Canonical suppression & enhancement using Master equations

Moritz Wortmann

Institut für Theoretische Physik, Goethe Universität Frankfurt am Main, Germany



MHD/Transport Meeting, $10^{\rm th}$ July 2025



- 2 Master equations
- 3 Canonical suppression in $\pi\pi \leftrightarrow K\overline{K}$ transition
- 4 Canonical enhancement in $c\bar{c} \leftrightarrow J/\Psi$ transition

5 Outlook!

- Particle number important observable
- Statistical calculations relevant for comparing with experiment and microscopical simulations
 - Differences between canonical and grand canonical approach?
- Usage of Master equation:
 - Numerically rather simple
 - Adaptable to different transitions

Differential equation providing the **time development of the probabilities** in a system.

General structure of a Master equation $\frac{dP_i}{dt} = \sum_j \left(\alpha_{j \to i} P_j - \alpha_{i \to j} P_i \right)$ (1)

 P_j probability of the system being in state j. $a_{j \rightarrow i}$ transition rate from state j to state i.

Simulation setup (canonical suppression)



Figure: Simulation setup for Kaon production [Phys. Rev. C, 74:034902]

Canonical suppression



Figure: Canonical suppression [Phys. Rev. C, 74:034902]

- Canonical approach demands exact conservation
 - Reduces the phase space for transition
- Suppressed production of particles in canonical approach compared to grand-canonical

Suppression factor

$$N_{eq}^{C} = N_{eq}^{GC} \frac{I_1(2N_{eq}^{GC})}{I_0(2N_{eq}^{GC})} = N_{eq}^{GC} \eta$$

$$\eta \equiv \frac{N_{eq}^C}{N_{eq}^{GC}} = \frac{I_1(2N_{eq}^{GC})}{I_0(2N_{eq}^{GC})}$$
(2)

 $P_i \rightarrow P_n$ probability for a system with *n* Kaons and Antikaons Transitions are of type $\alpha_{n \rightarrow n+1}$, $\alpha_{n \rightarrow n-1}$, a particle is **produced** or **decays**

Adapted Master equation

$$\frac{\mathrm{d}P_n}{\mathrm{d}t} = \alpha_{n-1 \to n} P_{n-1} + \alpha_{n+1 \to n} P_{n+1} - (\alpha_{n \to n+1} + \alpha_{n \to n-1}) P_n \tag{3}$$

 α can be divided into two parts:

Transition rate of one pair

 Γ_G rate for $\pi\pi \to K\overline{K}$ formation $(\alpha_{n\to n+1})$ Γ_L rate for $K\overline{K} \to \pi\pi$ decay $(\alpha_{n\to n-1})$

$$\Gamma \equiv \frac{\Gamma_G}{\Gamma_L} = \frac{Z_K^1 Z_{\overline{K}}^1}{Z_\pi^1 Z_\pi^1}$$

Number of possible transitions

 N_{pair}^2 possible combinations.

(4)

Time development of Kaon number probability

$$\frac{\mathrm{d}P_n}{\mathrm{d}t} = \Gamma_G \langle N_\pi \rangle \langle N_\pi \rangle [P_{n-1} - P_n] - \Gamma_L [n^2 P_n - (n+1)^2 P_{n+1}]
\frac{\mathrm{d}P_n}{\mathrm{d}\tau} = \Gamma \langle N_\pi \rangle \langle N_\pi \rangle [P_{n-1} - P_n] - [n^2 P_n - (n+1)^2 P_{n+1}]; \quad \tau = \Gamma_L t$$
(5)

Kaon number in equilibrium

$$\left\langle N_{K}^{eq}\right\rangle = \left\langle N_{\overline{K}}^{eq}\right\rangle = \sum_{n} nP_{n}$$

Comparing the results



Figure: Microscopical simulation

Naomi's simulation



Figure: Simulation setup (Design: Ian David Ortiz Salcedo)

- Initial charm-anticharm pairs are placed in a box
- Charm quarks move as Brownian particles in a thermal medium.
- Charm–anticharm pairs interact via a potential
 - Pair is bound if total relative energy is negative
- Number of charm and anticharm strictly conserved at any time

Microscopic Simulation

- Fixed number of initial charm pairs
- Charm is conserved exactly
- Complies with the canonical ensemble

Statistical Hadronization Model

- Framework describing hadron production in high energy collisions
 - Hadrons formed from thermalized QGP
- Multiplicity is derived from grand canonical ensemble
- Charm is not conserved exactly, only on average

Canonical enhancement



Figure: Canonical enhancement (data from [PhysRevD.111.074012]) Canonical approach:

- Particles always occur in pairs
- Grand-canonical approach
 - Particle and antiparticle number only the same on average

Low number $\langle N_c \rangle$ of initial charm present Numerical example: $\langle N_c \rangle = 0.1$

Canonical Ensemble	Grand-Canonical Ensemble
In every event: $N_c = N_{\bar{c}}$	Only on average : $\langle N_c \rangle = \langle N_{\bar{c}} \rangle$
$P((N_c \wedge N_{\bar{c}}) \ge 1) = P(N_c \ge 1) \approx 0.1$	$P(N_{\bar{c}} \ge 1) = P(N_c \ge 1) \approx 0.1$ $P((N_c \land N_{\bar{c}}) \ge 1) \approx 0.01$

 $P_n^{J/\Psi}$ probability for a system with $n J/\Psi$ P_m^c probability for a system with m c and \bar{c} Transitions are of type $\alpha_{n \to n+1}$, $\alpha_{n \to n-1}$, a particle is **produced** or **decays**

Adapted Master equation

$$\frac{\mathrm{d}P_{n}^{J/\Psi}}{\mathrm{d}t} = \alpha_{n-1\to n}P_{n-1}^{J/\Psi} + \alpha_{n+1\to n}P_{n+1}^{J/\Psi} - (\alpha_{n\to n+1} + \alpha_{n\to n-1})P_{n}^{J/\Psi}$$

$$\frac{\mathrm{d}P_{m}^{c}}{\mathrm{d}t} = \alpha_{m-1\to m}P_{m-1}^{c} + \alpha_{m+1\to m}P_{m+1}^{c} - (\alpha_{m\to m+1} + \alpha_{m\to m-1})P_{m}^{c}$$
(6)

Transition rate of one pair

 Γ_G rate for $c\bar{c} \to J/\Psi$ formation $(\alpha_{n \to n+1}, \alpha_{m \to m-1})$ Γ_L rate for $J/\Psi \to c\bar{c}$ decay $(\alpha_{n \to n-1}, \alpha_{m \to m+1})$

$$\Gamma \equiv \frac{\Gamma_G}{\Gamma_L} = \frac{Z^1_{J/\Psi}}{Z^1_c Z^1_{\bar{c}}}$$

Number of possible transitions

For $c\bar{c} \rightarrow J/\Psi$: N_c^2 possible combinations. For $J/\Psi \rightarrow c\bar{c}$: $N_{J/\Psi}$ possibilities. (7)

Master equation for $c\bar{c} \leftrightarrow J/\Psi$

$$\frac{\mathrm{d}P_{n}^{J/\Psi}}{\mathrm{d}t} = \Gamma_{G} \left(P_{n-1}^{J/\Psi} - P_{n}^{J/\Psi} \right) \left(\sum_{m=0}^{\infty} P_{m}^{c} m^{2} \right) - \Gamma_{L} \left(P_{n}^{J/\Psi} n - P_{n+1}^{J/\Psi} (n+1) \right)$$
$$\frac{\mathrm{d}P_{m}^{c}}{\mathrm{d}t} = \Gamma_{L} \left(\sum_{n=0}^{\infty} P_{n}^{J/\Psi} n \right) \left(P_{m-1}^{c} - P_{m}^{c} \right) - \Gamma_{G} \left(P_{m}^{c} m^{2} - P_{m+1}^{c} (m+1)^{2} \right)$$
(8)

Comparing the results



Figure: Microscopic model, the master equations and SHM.

Summary

Low number of particles: one has to consider canonical effects

- depending on the setup production is suppressed or enhanced
- Master equations can be used for the canonical approach
 - Seem to align with microscopic simulations

Outlook

- Including other charmonium states
- Other transitions e.g. $b\overline{b} \to \Upsilon$