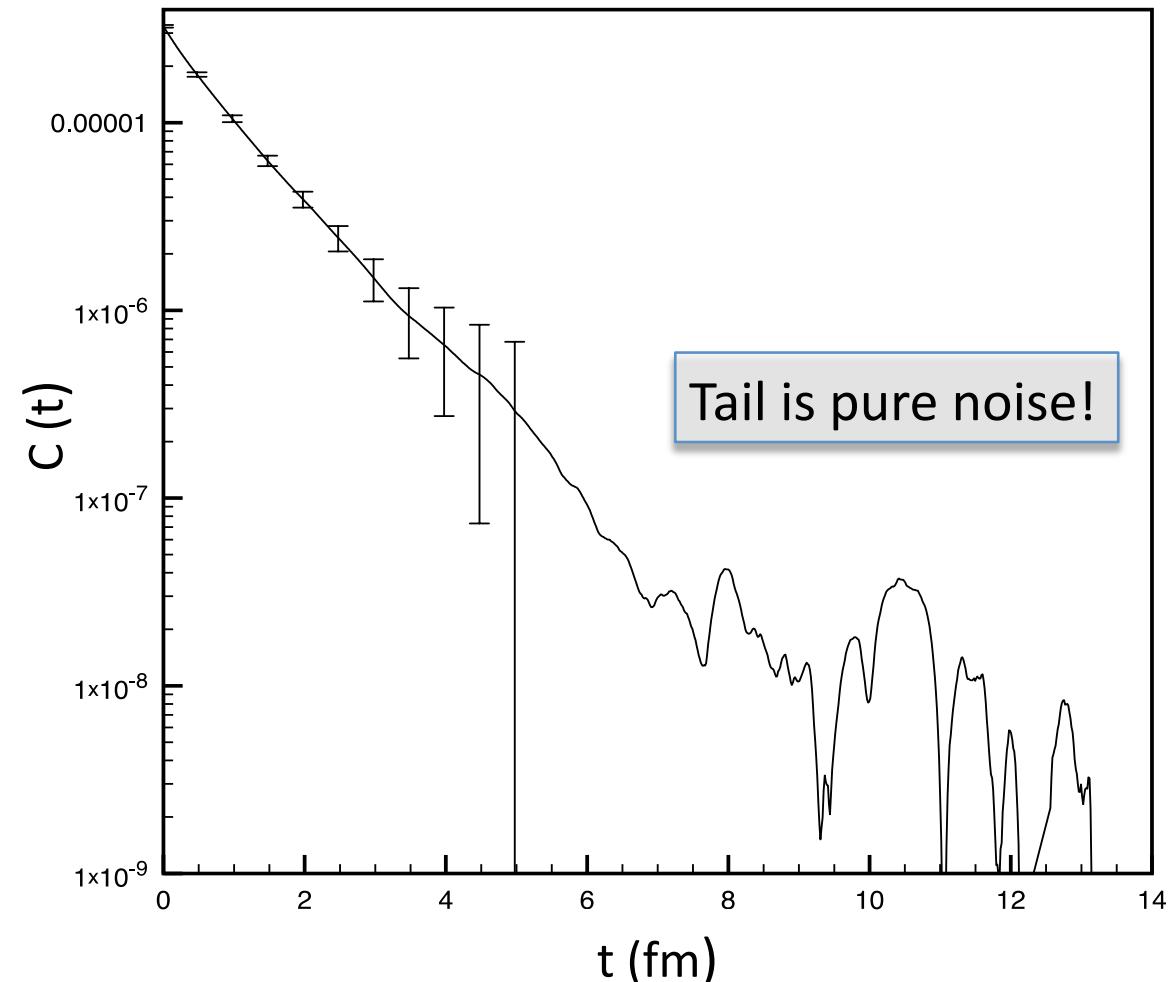
where

$$C^{xy}(t) = \frac{1}{N} \sum_s T^{xy}(s) T^{xy}(s+t)$$

and

$$T^{\mu\nu} = \frac{1}{V} \sum_i^{N_{part}} \frac{p_i^\mu p_i^\nu}{p_i^0}$$

N is the number of time steps, and N_{part} the number of particles



Green-Kubo Formalism

It has been shown that
the correlation

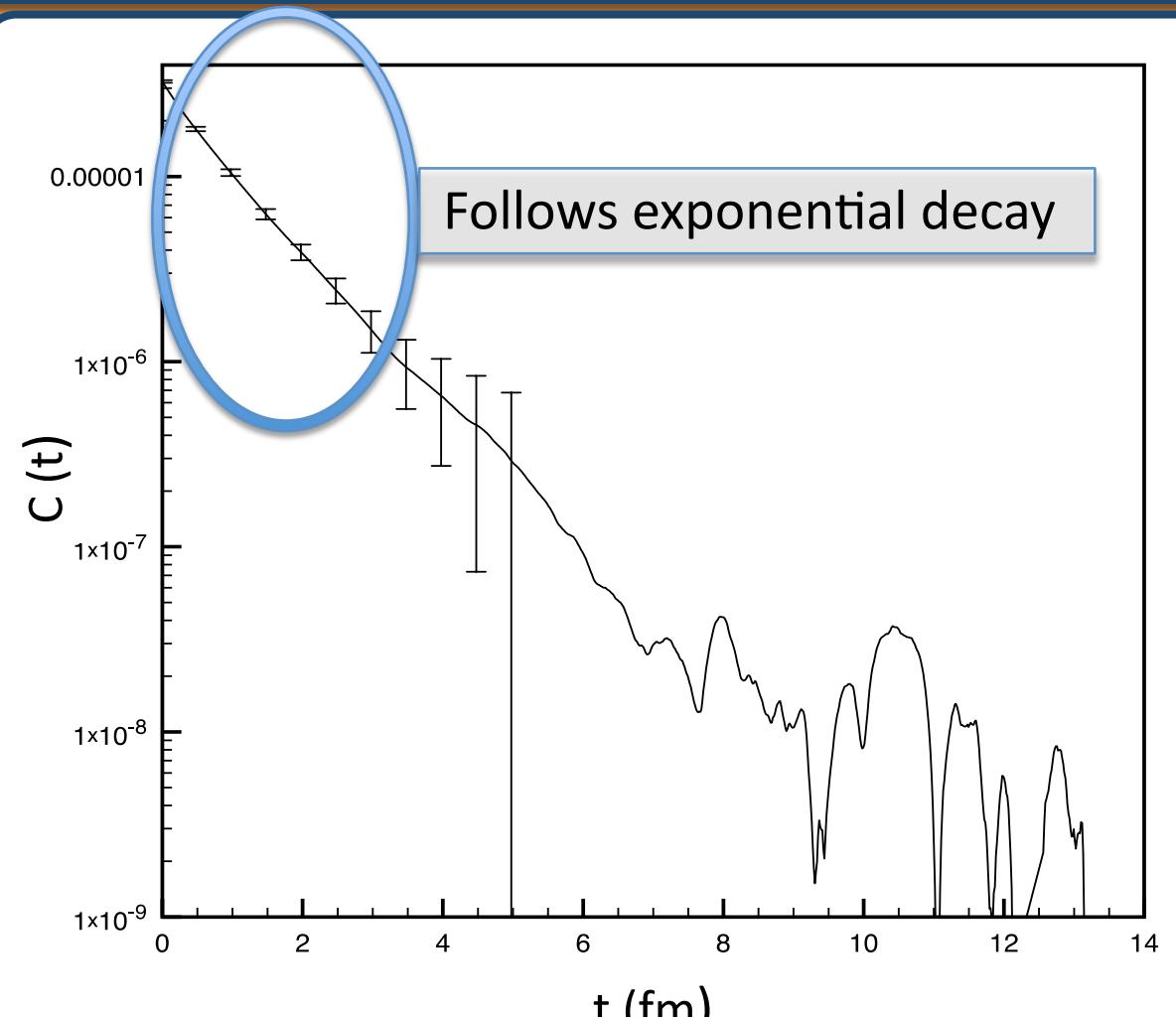
$$\eta = \frac{V}{T} \int_0^{\infty} C^{xy}(t) dt$$

Follows

$$C^{xy}(t) = C^{xy}(0) \exp\left(-\frac{t}{\tau}\right)$$

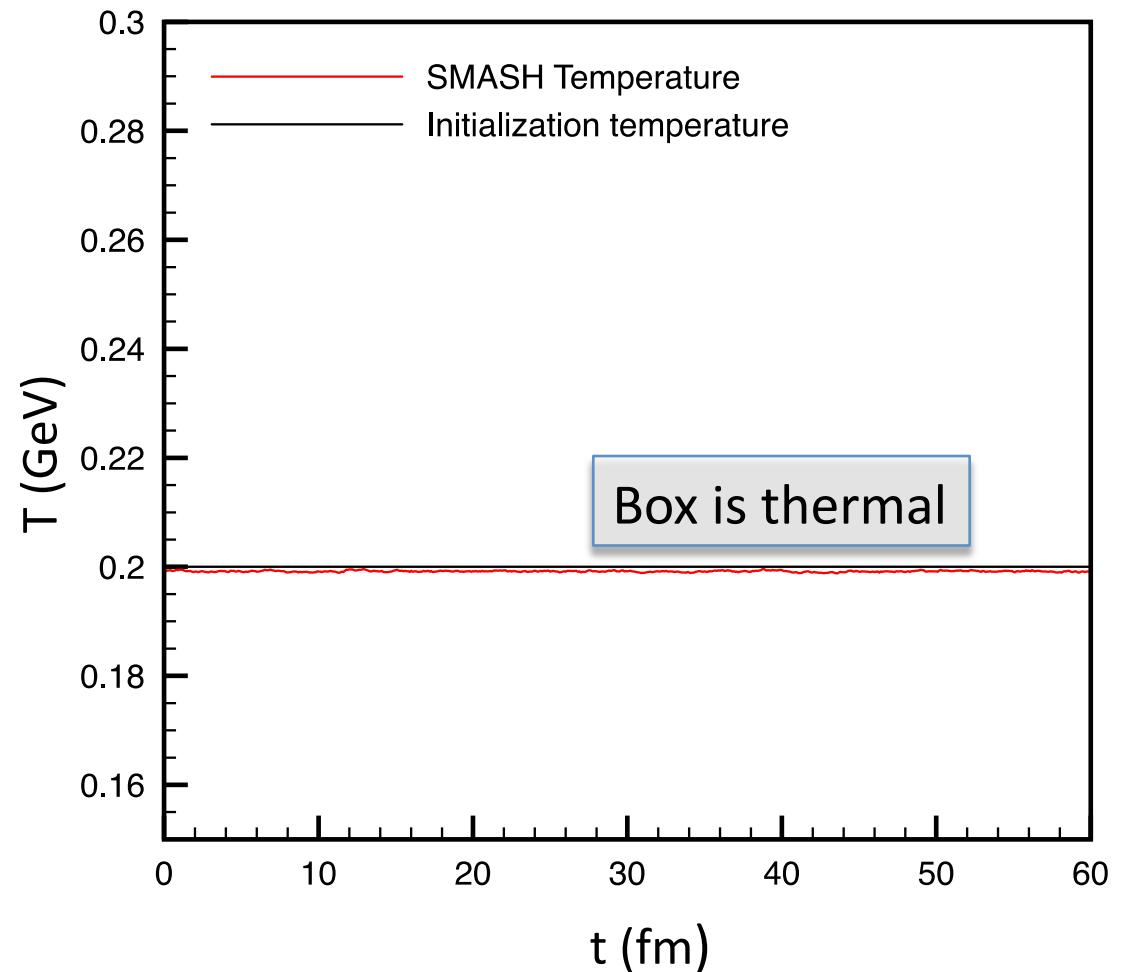
So that

$$\eta = \frac{VC^{xy}(0)\tau}{T}$$



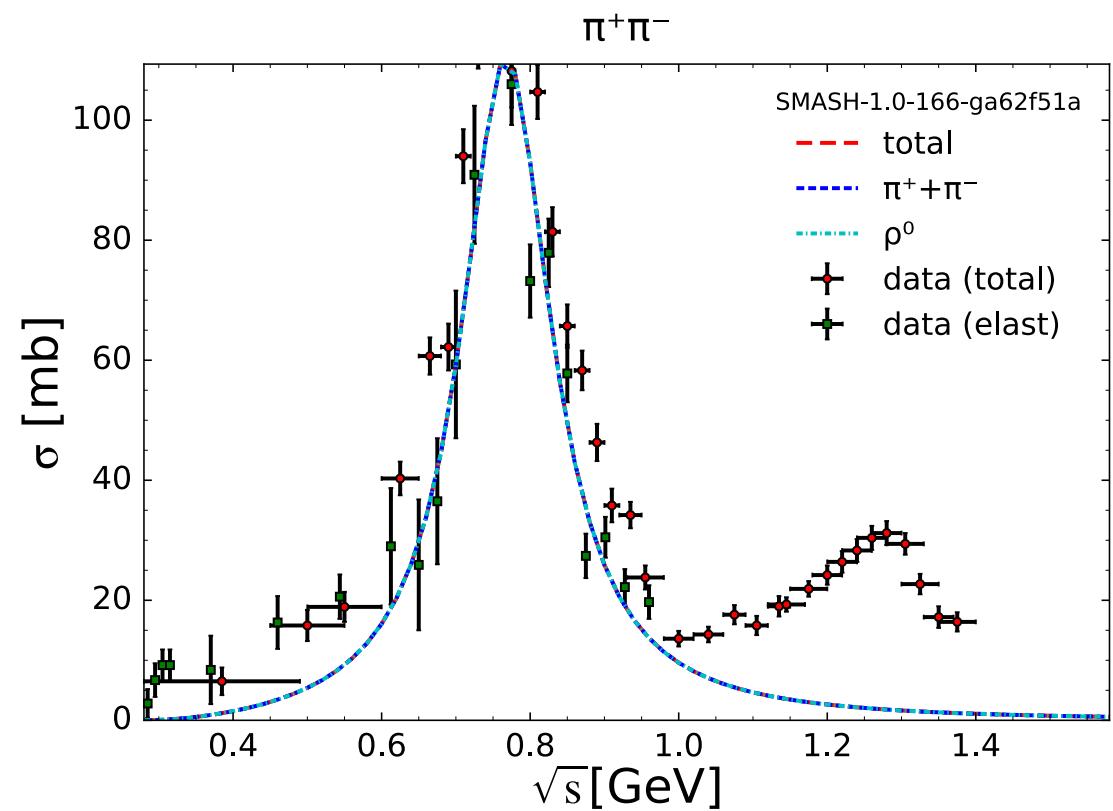
Test case #1: Constant σ

- Pions in a $(20 \text{ fm})^3$ box simulating infinite matter
- Constant, isotropic σ
- Runs for $t_{max}=200 \text{ fm}$
- Initialized with initial densities consistent with Boltzmann ideal gas



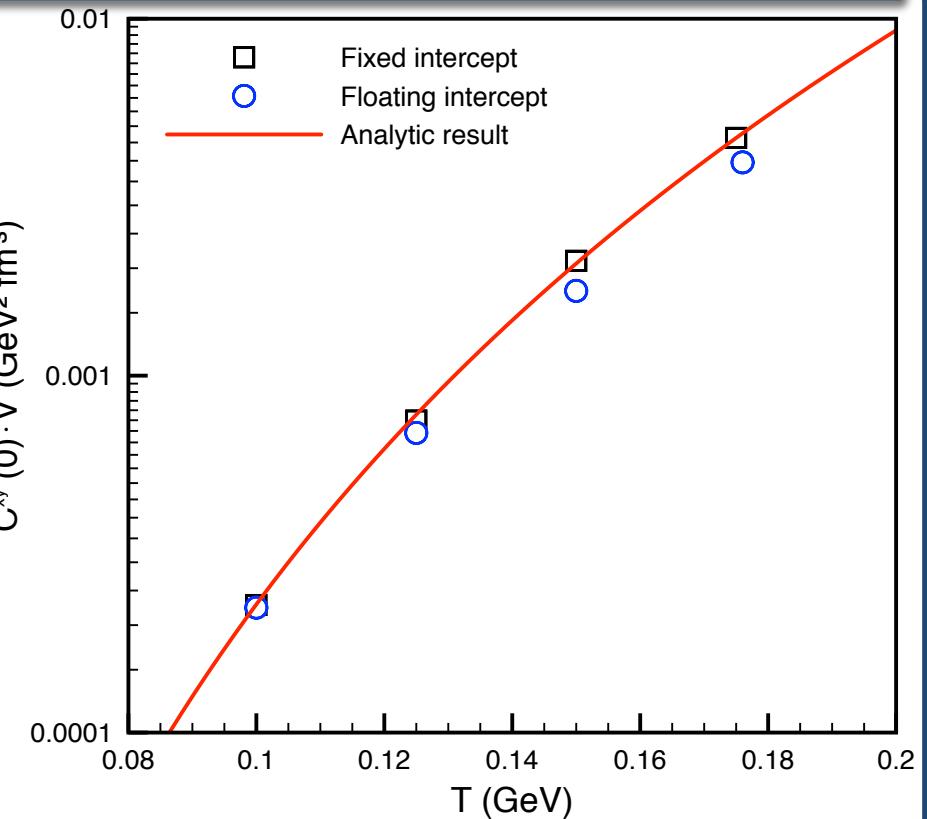
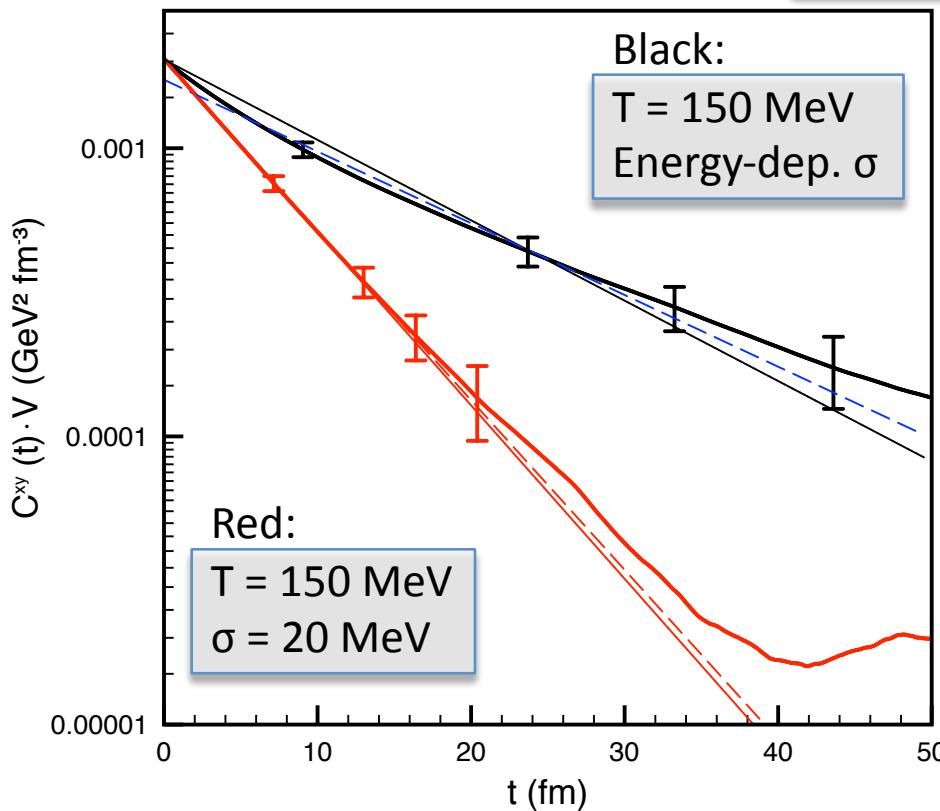
Test case #2: Energy-dependent σ

- Pions in a $(20 \text{ fm})^3$ box simulating infinite matter
- Cross-section uses ρ resonance
- Runs for $t_{max} = 200 \text{ fm}$
- Initialized with initial densities consistent with Boltzmann ideal gas



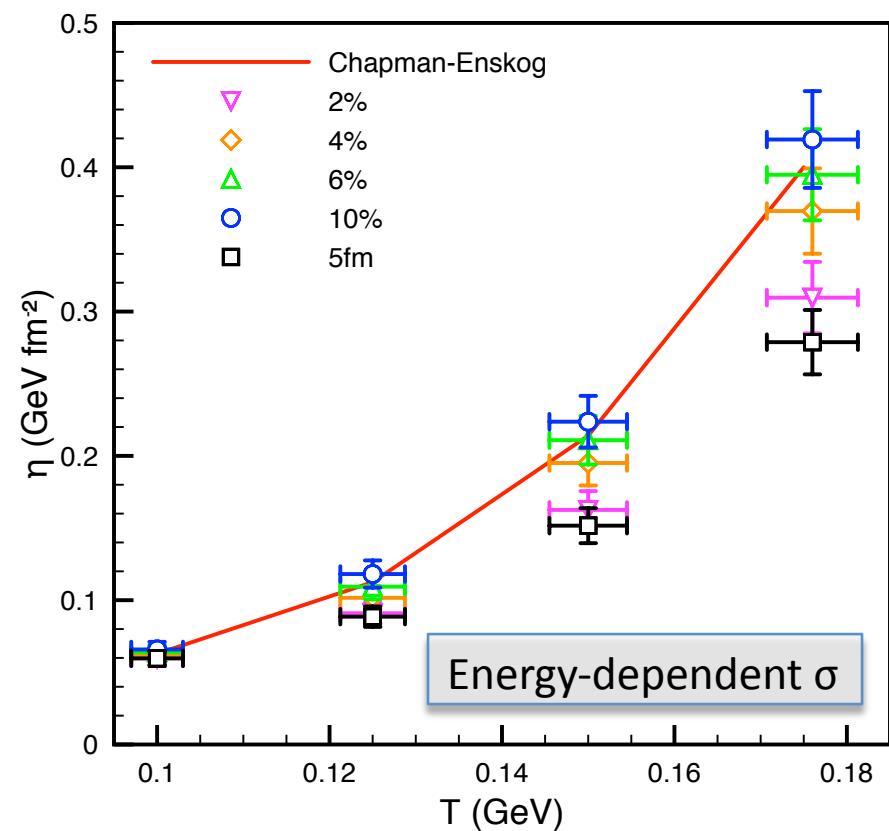
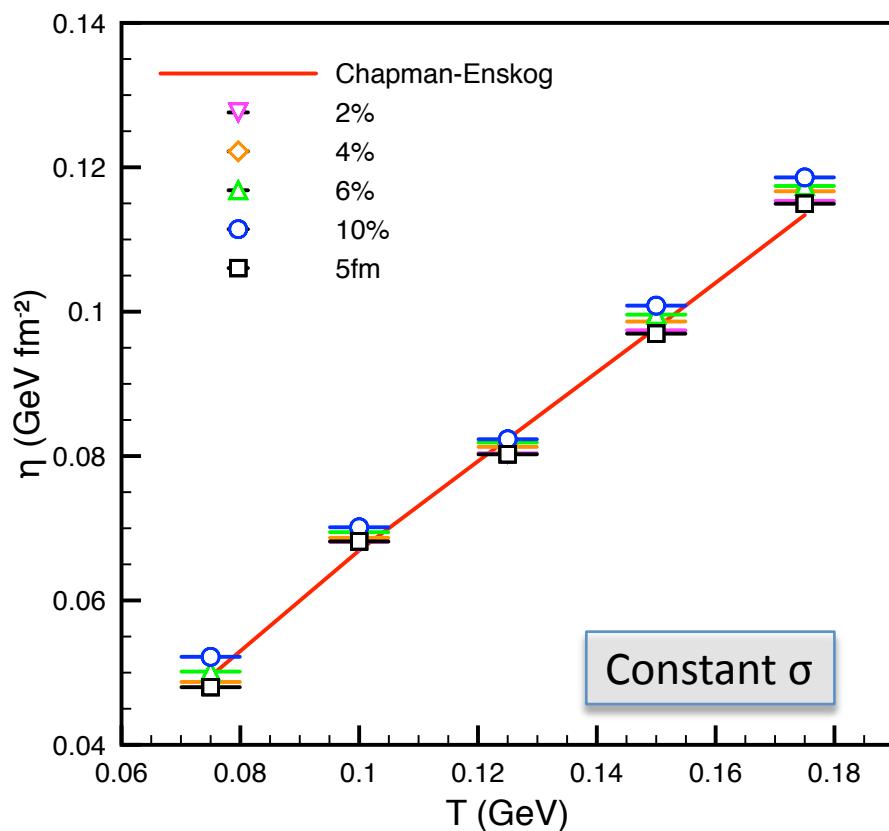
How to fit?

$$C^{xy}(0) = \frac{g \exp\left(\frac{\mu}{T}\right)}{30\pi^2 V} \int_0^\infty dp \frac{p^6}{m^2 + p^2} \exp\left(-\frac{\sqrt{m^2 + p^2}}{T}\right)$$

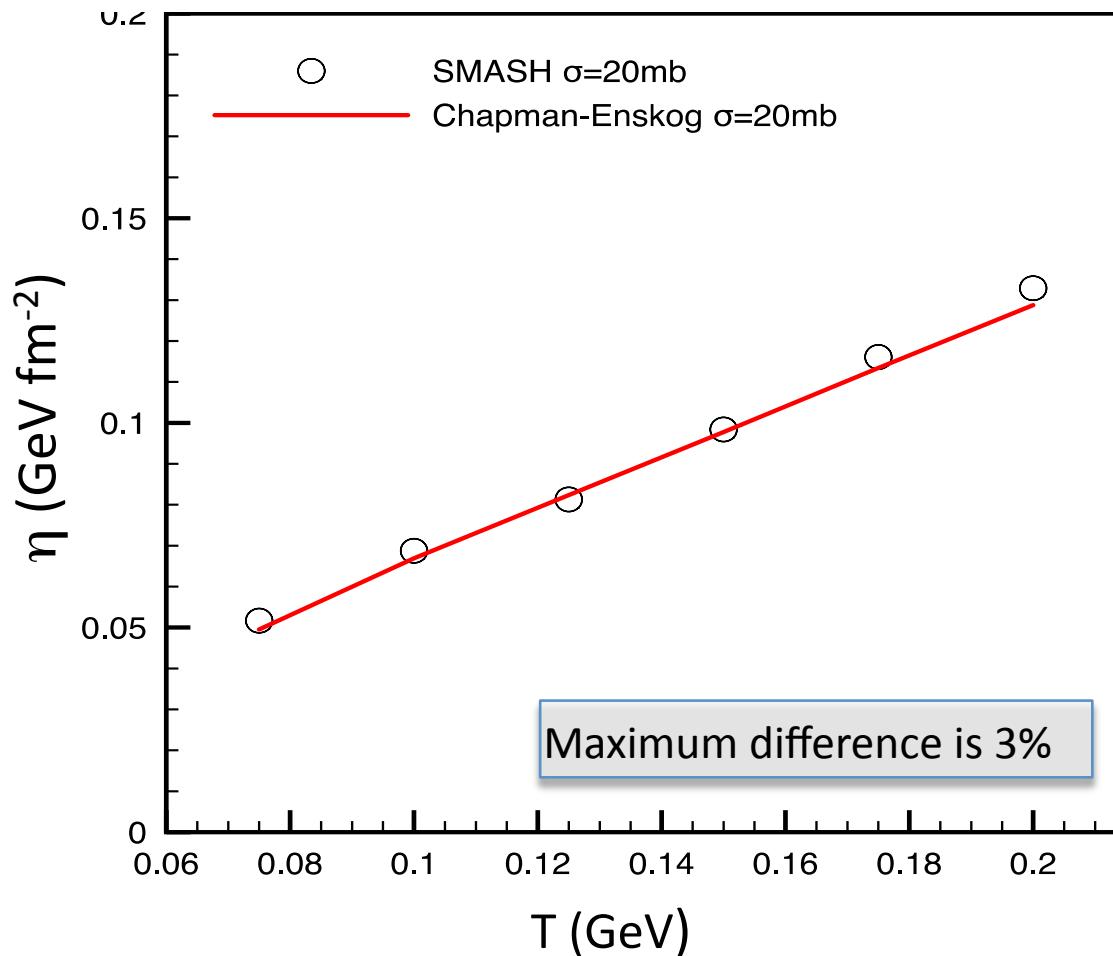


Where to stop?

J. Torres-Rincon, PhD dissertation (2012), *Hadronic Transport Coefficients from Effective Field Theories*

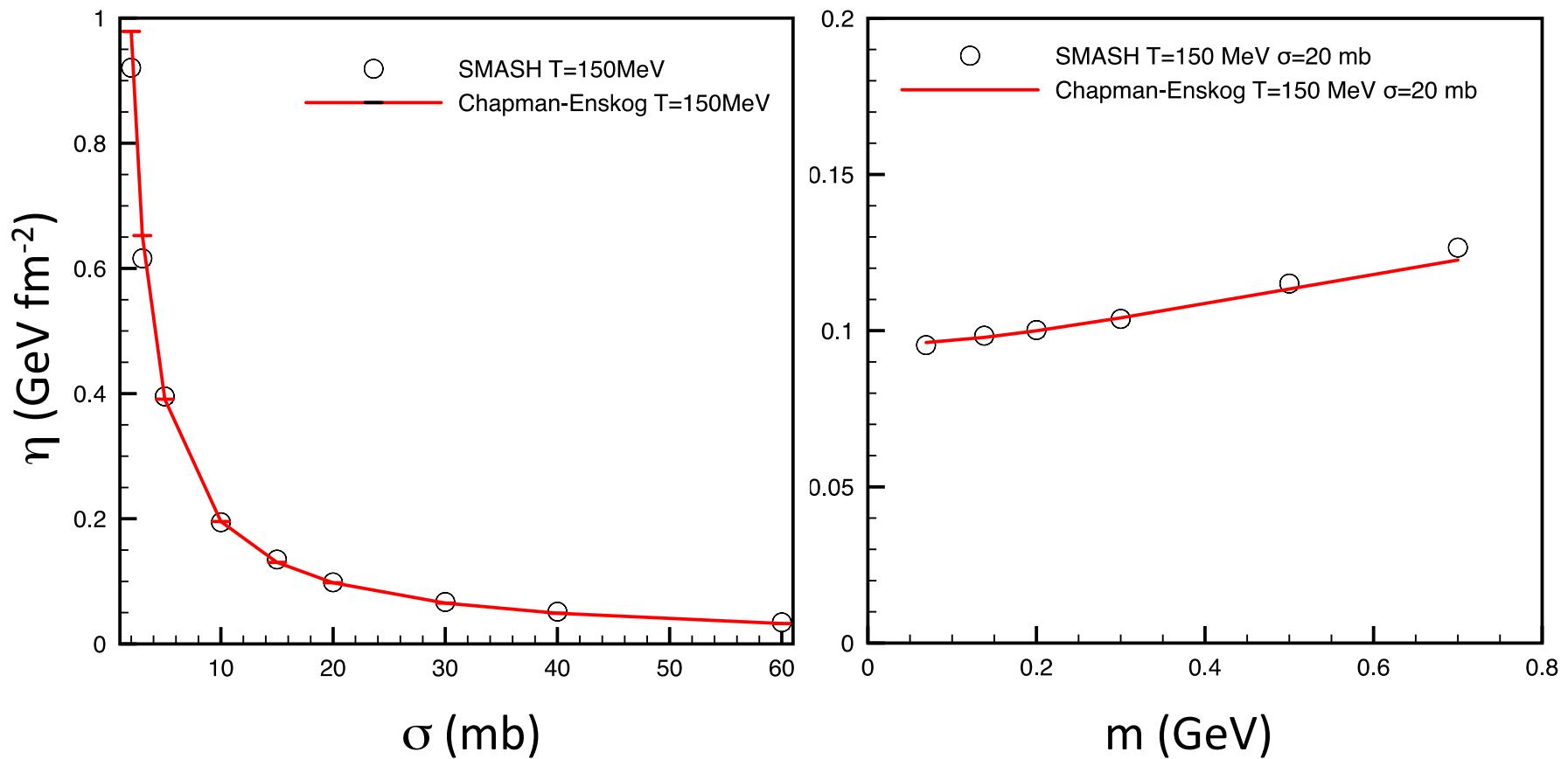


Constant σ : Temperature dependence



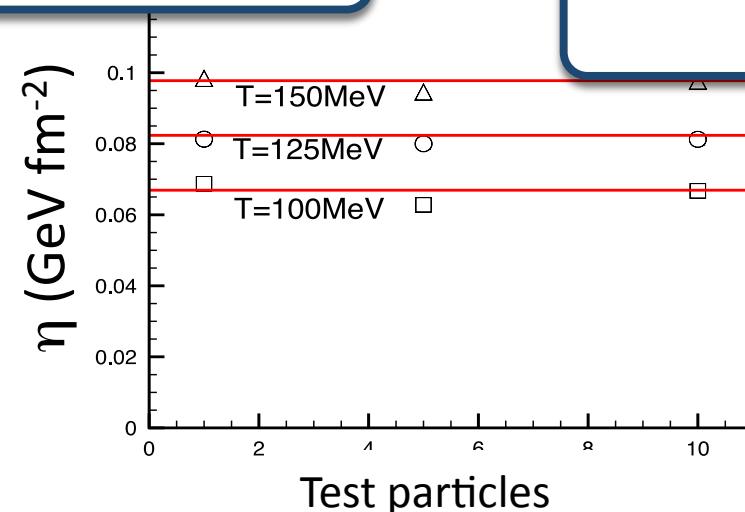
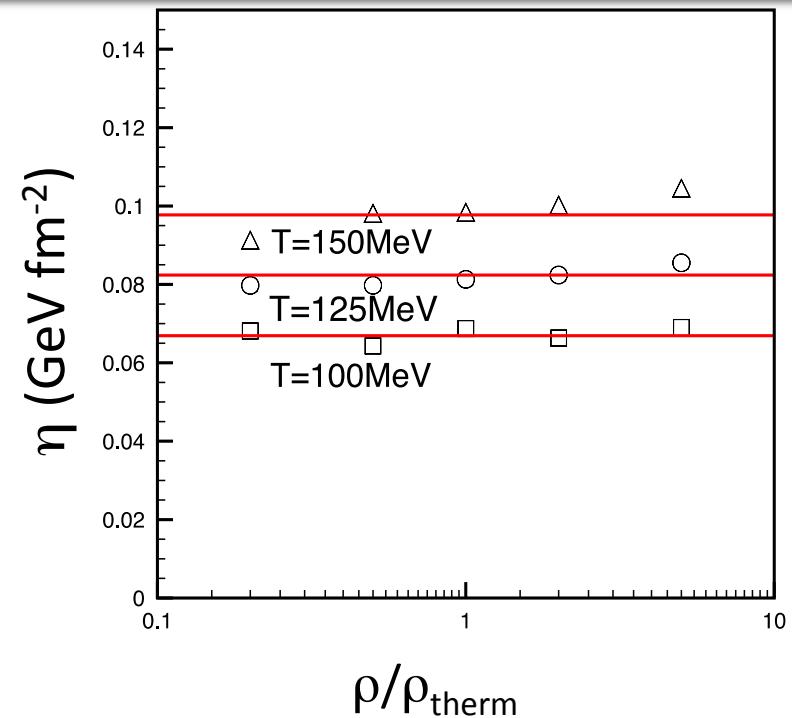
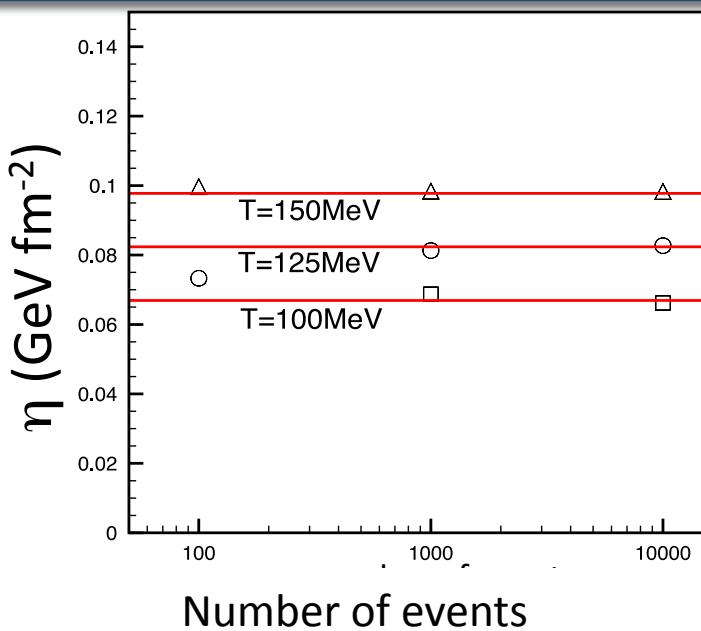
J. Torres-Rincon, PhD dissertation (2012), *Hadronic Transport Coefficients from Effective Field Theories*

Constant σ : Cross-section/mass dependence



Very good agreement with analytical calculations!

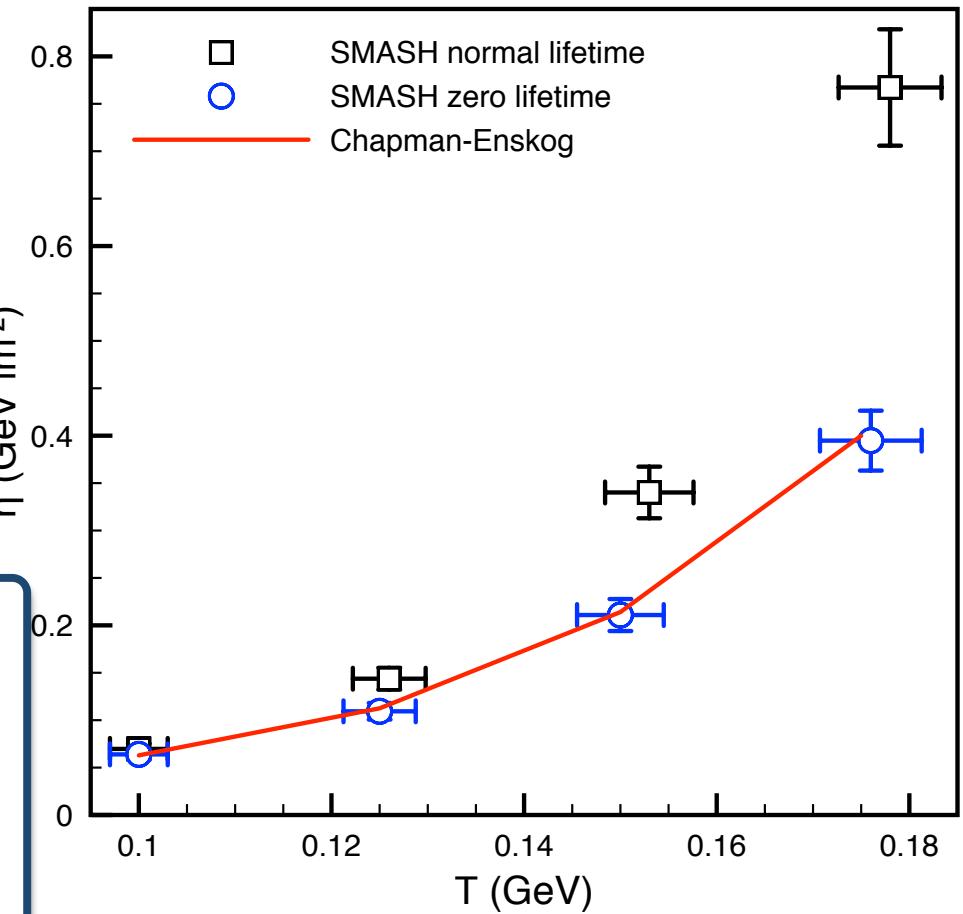
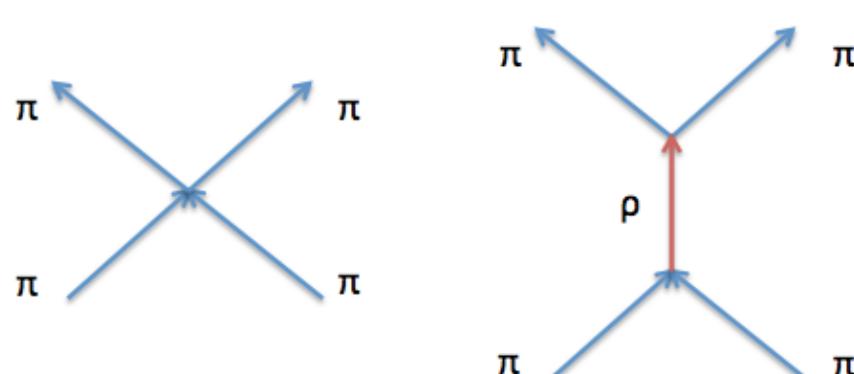
Constant σ : Systematics



Main take-away:
The method is relatively inelastic to variations of most parameters; maximum error is less than 10%

Energy-dependent σ : Resonance lifetimes

- Normal SMASH run does not coincide directly with Chapman-Enskog
 - Resonance lifetimes



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What about entropy?

The entropy density can be calculated from the Gibbs formula:

$$S = \frac{e + p - \mu n}{T} = \frac{w - \mu n}{T}$$

where the energy density and pressure can be taken from the average shear-stress tensor according to:

$$T^{\mu\nu} = \text{diag}(e, p, p, p)$$

Assuming a nearly ideal gas, one can fit the temperature and chemical potential with momentum distributions:

$$\frac{dN}{dp} = \frac{g}{2\pi^2} V p^2 \exp\left(-\frac{\sqrt{p^2 + m^2} - \mu}{T}\right)$$

What about entropy?

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When is this correct?

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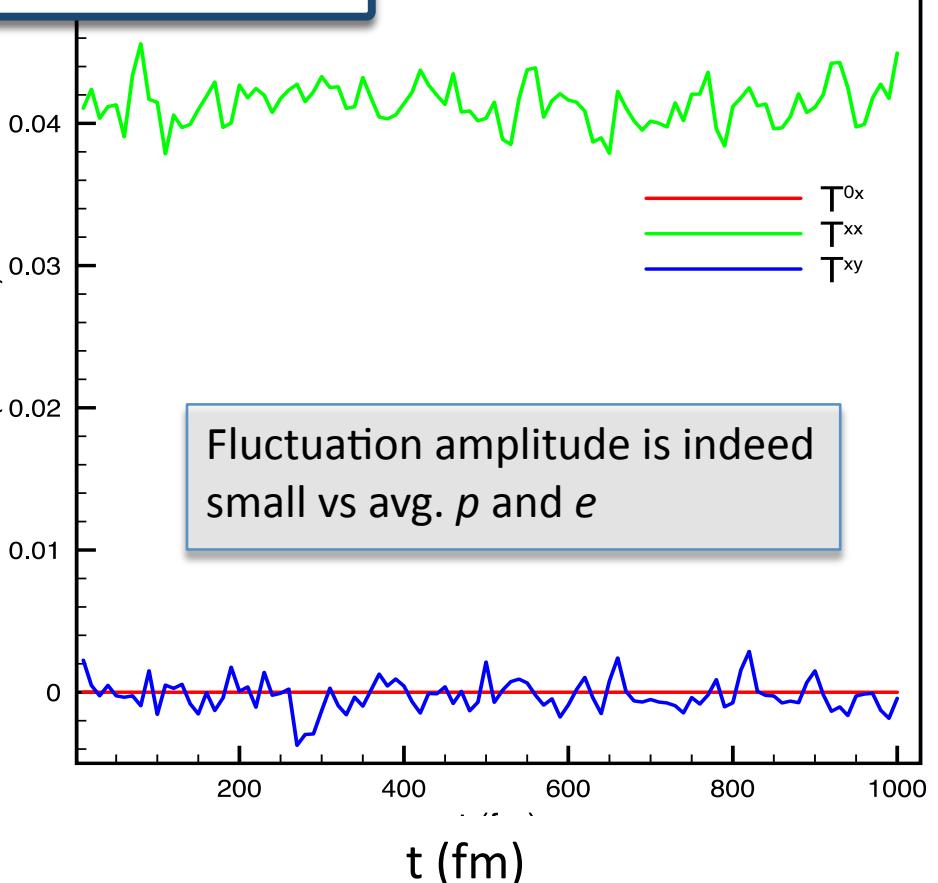
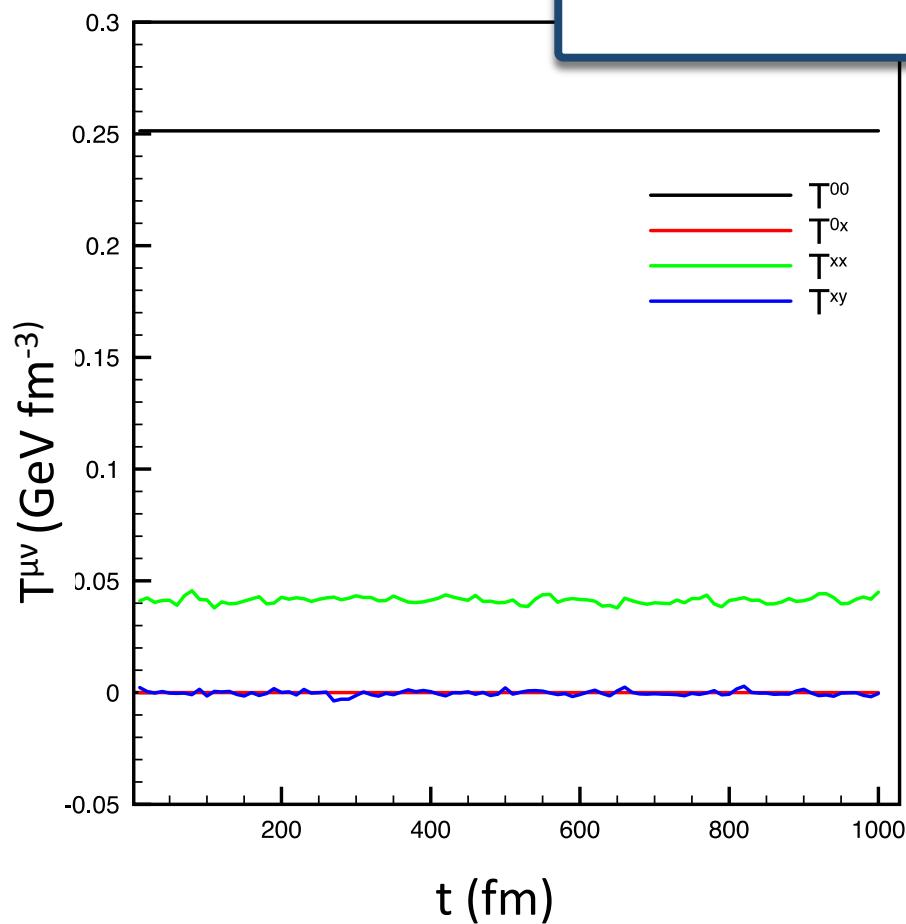
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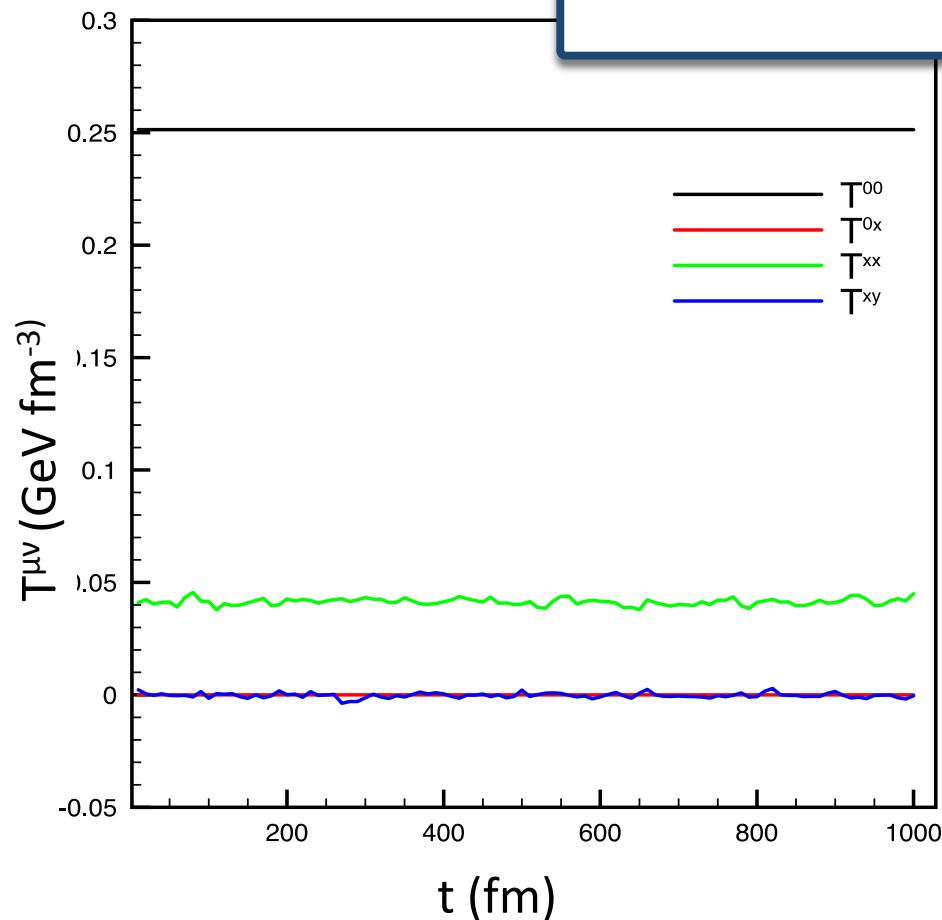
Energy density and pressure

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Energy density and pressure

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$$\frac{p}{e} \sim \frac{1}{6}$$

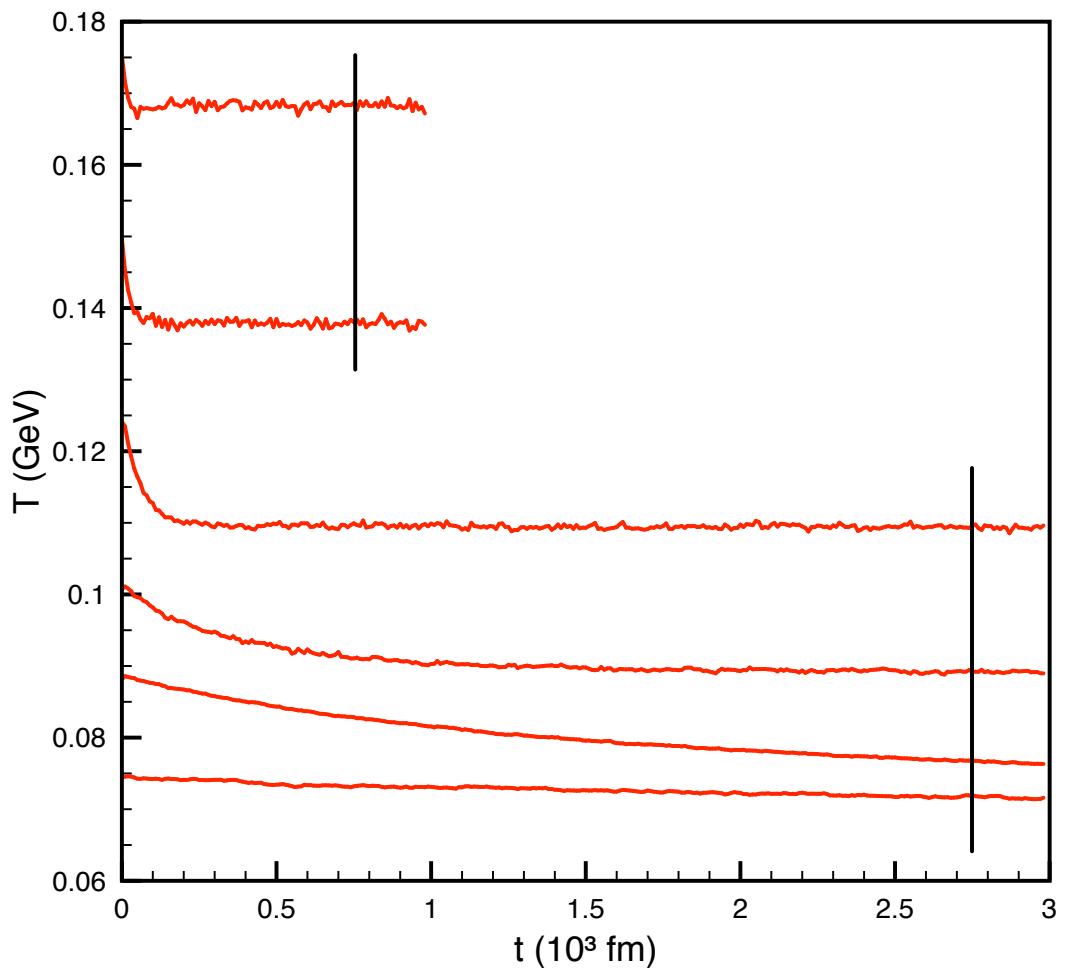
In the range of typical
hadron gas equation of state

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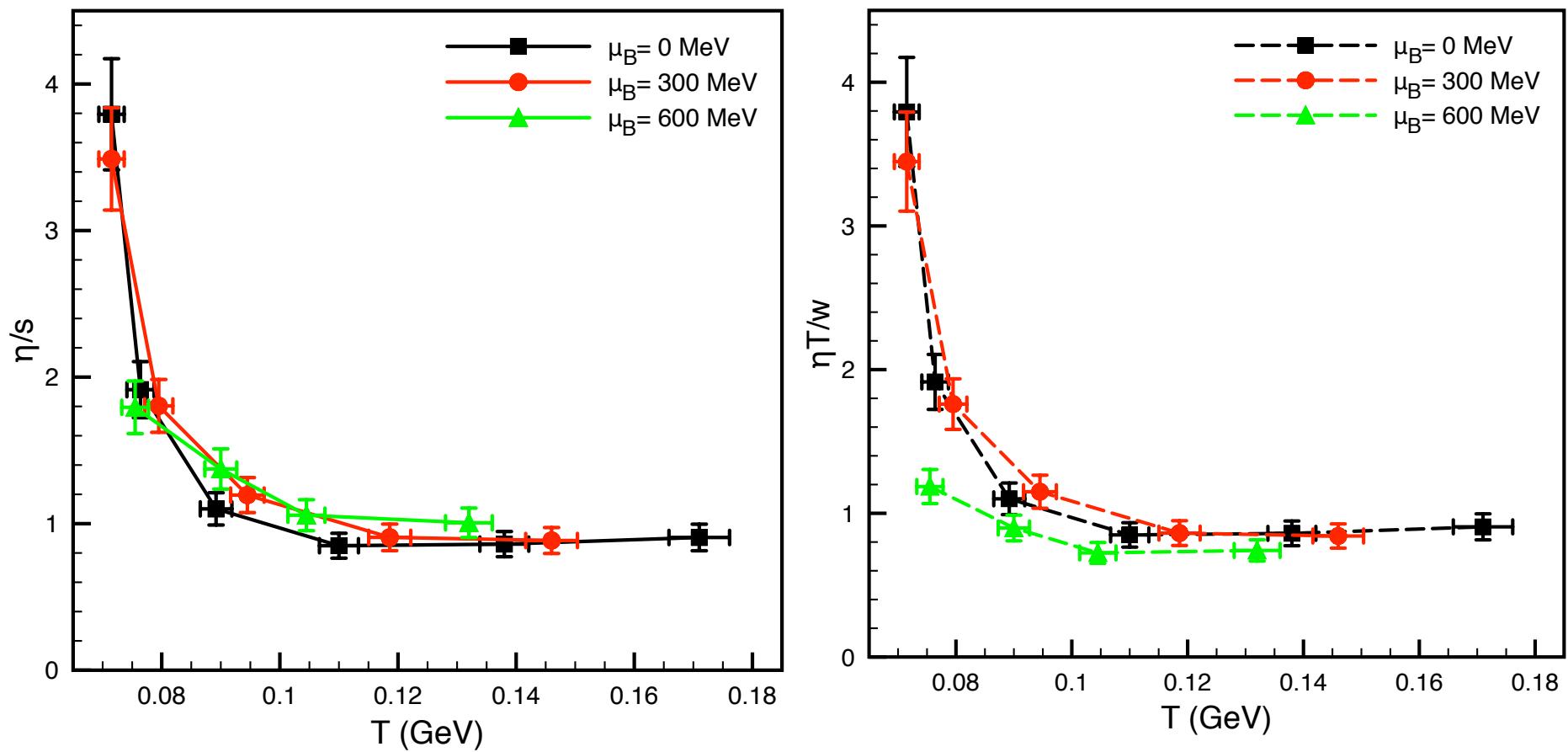
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Hadron Gas (HG)

- All particles and resonances initialized to thermal multiplicities
- Must wait for equilibration and compute T, μ once in equilibrium from most abundant particles
 - T fitted from weighted momentum spectra of π, K & N
 - μ_B obtained from N / anti- N ratio

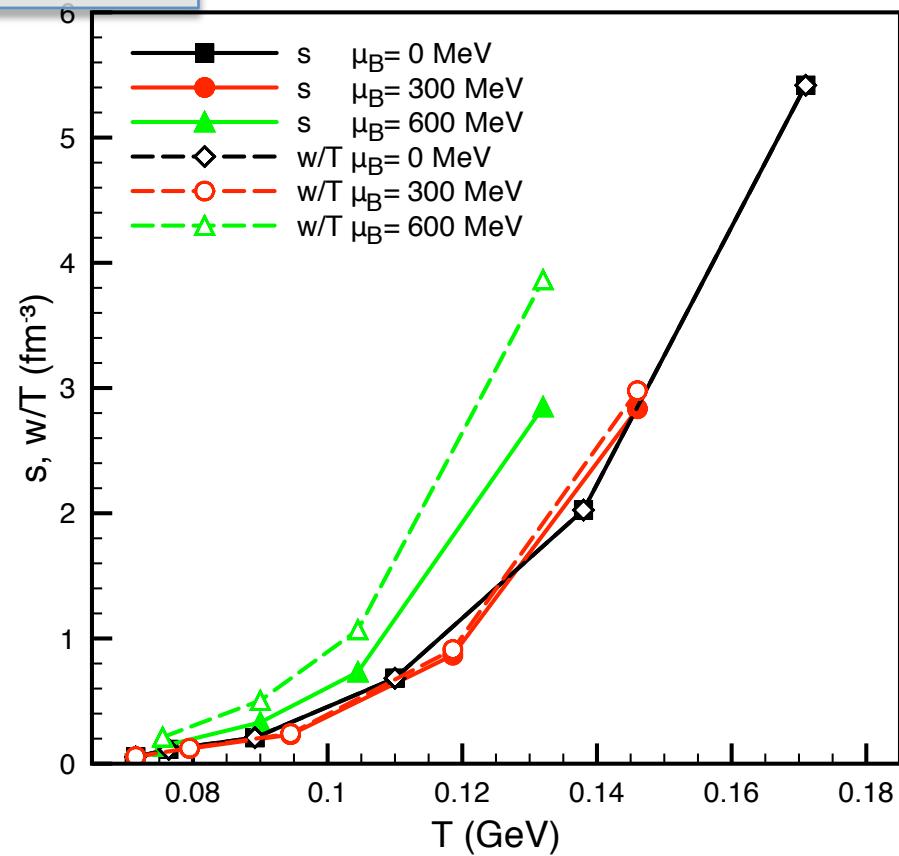
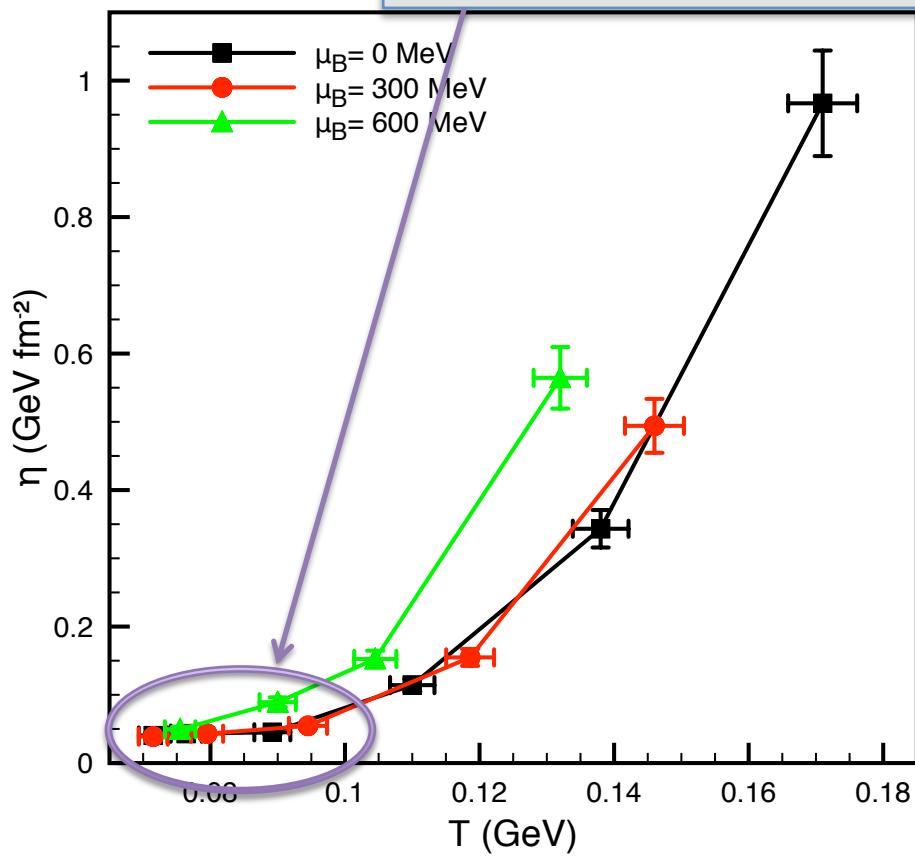


η/s and $\eta T/w$

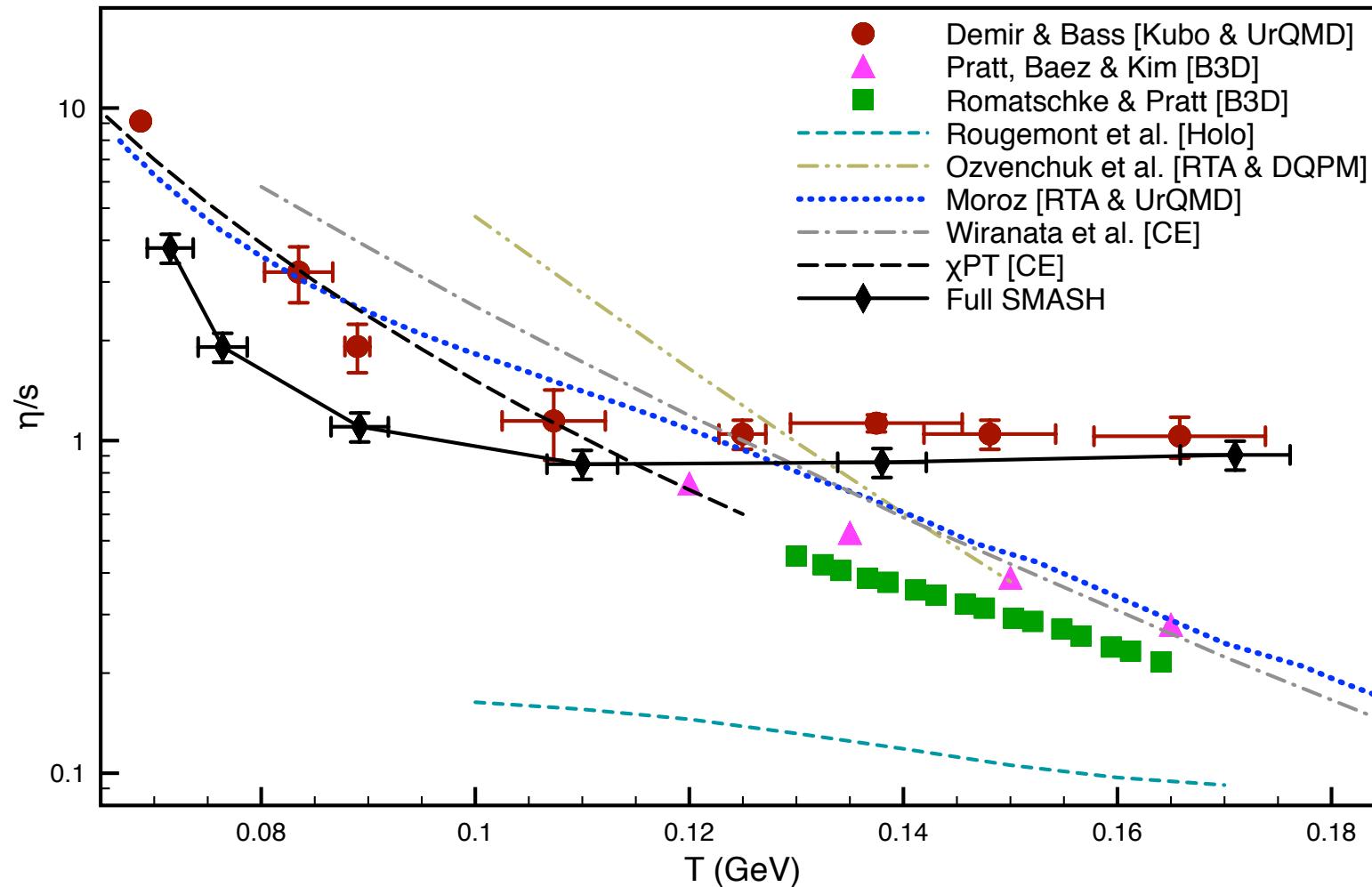


η , s and w/T

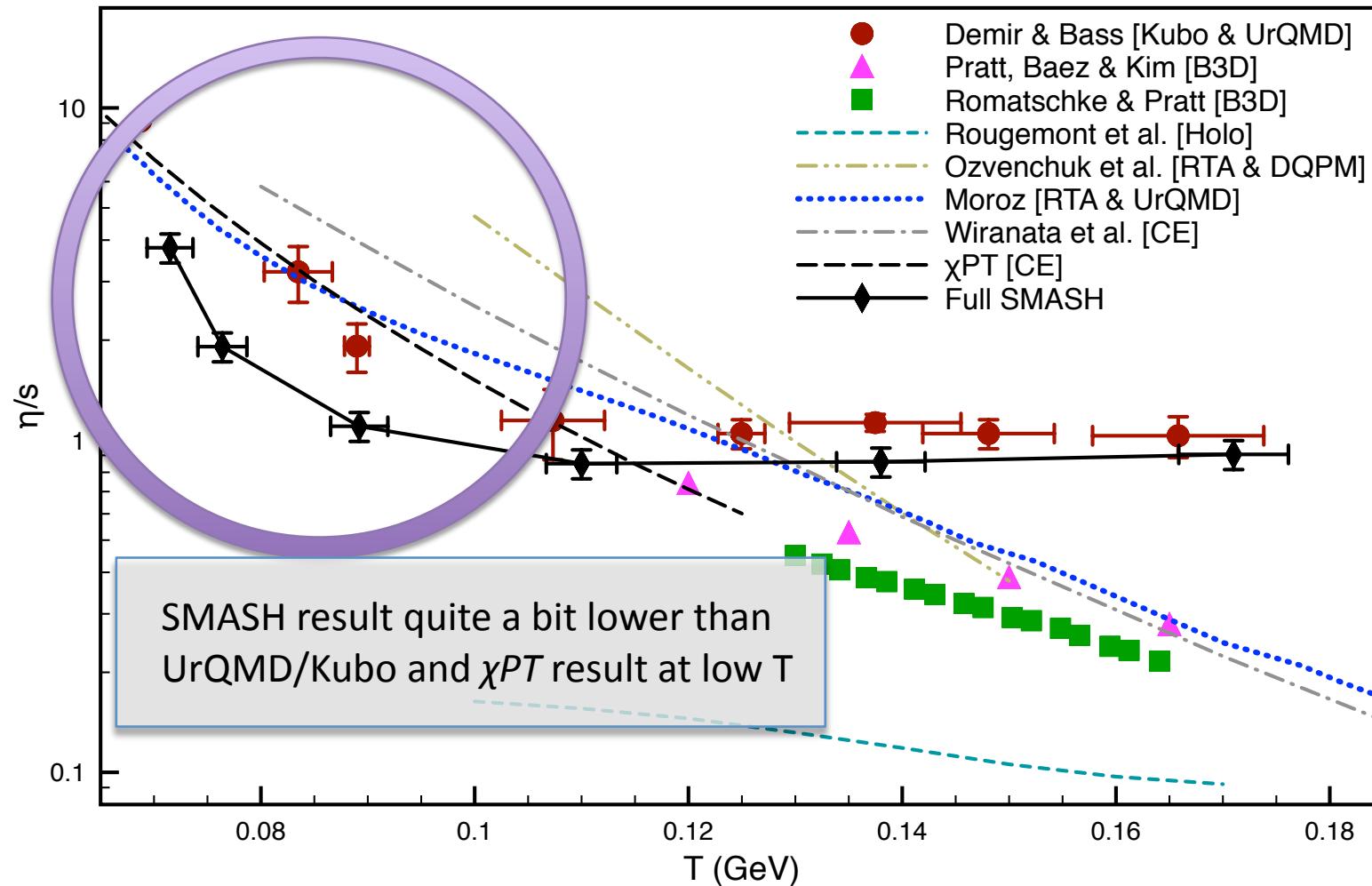
Viscosity decreases slower at small temperatures; explains rise of η/s



HG: Viscosity Comparison

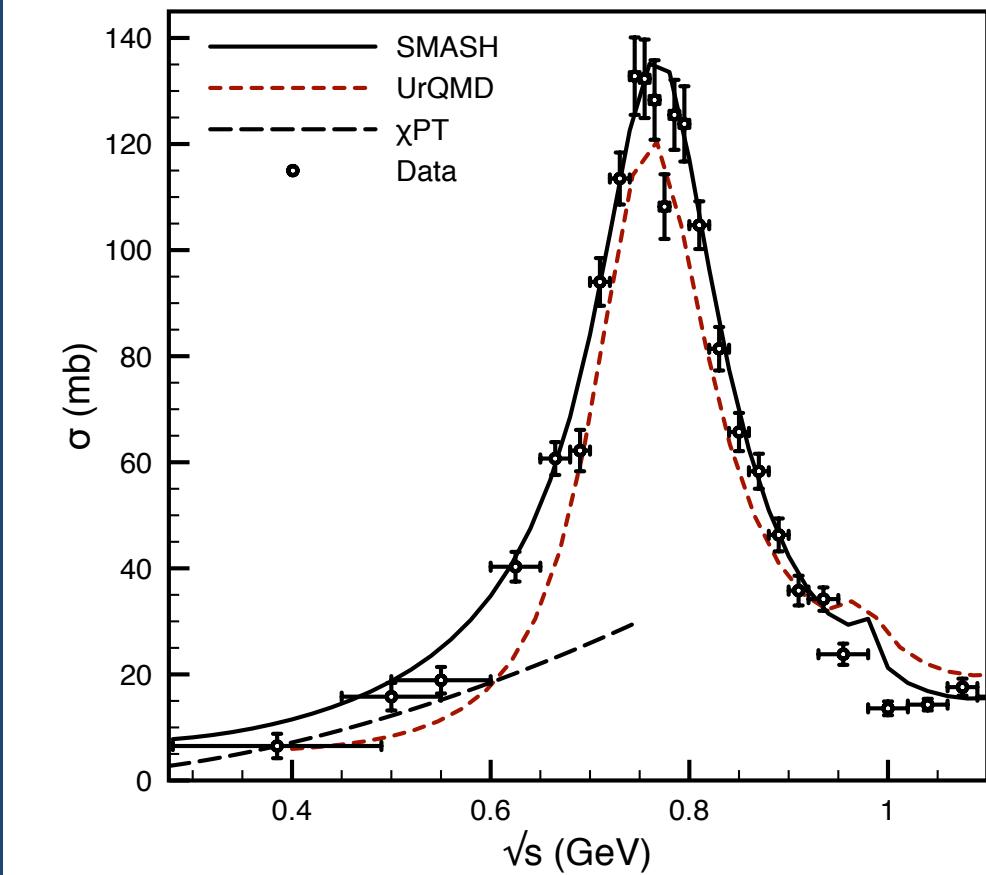


HG: Viscosity Comparison

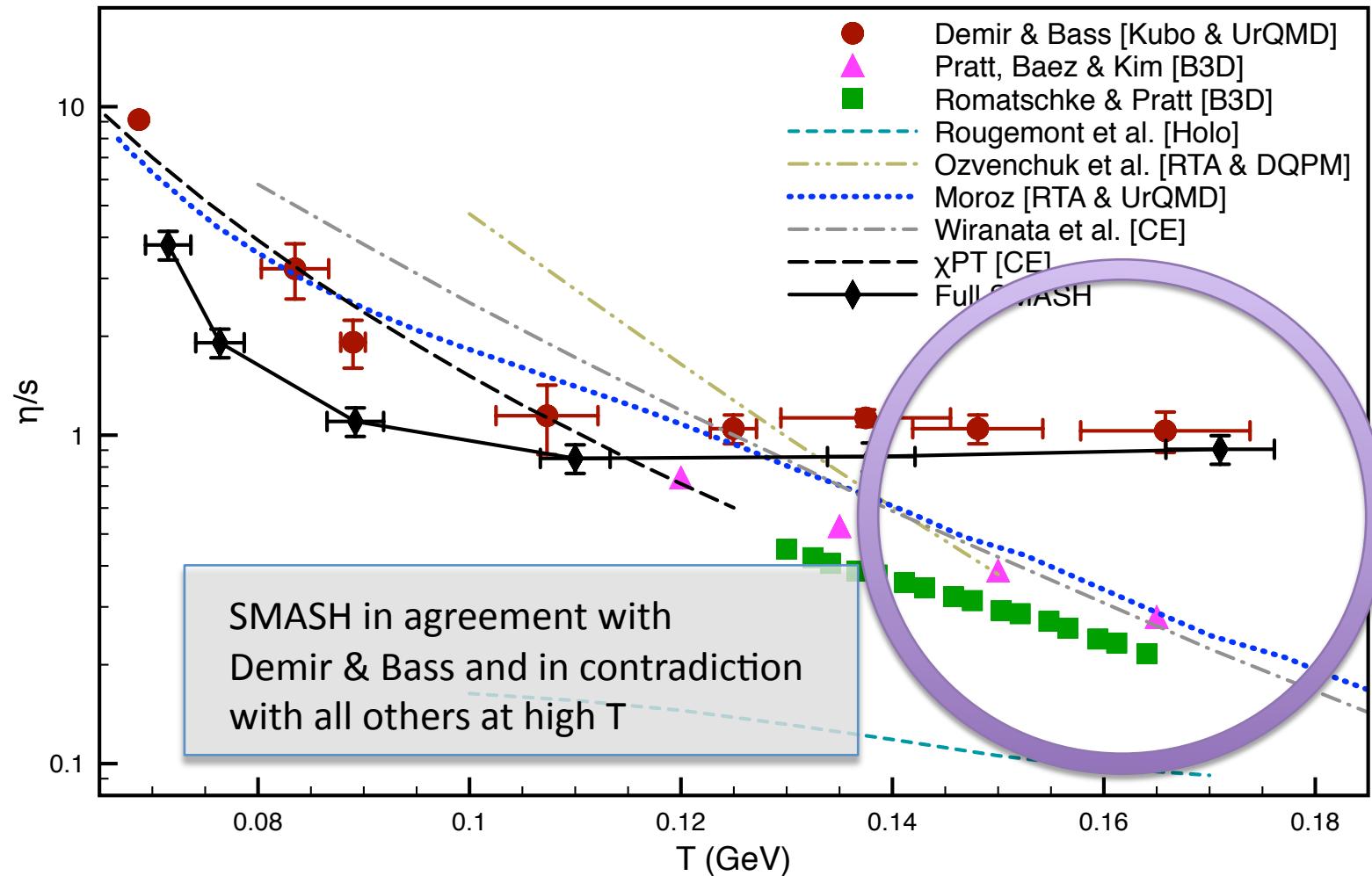


Low temperature η/s

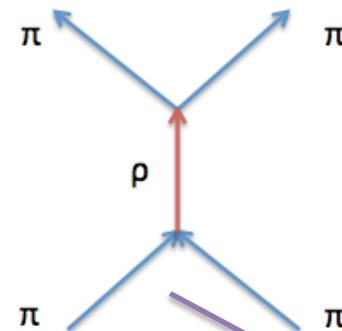
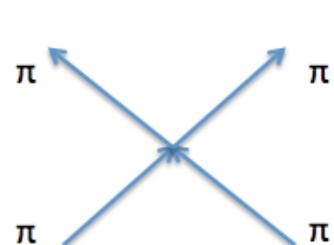
- Low temperature hadron gas is composed almost exclusively of pions
- $\pi\text{-}\pi$ cross-section is then most relevant
 - At very low energy, SMASH much higher than UrQMD/ χ PT
 - χ PT includes angular dependence, UrQMD&SMASH don't; increases viscosity by factor up to $5/3$ for ρ resonance



HG: Viscosity Comparison

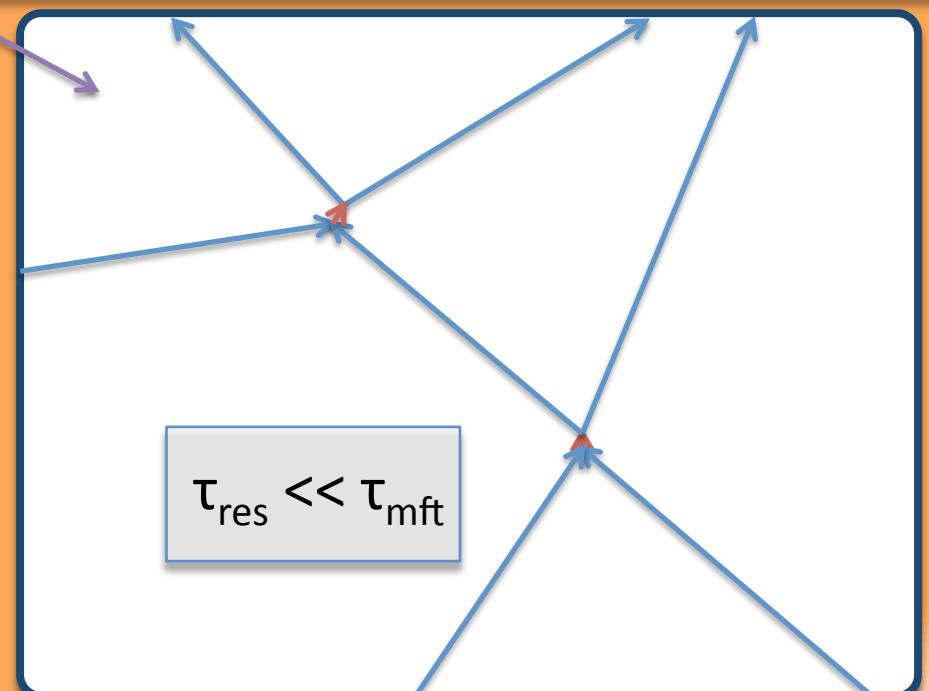
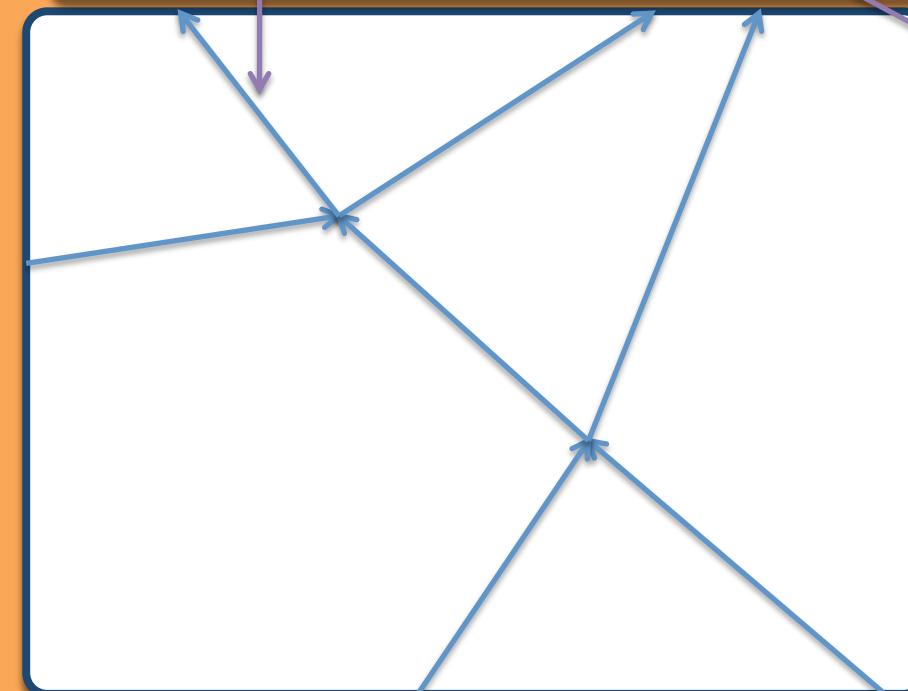


High temperature η/s : Resonance lifetimes

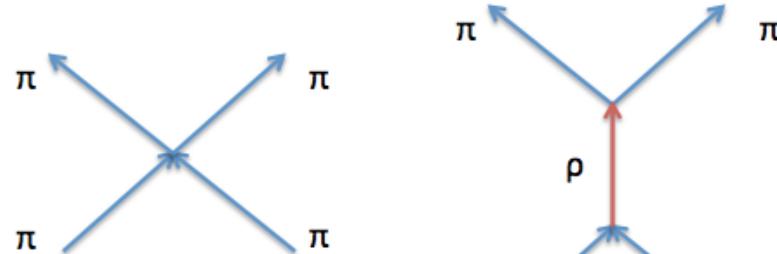


Must look at the microscopic picture from different descriptions

At low T and density:

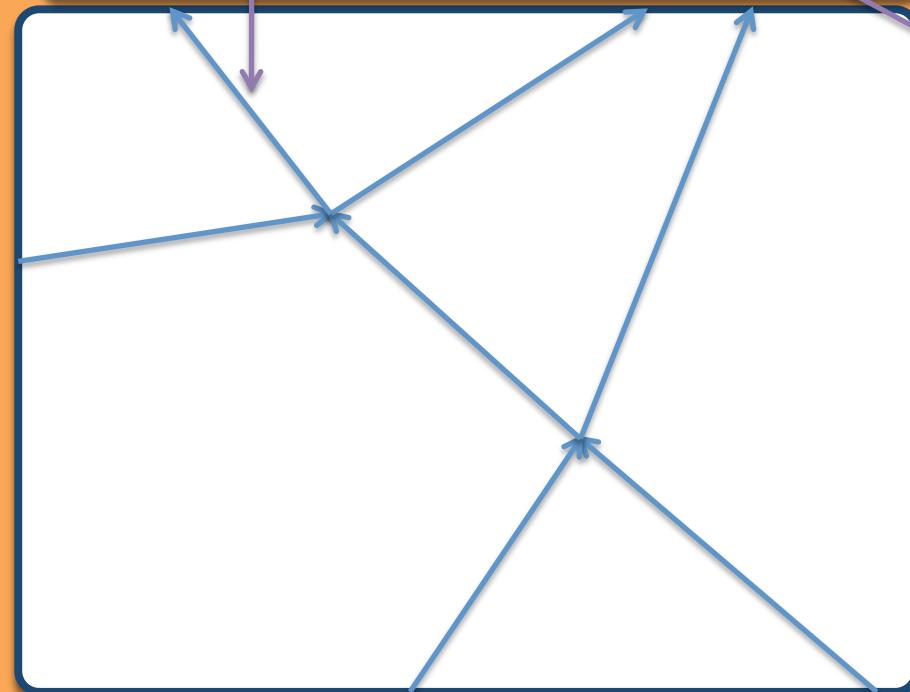


High temperature η/s : Resonance lifetimes

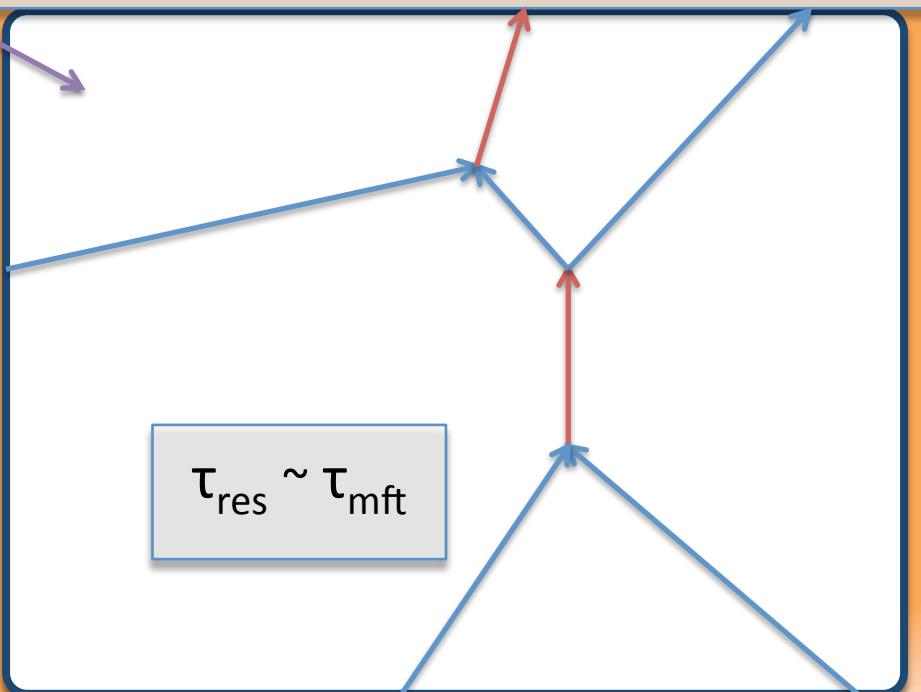


Must look at the microscopic picture from different descriptions

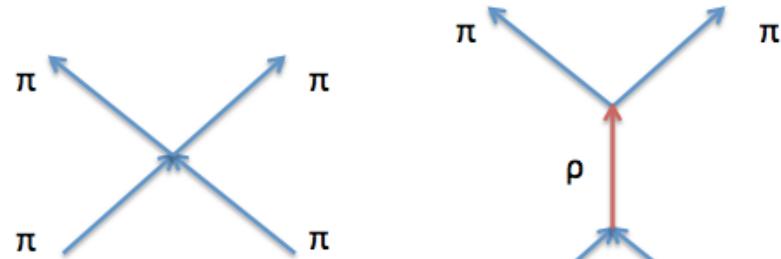
At high T and density:



$$\tau_{\text{res}} \sim \tau_{\text{mft}}$$



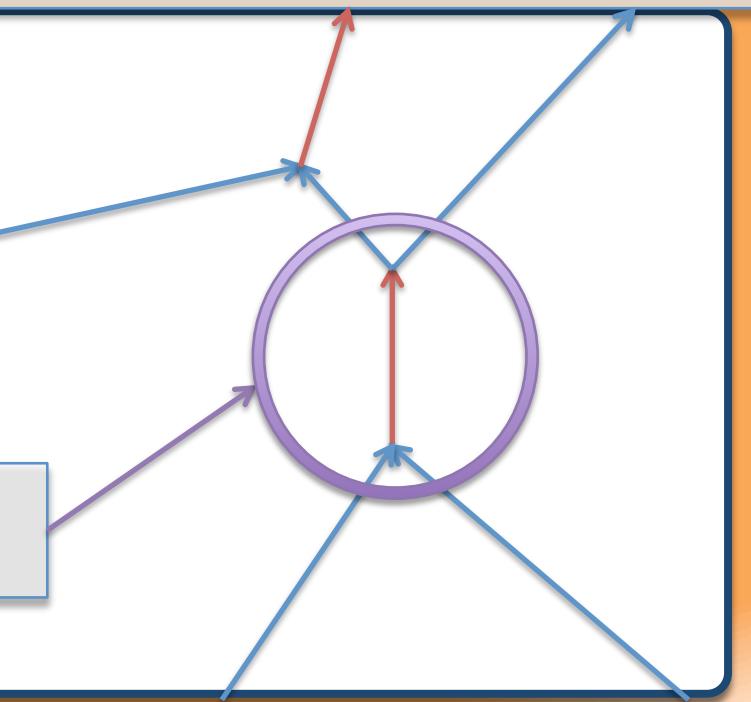
High temperature η/s : Resonance lifetimes



Must look at the microscopic picture from different descriptions

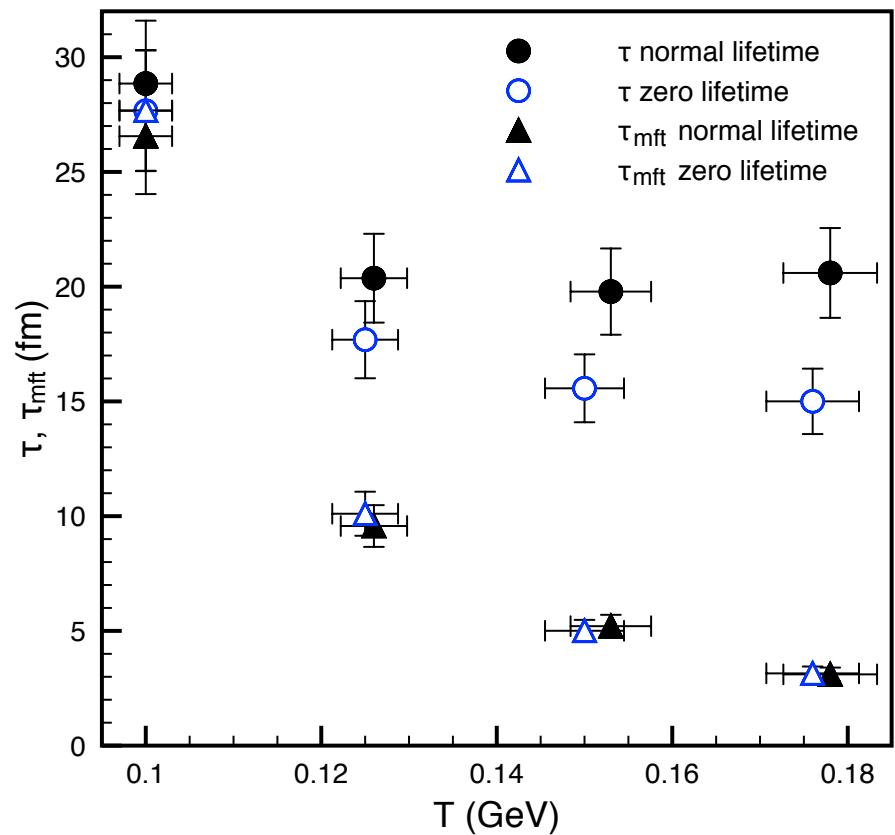
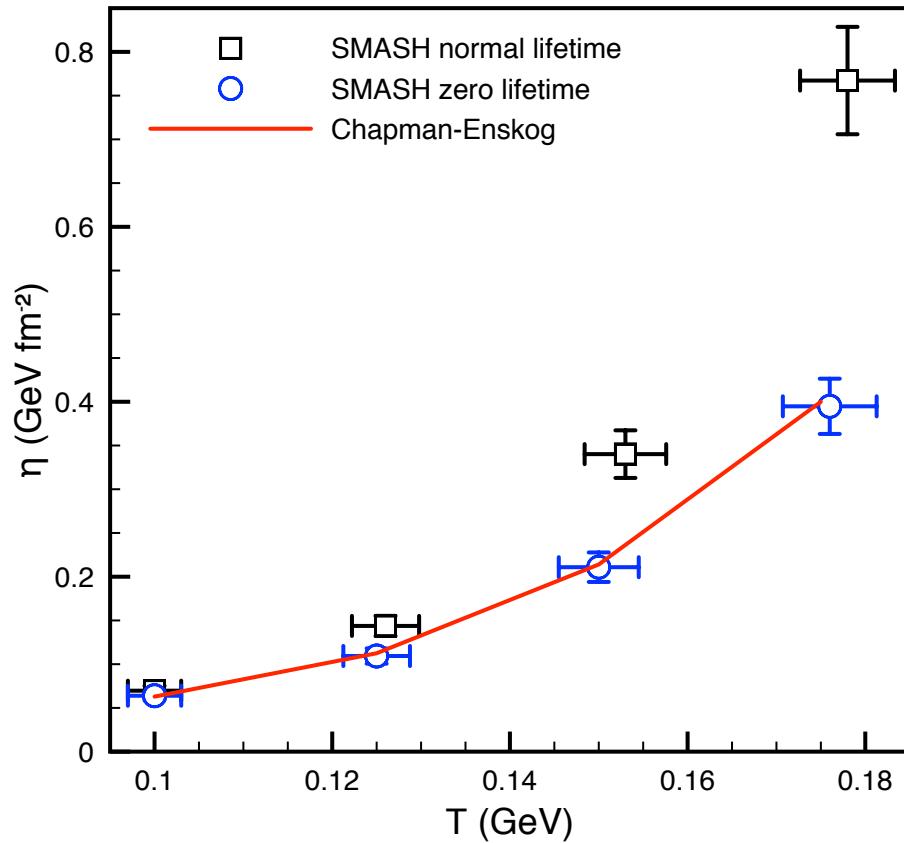
At high T and density:

Momentum transport is *delayed* until resonances decay!



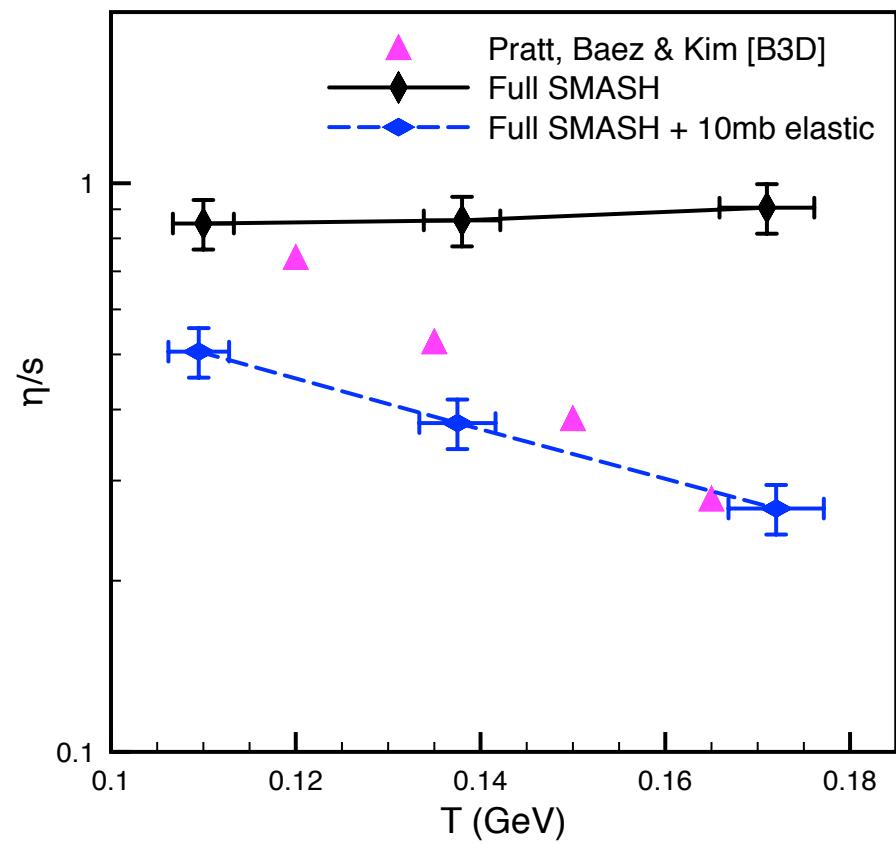
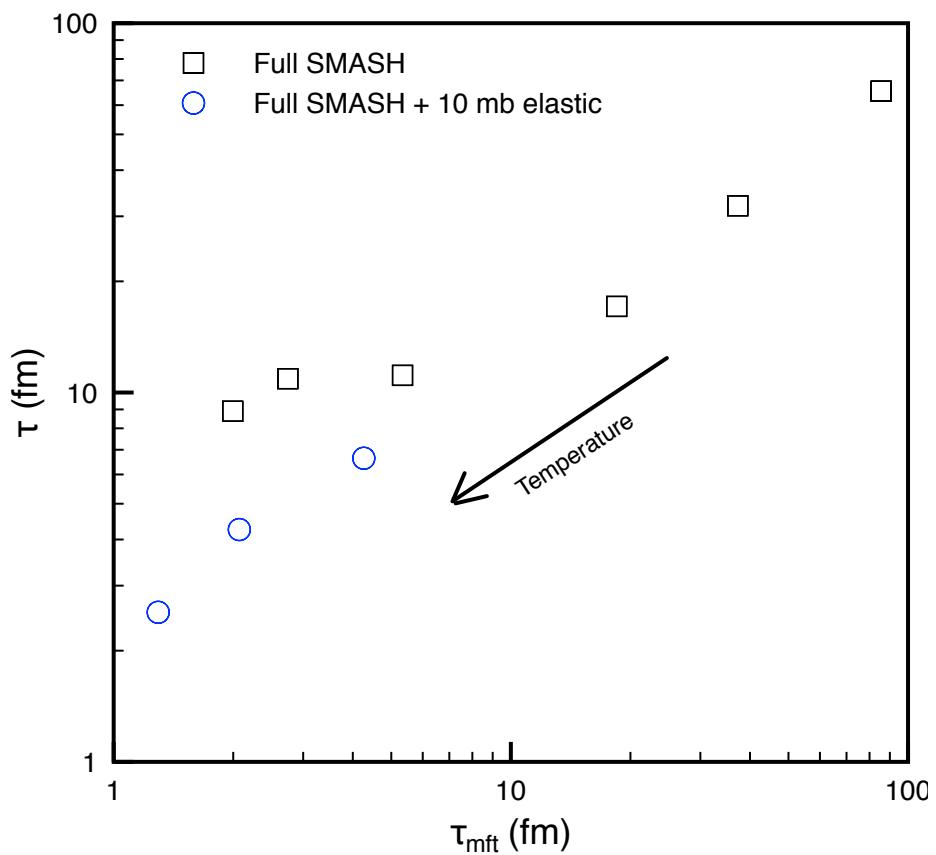
π - ρ : Zero lifetimes vs relaxation time

Large part of the difference explained from eliminating lifetimes



Effect of many non-resonant interactions

Introduce a constant elastic cross-section between all particles to add many non-resonant interactions



Outline

1. Introduction: Viscosity of the hadron gas
2. Transport
 - SMASH
3. Methodology
 - Viscosity considerations
 - Green-Kubo formalism
 - Test case #1: Constant isotropic cross-section
 - Test case #2: Energy-dependent cross-section
 - Entropy considerations
4. Results
 - Full hadron gas viscosity
 - Comparison & discussion
5. Conclusion

Summary & Outlook

- **Investigated temperature, cross-section and mass dependence of the shear viscosity in an elastic pion box**
 - Very good agreement with Chapman-Enskog approximation (within 3%)
 - Systematics show that method is robust to variation of technical parameters
- **Full hadron gas η/s calculated**
 - Slightly lower than other calculations at low T because of large π - π cross-section
 - High T discrepancy explained by looking at microscopic details of resonance modelling; finite lifetime increases viscosity
 - Could be used to constrain the treatment of resonances
- **Outlook:**
 - More rigorous analysis of the dependence between τ and τ_{res} needed
 - Investigation of angular dependent interactions on viscosity
 - At temperatures close to the phase transition, inclusion of multi-particle interaction will probably play a role, and needs to be investigated
 - Other transport coefficients (electrical conductivity, bulk viscosity, etc.)