

# Dissipative fluid dynamics in the relaxation-time approximation for an ideal gas of massive particles

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## Transport coefficients of second-order relativistic fluid dynamics in the relaxation-time approximation

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We derive the transport coefficients of second-order fluid dynamics with 14 dynamical moments using the method of moments and the Chapman-Enskog method in the relaxation-time approximation for the collision integral of the relativistic Boltzmann equation. Contrary to results previously reported in the literature, we find that the second-order transport coefficients derived using the two methods are in perfect agreement. Furthermore, we show that, unlike in the case of binary hard-sphere interactions, the diffusion-shear coupling coefficients  $\ell_{V\pi}$ ,  $\lambda_{V\pi}$ , and  $\tau_{V\pi}$  actually diverge in some approximations when the expansion order  $N_\ell \rightarrow \infty$ . Here we show how to circumvent such a problem in multiple ways, recovering the correct transport coefficients of second-order fluid dynamics with 14 dynamical moments. We also validate our results for the diffusion-shear coupling by comparison to a numerical solution of the Boltzmann equation for the propagation of sound waves in an ultrarelativistic ideal gas.

## Introduction and Conclusions II.

## Relativistic second-order dissipative and anisotropic fluid dynamics in the relaxation-time approximation for an ideal gas of massive particles

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In this paper, we study all transport coefficients of second-order dissipative fluid dynamics derived by V. E. Ambruş *et al.* [*Phys. Rev. D* **106**, 076005 (2022)] from the relativistic Boltzmann equation in the relaxation-time approximation for the collision integral. These transport coefficients are computed for a classical ideal gas of massive particles, with and without taking into account the conservation of intrinsic quantum numbers. Through rigorous comparison between kinetic theory, second-order dissipative fluid dynamics, and leading-order anisotropic fluid dynamics for a  $(0 + 1)$ -dimensional boost-invariant flow scenario, we show that both fluid-dynamical theories describe the early far-from-equilibrium stage of the expansion reasonably well.

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# The relativistic Boltzmann equation I.

The Boltzmann equation describes the space-time evolution of the single-particle distribution function  $f(x^\mu, k^\mu) = f_{\mathbf{k}}$

## The relativistic Boltzmann equation

$$k^\mu \partial_\mu f_{\mathbf{k}} \equiv C[f_{\mathbf{k}}] = \frac{1}{2} \int dK' dP dP' W_{\mathbf{k}\mathbf{k}' \rightarrow \mathbf{p}\mathbf{p}'} \left( f_{\mathbf{p}} f_{\mathbf{p}'} \tilde{f}_{\mathbf{k}} \tilde{f}_{\mathbf{k}'} - f_{\mathbf{k}} f_{\mathbf{k}'} \tilde{f}_{\mathbf{p}} \tilde{f}_{\mathbf{p}'} \right), \quad (1)$$

$k^\mu = (k^0, \mathbf{k})$  is the four-momenta of particles with mass  $m_0 = \sqrt{k^\mu k_\mu}$ , where  $dK = g d^3\mathbf{k} / [(2\pi)^3 k^0]$ , while  $\tilde{f}_{\mathbf{k}} = 1 - a f_{\mathbf{k}}$ , with  $a = 0/a = 1/a = -1$  for Boltzmann/Fermi/Bose statistics. The invariant transition rate is

$$W_{\mathbf{k}\mathbf{k}' \rightarrow \mathbf{p}\mathbf{p}'} \equiv \frac{s}{g^2} (2\pi)^6 \frac{d\sigma(\sqrt{s}, \Omega)}{d\Omega} \delta(k^\mu + k'^\mu - p^\mu - p'^\mu), \quad (2)$$

which for an isotropic and energy independent diff. cross-section

## the hard-sphere approximation

$$\sigma_T \equiv 2\pi \frac{d\sigma(\sqrt{s}, \Omega)}{d\Omega} = \frac{1}{n_0 \lambda_{\text{mfP}}}. \quad (3)$$

# The relativistic Boltzmann equation II.

## Local thermal equilibrium - Jüttner distribution

$$f_{\mathbf{k}} \rightarrow f_{0\mathbf{k}} \equiv [\exp(-\alpha_0 + \beta_0 E_{\mathbf{k}}) + a]^{-1}, \quad (4)$$

Not a "true" solution, where  $\alpha_0 = \mu_0/T_0$ ,  $\beta_0 = 1/T_0$  and  $E_{\mathbf{k}} = k^\mu u_\mu$ . The solution

$$f_{\mathbf{k}} \equiv f_{0\mathbf{k}} + \delta f_{\mathbf{k}}, \quad (5)$$

where  $\delta f_{\mathbf{k}}$  is the non-equilibrium correction.

## The Anderson-Witting approximation to the collision integral

$$C[f_{\mathbf{k}}] \equiv -k^\mu u_\mu \frac{f_{\mathbf{k}} - f_{0\mathbf{k}}}{\tau_R} = -E_{\mathbf{k}} \frac{\delta f_{\mathbf{k}}}{\tau_R}, \quad (6)$$

where the relaxation time  $\tau_R(x^\mu)$  is a momentum-independent parameter proportional to the mean free time between collisions. This is the relativistic generalization of the BGK model;  $C_{BGK}[f_{\mathbf{k}}] \equiv -\frac{f_{\mathbf{k}} - f_{0\mathbf{k}}}{\tau_R}$ , used in Lattice Boltzmann methods.

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# Fluids dynamics from the Boltzmann equation

The momentum-space integral of the distribution function  $f_{\mathbf{k}}$  with the four-momentum and the energy of particles forming a rank- $\ell$  tensor defines the moment

## Rank- $\ell$ tensor moment

$$M_r^{\mu_1 \dots \mu_\ell} \equiv \langle E_{\mathbf{k}}^r k^{\mu_1} \dots k^{\mu_\ell} \rangle = \int dK E_{\mathbf{k}}^r k^{\mu_1} \dots k^{\mu_\ell} f_{\mathbf{k}} . \quad (7)$$

## The particle four-current and the energy-momentum tensor

$$N^\mu \equiv M_0^\mu = \int dK k^\mu f_{\mathbf{k}} = \int dK k^\mu (f_{0\mathbf{k}} + \delta f_{\mathbf{k}}) , \quad (8)$$

$$T^{\mu\nu} \equiv M_0^{\mu\nu} = \int dK k^\mu k^\nu f_{\mathbf{k}} = \int dK k^\mu k^\nu (f_{0\mathbf{k}} + \delta f_{\mathbf{k}}) . \quad (9)$$

## Moments of the relativistic Boltzmann equation (RTA)

$$\begin{aligned} \int dK k^{\lambda_1} \dots k^{\lambda_\ell} [k^\mu \partial_\mu f_{\mathbf{k}}] &\equiv \partial_\mu M_0^{\mu\lambda_1 \dots \lambda_\ell} = C^{\lambda_1 \dots \lambda_\ell} [f_{\mathbf{k}}] , \\ &= -\frac{1}{T_R} \left( M_1^{\lambda_1 \dots \lambda_\ell} - \int dK E_{\mathbf{k}} k^{\lambda_1} \dots k^{\lambda_\ell} f_{0\mathbf{k}} \right) \end{aligned} \quad (10)$$

Note that the particle number and particle four-momenta are collision invariants, i.e.,  $C = \int dK C[f] = 0$  and  $C^\alpha = \int dK k^\alpha C[f] = 0$

### The particle number conservation

$$\begin{aligned}\partial_\mu N^\mu &\equiv \partial_\mu \int dK k^\mu f_{\mathbf{k}} = -\frac{1}{\tau_R} \int dK E_{\mathbf{k}} (f_{\mathbf{k}} - f_{0\mathbf{k}}) \\ &= -\frac{1}{\tau_R} (n - n_0) \equiv 0,\end{aligned}\tag{11}$$

### The energy-momentum conservation

$$\begin{aligned}u_\nu \partial_\mu T^{\mu\nu} &\equiv u_\nu \partial_\mu \int dK k^\mu k^\nu f_{\mathbf{k}} = -\frac{1}{\tau_R} \int dK E_{\mathbf{k}}^2 (f_{\mathbf{k}} - f_{0\mathbf{k}}) \\ &= -\frac{1}{\tau_R} (e - e_0) \equiv 0,\end{aligned}\tag{12}$$

$$\begin{aligned}\Delta_\nu^\lambda \partial_\mu T^{\mu\nu} &\equiv \Delta_\nu^\lambda \partial_\mu \int dK k^\mu k^\nu f_{\mathbf{k}} = -\frac{1}{\tau_R} \int dK E_{\mathbf{k}} k^{\langle\lambda} (f_{\mathbf{k}} - f_{0\mathbf{k}}) \\ &= -\frac{1}{\tau_R} W^\lambda \equiv 0,\end{aligned}\tag{13}$$

### Landau matching conditions, and Landau frame

$$n = n_0, \quad e = e_0, \tag{14}$$

$$W^\lambda = 0, \tag{15}$$

## Ideal Fluids I.

Conservation laws for a simple (single component) perfect fluid (no dissipation)

$$\begin{aligned} \partial_\mu N_0^\mu &= 0 & \text{charge conservation} & \Rightarrow \mathbf{1 \text{ eq.}} \\ \partial_\mu T_0^{\mu\nu} &= 0 & \text{energy-momentum conservation} & \Rightarrow \mathbf{4 \text{ eqs.}} \\ u_\nu \partial_\mu T_0^{\mu\nu} &= 0, & \Delta_\nu^\lambda \partial_\mu T_0^{\mu\nu} &= 0 \end{aligned}$$

Perfect fluid decomposition with respect to  $u^\mu$

$$\begin{aligned} N_0^\mu &= n_0 u^\mu \\ T_0^{\mu\nu} &= \epsilon_0 u^\mu u^\nu - p_0 \Delta^{\mu\nu} \\ n_0 &= N_0^\mu u_\mu & \text{(net)charge density} \\ \epsilon_0 &= T_0^{\mu\nu} u_\mu u_\nu & \text{energy density} \\ P_0 &= -\frac{1}{3} \Delta_{\mu\nu} T_0^{\mu\nu} & \text{equilibrium pressure} \end{aligned}$$

- The time-like normalized flow velocity is  $u^\mu(t, \vec{x}) = \gamma(1, \mathbf{v})$ , where  $u^\mu u_\mu = 1$
- The projection tensor  $\Delta^{\mu\nu} = g^{\mu\nu} - u^\mu u^\nu$ , where  $g^{\mu\nu} = \text{diag}(1, -1, -1, -1)$
- We have **5** equations for **6** unknowns  $n_0(1)$ ,  $\epsilon_0(1)$ ,  $p_0(1)$  and  $u^\mu(3)$ . Closed by an Equation of State (EoS)  $P_0 = P_0(\epsilon_0, n_0)$ .

## Dissipative Fluids I.

Conservation laws for a simple (single component) dissipative fluid

$$\partial_\mu N^\mu = 0 \quad \text{charge conservation} \quad \Rightarrow \mathbf{1 \text{ eq.}}$$

$$\partial_\mu T^{\mu\nu} = 0 \quad \text{energy-momentum conservation} \quad \Rightarrow \mathbf{4 \text{ eqs.}}$$

General decomposition  $N^\mu = N_0^\mu + \delta N^\mu$  and  $T^{\mu\nu} = T_0^{\mu\nu} + \delta T^{\mu\nu}$ 

$$N^\mu = n u^\mu + V^\mu$$

$$T^{\mu\nu} = e u^\mu u^\nu - (P_0 + \Pi) \Delta^{\mu\nu} + W^\mu u^\nu + W^\nu u^\mu + \pi^{\mu\nu}$$

$$n = N^\mu u_\mu \quad \text{charge density}$$

$$e = T^{\mu\nu} u_\mu u_\nu \quad \text{energy density}$$

$$P \equiv P_0 + \Pi \equiv -\frac{1}{3} \Delta_{\mu\nu} T^{\mu\nu} \quad \text{isotropic pressure}$$

$$V^\mu = \Delta^{\mu\alpha} N_\alpha \quad \text{charge flow}$$

$$W^\mu = \Delta^{\mu\alpha} u^\beta T_{\alpha\beta} \quad \text{energy-momentum flow}$$

$$\pi^{\mu\nu} = \left[ \frac{1}{2} \left( \Delta^{\mu\alpha} \Delta^{\nu\beta} + \Delta^{\mu\beta} \Delta^{\alpha\nu} \right) - \frac{1}{3} \Delta^{\mu\nu} \Delta^{\alpha\beta} \right] T_{\alpha\beta} \quad \text{stress tensor}$$

- We only have **5** equations and an EoS for **17** unknowns,  $n(1)$ ,  $e(1)$ ,  $u^\mu(3)$  and  $\Pi(1)$ ,  $V^\mu(3)$ ,  $W^\mu(3)$ ,  $\pi^{\mu\nu}(5)$ .

# General equations of motion I.

Using  $f_{\mathbf{k}} = f_{0\mathbf{k}} + \delta f_{\mathbf{k}}$  where  $\delta f_{\mathbf{k}} = f_{0\mathbf{k}}(1 - af_{0\mathbf{k}})\phi_{\mathbf{k}}$  we define the **irreducible moment**

$$\rho_r^{\mu_1 \dots \mu_\ell} \equiv \left\langle E_{\mathbf{k}u}^r k^{\langle \mu_1} \dots k^{\mu_\ell \rangle} \right\rangle_\delta \quad (16)$$

$$\langle \dots \rangle_\delta \equiv \langle \dots \rangle - \langle \dots \rangle_0 = \int dK (\dots) \delta f_{\mathbf{k}} \quad \text{and} \quad k^{\langle \mu_1} \dots k^{\mu_\ell \rangle} = \Delta_{\nu_1 \dots \nu_\ell}^{\mu_1 \dots \mu_\ell} k^{\nu_1} \dots k^{\nu_\ell}$$

The primary (14) dynamical moments in  $N^\mu$  and  $T^{\mu\nu}$

$$\begin{aligned} \rho_1 &\equiv \delta n = 0, & \rho_2 &\equiv \delta e = 0, & \rho_0 &\equiv -\frac{3}{m^2} \Pi, \\ \rho_0^\mu &\equiv V^\mu, & \rho_1^\mu &\equiv W^\mu = 0, & \rho_0^{\mu\nu} &\equiv \pi^{\mu\nu}. \end{aligned}$$

Now, writing the Boltzmann equation in the following form

$$D\delta f_{\mathbf{k}} = -Df_{0\mathbf{k}} - E_{\mathbf{k}u}^{-1} k_\nu \nabla^\nu (f_{0\mathbf{k}} + \delta f_{\mathbf{k}}) + E_{\mathbf{k}u}^{-1} C [f_{0\mathbf{k}} + \delta f_{\mathbf{k}}], \quad (17)$$

where  $D \equiv u^\mu \partial_\mu = \frac{d}{d\tau}$  and  $\nabla_\mu = \Delta_{\nu}^{\mu} \partial_\nu$ , the equations for  $D\rho_r^{\mu_1 \dots \mu_\ell}$  follow from,

$$D\rho_r^{\langle \mu_1 \dots \mu_\ell \rangle} \equiv \Delta_{\nu_1 \dots \nu_\ell}^{\mu_1 \dots \mu_\ell} \frac{d}{d\tau} \int dK E_{\mathbf{k}u}^r k^{\langle \nu_1} \dots k^{\nu_\ell \rangle} \delta f_{\mathbf{k}}. \quad (18)$$

# General equations of motion II.

Infinitely many coupled equations for  $\rho_r^{\mu_1 \dots \mu_\ell}$  equivalent to the Boltzmann equation !

## Scalar, vector, and tensor equations

$$\begin{aligned} \dot{\rho}_r - C_{r-1} &= \alpha_r^{(0)} \theta + \frac{\theta}{3} \left[ m_0^2 (r-1) \rho_{r-2} - (r+2) \rho_r - 3 \frac{G_{2r}}{D_{20}} \Pi \right] \\ &+ \frac{G_{3r}}{D_{20}} \partial_\mu V^\mu - \nabla_\mu \rho_{r-1}^\mu + r \rho_{r-1}^\mu \dot{u}_\mu + \left[ (r-1) \rho_{r-2}^{\mu\nu} + \frac{G_{2r}}{D_{20}} \pi^{\mu\nu} \right] \sigma_{\mu\nu}, \end{aligned}$$

$$\begin{aligned} \dot{\rho}_r^{\langle \mu \rangle} - C_{r-1}^{\langle \mu \rangle} &= \alpha_r^{(1)} \nabla^\mu \alpha + r \rho_{r-1}^{\mu\nu} \dot{u}_\nu - \frac{1}{3} \nabla^\mu \left( m_0^2 \rho_{r-1} - \rho_{r+1} \right) - \Delta_\alpha^\mu \left( \nabla_\nu \rho_{r-1}^{\alpha\nu} + \alpha_r^h \partial_\kappa \pi^{\kappa\alpha} \right) \\ &+ \frac{1}{3} \left[ m_0^2 (r-1) \rho_{r-2}^\mu - (r+3) \rho_r^\mu \right] \theta + \frac{1}{5} \sigma^{\mu\nu} \left[ 2m_0^2 (r-1) \rho_{r-2,\nu} - (2r+3) \rho_{r,\nu} \right] \\ &+ \frac{1}{3} \left[ m_0^2 r \rho_{r-1} - (r+3) \rho_{r+1} - 3\alpha_r^h \Pi \right] \dot{u}^\mu + \alpha_r^h \nabla^\mu \Pi + \rho_{r,\nu} \omega^{\mu\nu} + (r-1) \rho_{r-2}^{\mu\nu\lambda} \sigma_{\nu\lambda}, \end{aligned}$$

$$\begin{aligned} \dot{\rho}_r^{\langle \mu\nu \rangle} - C_{r-1}^{\langle \mu\nu \rangle} &= 2\alpha_r^{(2)} \sigma^{\mu\nu} + \frac{2}{15} \left[ m_0^4 (r-1) \rho_{r-2} - m_0^2 (2r+3) \rho_r + (r+4) \rho_{r+2} \right] \sigma^{\mu\nu} + 2\rho_r^{\lambda\langle \mu} \omega_{\lambda}^{\nu \rangle} \\ &+ \frac{2}{5} \dot{u}^{\langle \mu} \left[ m_0^2 r \rho_{r-1}^{\nu \rangle} - (r+5) \rho_{r+1}^{\nu \rangle} \right] - \frac{2}{5} \nabla^{\langle \mu} \left( m_0^2 \rho_{r-1}^{\nu \rangle} - \rho_{r+1}^{\nu \rangle} \right) \\ &+ \frac{1}{3} \left[ m_0^2 (r-1) \rho_{r-2}^{\mu\nu} - (r+4) \rho_r^{\mu\nu} \right] \theta + \frac{2}{7} \left[ 2m_0^2 (r-1) \rho_{r-2}^{\lambda\langle \mu} - (2r+5) \rho_r^{\lambda\langle \mu} \right] \sigma_{\lambda}^{\nu \rangle} \\ &+ r \rho_{r-1}^{\mu\nu\gamma} \dot{u}_\gamma - \Delta_{\alpha\beta}^{\mu\nu} \nabla_\lambda \rho_{r-1}^{\alpha\beta\lambda} + (r-1) \rho_{r-2}^{\mu\nu\lambda\kappa} \sigma_{\lambda\kappa}. \end{aligned} \quad (19)$$

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Phys. Rev. D **85**, 114047 (2012)

# General equations of motion III.

The Dirk Not Mother (DNMR) recipe:

## Expansion around equilibrium

$$f_{\mathbf{k}} = f_{0\mathbf{k}} + f_{0\mathbf{k}} (1 - af_{0\mathbf{k}}) \sum_{\ell=0}^{\infty} \sum_{n=0}^{N_{\ell}} \rho_n^{\mu_1 \dots \mu_{\ell}} k_{\langle \mu_1} \dots k_{\mu_{\ell} \rangle} \mathcal{H}_{\mathbf{k}n}^{(\ell)}, \quad (20)$$

$$\mathcal{H}_{\mathbf{k}n}^{(\ell)} = \frac{(-1)^{\ell}}{\ell! J_{2\ell, \ell}} \sum_{i=n}^{N_{\ell}} \sum_{m=0}^i a_{in}^{(\ell)} a_{im}^{(\ell)} E_{\mathbf{k}}^m. \quad (21)$$

The coefficients  $a_{ij}^{(\ell)}$  are calculated via the Gram-Schmidt orthogonalization procedure

## Define the negative-order moments $r < 0$

$$\rho_{\pm r}^{\mu_1 \dots \mu_{\ell}} = \sum_{n=0}^{N_{\ell}} \rho_n^{\mu_1 \dots \mu_{\ell}} \mathcal{F}_{\mp r, n}^{(\ell)}, \quad (22)$$

$$\mathcal{F}_{\mp r, n}^{(\ell)} \equiv \frac{\ell!}{(2\ell + 1)!!} \int d\mathbf{k} E_{\mathbf{k}}^{\pm r} \left( \Delta^{\alpha\beta} k_{\alpha} k_{\beta} \right)^{\ell} \mathcal{H}_{\mathbf{k}n}^{(\ell)} f_{0\mathbf{k}} (1 - af_{0\mathbf{k}}). \quad (23)$$

The Grad or Israel-Stewart approximations have no negative-moments!

## General equations of motion IV.

## The moments of the linearized collision integral

$$C_{r-1}^{\langle \mu_1 \dots \mu_\ell \rangle} \equiv - \sum_{n=0}^{N_\ell} \mathcal{A}_{rn}^{(\ell)} \rho_n^{\mu_1 \dots \mu_\ell}, \quad \sum_{r=0}^{N_\ell} \tau_{nr}^{(\ell)} C_{r-1}^{\langle \mu_1 \dots \mu_\ell \rangle} = -\rho_n^{\langle \mu_1 \dots \mu_\ell \rangle}, \quad (24)$$

where  $\mathcal{A}_{in}^{(\ell)}$  is the collision matrix while its inverse

$$\tau_{rn}^{(\ell)} \equiv \left( \mathcal{A}^{(\ell)} \right)_{rn}^{-1} = \sum_{m=0}^{N_\ell} \Omega_{rm}^{(\ell)} \frac{1}{\chi_m^{(\ell)}} \left( \Omega^{(\ell)} \right)_{mn}^{-1}, \quad (25)$$

is the microscopic time scales proportional to  $\tau_{\text{mfp}} = \lambda_{\text{mfp}}/c$  between collisions.

Here  $\Omega^{(\ell)}$  diagonalizes  $\mathcal{A}^{(\ell)}$  according to;  $(\Omega^{(\ell)})^{-1} \mathcal{A}^{(\ell)} \Omega^{(\ell)} = \text{diag} \left( \chi_0^{(\ell)}, \chi_1^{(\ell)}, \dots \right)$

while we set  $\Omega_{00}^{(\ell)} = 1$ , while arrange the eigenvalues in increasing order,  $\chi_r^{(\ell)} \leq \chi_{r+1}^{(\ell)}$ .

## The RTA collision integral

$$C_{r-1}^{\langle \mu_1 \dots \mu_\ell \rangle} \equiv - \sum_{n=0}^{N_\ell} \frac{\delta_{rn}^{(\ell)}}{\tau_R} \rho_n^{\mu_1 \dots \mu_\ell} = - \frac{1}{\tau_R} \rho_r^{\mu_1 \dots \mu_\ell}, \quad \tau_{rn}^{(\ell)} = \tau_R \delta_{rn}, \quad \Omega_{rn}^{(\ell)} = \delta_{rn}. \quad (26)$$

## General equations of motion V.

## The moment equations

$$\dot{\rho}_r + \sum_{n=0}^{N_0} \mathcal{A}_{rn}^{(0)} \rho_n = \alpha_r^{(0)} \theta + (\text{higher-order terms}), \quad (27)$$

$$\dot{\rho}_r^{(\mu)} + \sum_{n=0}^{N_1} \mathcal{A}_{rn}^{(1)} \rho_n^\mu = \alpha_r^{(1)} \nabla^\mu \alpha + (\text{higher-order terms}), \quad (28)$$

$$\dot{\rho}_r^{(\mu\nu)} + \sum_{n=0}^{N_2} \mathcal{A}_{rn}^{(2)} \rho_n^{\mu\nu} = 2\alpha_r^{(2)} \sigma^{\mu\nu} + (\text{higher-order terms}), \quad (29)$$

## The moment equations - final form

$$\sum_{r=0}^{N_0} \tau_{nr}^{(0)} \dot{\rho}_r + \rho_n = \theta \sum_{r=0}^{N_0} \tau_{nr}^{(0)} \alpha_r^{(0)} + \sum_{r=0}^{N_0} \tau_{nr}^{(0)} (\text{higher-order terms}), \quad (30)$$

$$\sum_{r=0}^{N_1} \tau_{nr}^{(1)} \dot{\rho}_r^{(\mu)} + \rho_n^\mu = \nabla^\mu \alpha \sum_{r=0}^{N_1} \tau_{nr}^{(1)} \alpha_r^{(1)} + \sum_{r=0}^{N_1} \tau_{nr}^{(1)} (\text{higher-order terms}), \quad (31)$$

$$\sum_{r=0}^{N_2} \tau_{nr}^{(2)} \dot{\rho}_r^{(\mu\nu)} + \rho_n^{\mu\nu} = 2\sigma^{\mu\nu} \sum_{r=0}^{N_2} \tau_{nr}^{(2)} \alpha_r^{(2)} + \sum_{r=0}^{N_2} \tau_{nr}^{(2)} (\text{higher-order terms}), \quad (32)$$

## General equations of motion VI.

The 14 dynamical moments are  $\rho_0 = -\frac{3}{m_0^2}\Pi$ ,  $\rho_0^\mu = V^\mu$ , and  $\rho_0^{\mu\nu} = \pi^{\mu\nu}$ ,

The DNMR approximation for the non-dynamical moments

$$\rho_{r>0} \simeq -\frac{3}{m_0^2}\Omega_{r0}^{(0)}\Pi + \frac{3}{m_0^2}(\zeta_r - \Omega_{r0}^{(0)}\zeta_0)\theta, \quad \rho_{-r} \simeq -\frac{3}{m_0^2}\gamma_{r0}^{(0)}\Pi + O(Kn), \quad (33)$$

$$\rho_{r>0}^\mu \simeq \Omega_{r0}^{(1)}V^\mu + (\kappa_r - \Omega_{r0}^{(1)}\kappa_0)\nabla^\mu\alpha, \quad \rho_{-r}^\mu \simeq \gamma_{r0}^{(1)}V^\mu + O(Kn), \quad (34)$$

$$\rho_{r>0}^{\mu\nu} \simeq \Omega_{r0}^{(2)}\pi^{\mu\nu} + 2(\eta_r - \Omega_{r0}^{(2)}\eta_0)\sigma^{\mu\nu}, \quad \rho_{-r}^{\mu\nu} \simeq \gamma_{r0}^{(2)}\pi^{\mu\nu} + O(Kn). \quad (35)$$

where the **first-order transport coefficients** and the  $\gamma_{r0}^{(\ell)}$  coefficients are:

$$\zeta_r \equiv \frac{m_0^2}{3} \sum_{n=0, \neq 1, 2}^{N_0} \tau_m^{(0)} \alpha_n^{(0)}, \quad \kappa_r \equiv \sum_{n=0, \neq 1}^{N_1} \tau_m^{(1)} \alpha_n^{(1)}, \quad \eta_r \equiv \sum_{n=0}^{N_2} \tau_m^{(2)} \alpha_n^{(2)}, \quad (36)$$

$$\gamma_{r0}^{(0)} \equiv \sum_{n=0, \neq 1, 2}^{N_0} \mathcal{F}_m^{(0)} \Omega_{n0}^{(0)}, \quad \gamma_{r0}^{(1)} \equiv \sum_{n=0, \neq 1}^{N_1} \mathcal{F}_m^{(1)} \Omega_{n0}^{(1)}, \quad \gamma_{r0}^{(2)} \equiv \sum_{n=0}^{N_2} \mathcal{F}_m^{(2)} \Omega_{n0}^{(2)}. \quad (37)$$

For  $r = 0$  one obtains the usual Navier-Stokes coefficients:  $\zeta_0 \equiv \tau_{00}^{(0)} \alpha_0^{(0)} = \tau_\Pi \alpha_0^{(0)}$ ,  $\kappa_0 = \tau_V \alpha_0^{(1)}$ , and  $\eta_0 = \tau_\pi \alpha_0^{(2)}$ .

## General equations of motion VII.

The corrected DNMR approximation for the non-dynamical moments

$$\rho_{-r} \simeq -\frac{3}{m_0^2} \gamma_{r0}^{(0)} \Pi + \frac{3}{m_0^2} (\Gamma_{r0}^{(0)} - \gamma_{r0}^{(0)}) \xi_0 \theta \quad \Rightarrow \quad \rho_{-r} \simeq \frac{3}{m_0^2} \Gamma_{r0}^{(0)}, \quad (38)$$

$$\rho_{-r}^\mu \simeq \gamma_{r0}^{(1)} V^\mu + (\Gamma_{r0}^{(1)} - \gamma_{r0}^{(1)}) \kappa_0 \nabla^\mu \alpha \quad \Rightarrow \quad \rho_{-r}^\mu \simeq \Gamma_{r0}^{(1)} V^\mu, \quad (39)$$

$$\rho_{-r}^{\mu\nu} \simeq \gamma_{r0}^{(2)} \pi^{\mu\nu} + 2(\Gamma_{r0}^{(2)} - \gamma_{r0}^{(2)}) \eta_0 \sigma^{\mu\nu} \quad \Rightarrow \quad \rho_{-r}^{\mu\nu} \simeq \Gamma_{r0}^{(2)} \pi^{\mu\nu}. \quad (40)$$

where the  $O(Kn)$  corrections are now explicitly included, *cDNMR*

$$\gamma_{r0}^{(0)} \equiv \sum_{n=0, \neq 1, 2}^{N_0} \mathcal{F}_m^{(0)} \Omega_{n0}^{(0)}, \quad \gamma_{r0}^{(1)} \equiv \sum_{n=0, \neq 1}^{N_1} \mathcal{F}_m^{(1)} \Omega_{n0}^{(1)}, \quad \gamma_{r0}^{(2)} \equiv \sum_{n=0}^{N_2} \mathcal{F}_m^{(2)} \Omega_{n0}^{(2)}, \quad (41)$$

$$\Gamma_{r0}^{(0)} \equiv \sum_{n=0, \neq 1, 2}^{N_0} \mathcal{F}_m^{(0)} \frac{\zeta_n}{\zeta_0}, \quad \Gamma_{r0}^{(1)} \equiv \sum_{n=0, \neq 1}^{N_1} \mathcal{F}_m^{(1)} \frac{\kappa_n}{\kappa_0}, \quad \Gamma_{r0}^{(2)} \equiv \sum_{n=0}^{N_2} \mathcal{F}_m^{(2)} \frac{\eta_n}{\eta_0}. \quad (42)$$

V. E. Ambrus, E. Molnár and D. H. Rischke, "Transport coefficients of second-order relativistic fluid dynamics in the relaxation-time approximation," Phys. Rev. D **106**, no.7, 076005 (2022)

## General equations of motion VIII.

This is the way in RTA!

$$\tau_R \dot{\rho}_r + \rho_r = \tau_R \alpha_r^{(0)} \theta + O(2) = \frac{3}{m_0^2} \xi_r \theta + O(\text{Re}^{-1} \text{Kn}) \quad (43)$$

$$\tau_R \dot{\rho}_r^{\langle \mu \rangle} + \rho_r^{\langle \mu \rangle} = \tau_R \alpha_r^{(1)} \nabla^\mu \alpha + O(2) = \kappa_r \nabla^\mu \alpha + O(\text{Re}^{-1} \text{Kn}), \quad (44)$$

$$\tau_R \dot{\rho}_r^{\langle \mu \nu \rangle} + \rho_r^{\langle \mu \nu \rangle} = 2 \tau_R \alpha_r^{(2)} \sigma^{\mu \nu} + O(2) = 2 \eta_r \sigma^{\mu \nu} + O(\text{Re}^{-1} \text{Kn}), \quad (45)$$

where  $\zeta_r = \tau_R \frac{m_0^2}{3} \alpha_r^{(0)}$ ,  $\kappa_r = \tau_R \alpha_r^{(1)}$  and  $\eta_r = \tau_R \alpha_r^{(2)}$ , hence the ratio of coefficients,

$$\mathcal{R}_{r0}^{(\ell)} \equiv \frac{\alpha_r^{(\ell)}}{\alpha_0^{(\ell)}} \implies \mathcal{R}_{r0}^{(0)} = \frac{\zeta_r}{\zeta_0}, \quad \mathcal{R}_{r0}^{(1)} = \frac{\kappa_r}{\kappa_0}, \quad \mathcal{R}_{r0}^{(2)} = \frac{\eta_r}{\eta_0}. \quad (46)$$

The order of things in RTA!

$$\rho_{r \neq 0} \simeq -\frac{3}{m_0^2} \mathcal{R}_{r0}^{(0)} \Pi, \quad \rho_{r \neq 0}^\mu \simeq \mathcal{R}_{r0}^{(1)} V^\mu, \quad \rho_{r \neq 0}^{\mu \nu} \simeq \mathcal{R}_{r0}^{(2)} \pi^{\mu \nu}, \quad (47)$$

V. E. Ambrus, E. Molnár and D. H. Rischke, "Transport coefficients of second-order relativistic fluid dynamics in the relaxation-time approximation," Phys. Rev. D **106**, no.7, 076005 (2022)

# Second-order fluid dynamics I.

## Equation for the bulk pressure

$$\begin{aligned} \tau_{\Pi} \dot{\Pi} + \Pi = & -\zeta\theta - \ell_{\Pi V} \nabla_{\mu} V^{\mu} - \tau_{\Pi V} V_{\mu} \dot{u}^{\mu} - \delta_{\Pi\Pi} \Pi\theta \\ & - \lambda_{\Pi V} V_{\mu} \nabla^{\mu} \alpha + \lambda_{\Pi\pi} \pi^{\mu\nu} \sigma_{\mu\nu} , \end{aligned} \quad (48)$$

## Equation for the diffusion current

$$\begin{aligned} \tau_V \dot{V}^{\langle\mu\rangle} + V^{\mu} = & \kappa \nabla^{\mu} \alpha + -\tau_V V_{\nu} \omega^{\nu\mu} - \delta_{VV} V^{\mu} \theta - \ell_{V\Pi} \nabla^{\mu} \Pi \\ & + \ell_{V\pi} \Delta^{\mu\nu} \nabla_{\lambda} \pi^{\lambda}_{\nu} + \tau_{V\Pi} \Pi \dot{u}^{\mu} - \tau_{V\pi} \pi^{\mu\nu} \dot{u}_{\nu} \\ & - \lambda_{VV} V_{\nu} \sigma^{\mu\nu} + \lambda_{V\Pi} \Pi \nabla^{\mu} \alpha - \lambda_{V\pi} \pi^{\mu\nu} \nabla_{\nu} \alpha , \end{aligned} \quad (49)$$

## Equation for the shear-stress tensor

$$\begin{aligned} \tau_{\pi} \dot{\pi}^{\langle\mu\nu\rangle} + \pi^{\mu\nu} = & 2\eta\sigma^{\mu\nu} + 2\tau_{\pi} \pi_{\lambda}^{\langle\mu} \omega^{\nu\rangle\lambda} - \delta_{\pi\pi} \pi^{\mu\nu} \theta - \tau_{\pi\pi} \pi^{\lambda\langle\mu} \sigma_{\lambda}^{\nu\rangle} + \lambda_{\pi\Pi} \Pi \sigma^{\mu\nu} \\ & - \tau_{\pi V} V^{\langle\mu} \dot{u}^{\nu\rangle} + \ell_{\pi V} \nabla^{\langle\mu} V^{\nu\rangle} + \lambda_{\pi V} V^{\langle\mu} \nabla^{\nu\rangle} \alpha . \end{aligned} \quad (50)$$

## First-order transport coefficients

$$\zeta = \tau_{\Pi} \frac{m_0^2}{3} \alpha_0^{(0)} , \quad \kappa = \tau_V \alpha_0^{(1)} , \quad \eta = \tau_{\pi} \alpha_0^{(2)} . \quad (51)$$

## Second-order fluid dynamics II.

The second-order coefficients in the 14-moment approximation are:

$$\delta_{\Pi\Pi} = \tau_{\Pi} \left( \frac{2}{3} - \frac{m_0^2}{3} \frac{G_{20}}{D_{20}} + \frac{m_0^2}{3} \mathcal{R}_{-2,0}^{(0)} \right), \quad \lambda_{\Pi V} = -\tau_{\Pi} \frac{m_0^2}{3} \left( \frac{\partial \mathcal{R}_{-1,0}^{(1)}}{\partial \alpha} + \frac{1}{h} \frac{\partial \mathcal{R}_{-1,0}^{(1)}}{\partial \beta} \right),$$

$$\ell_{\Pi V} = \tau_{\Pi} \frac{m_0^2}{3} \left( \frac{G_{30}}{D_{20}} - \mathcal{R}_{-1,0}^{(1)} \right), \quad \tau_{\Pi V} = -\tau_{\Pi} \frac{m_0^2}{3} \left( \frac{G_{30}}{D_{20}} - \frac{\partial \mathcal{R}_{-1,0}^{(1)}}{\partial \ln \beta} \right), \quad \lambda_{\Pi\pi} = -\tau_{\Pi} \frac{m_0^2}{3} \left( \frac{G_{20}}{D_{20}} - \mathcal{R}_{-2,0}^{(2)} \right).$$

$$\delta_{VV} = \tau_V \left( 1 + \frac{m_0^2}{3} \mathcal{R}_{-2,0}^{(1)} \right), \quad \ell_{V\Pi} = \frac{\tau_V}{h} \left( 1 - h \mathcal{R}_{-1,0}^{(0)} \right), \quad \ell_{V\pi} = \frac{\tau_V}{h} \left( 1 - h \mathcal{R}_{-1,0}^{(2)} \right),$$

$$\tau_{V\Pi} = \frac{\tau_V}{h} \left( 1 - h \frac{\partial \mathcal{R}_{-1,0}^{(0)}}{\partial \ln \beta} \right), \quad \tau_{V\pi} = \frac{\tau_V}{h} \left( 1 - h \frac{\partial \mathcal{R}_{-1,0}^{(2)}}{\partial \ln \beta} \right), \quad \lambda_{VV} = \tau_V \left( \frac{3}{5} + \frac{2m_0^2}{5} \mathcal{R}_{-2,0}^{(1)} \right),$$

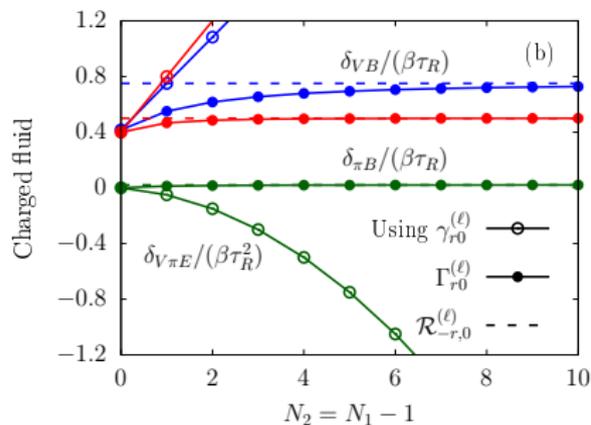
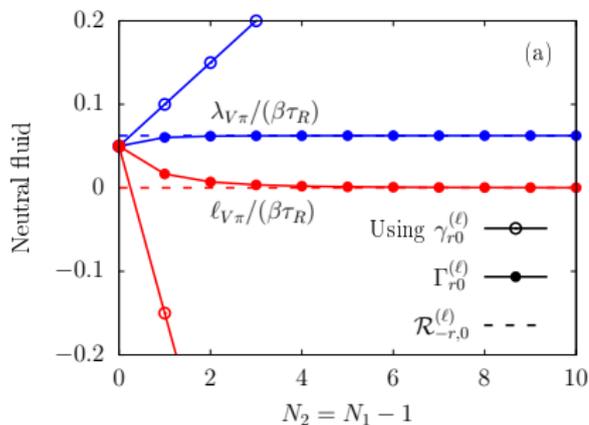
$$\lambda_{V\Pi} = \tau_V \left( \frac{\partial \mathcal{R}_{-1,0}^{(0)}}{\partial \alpha} + \frac{1}{h} \frac{\partial \mathcal{R}_{-1,0}^{(0)}}{\partial \beta} \right), \quad \lambda_{V\pi} = \tau_V \left( \frac{\partial \mathcal{R}_{-1,0}^{(2)}}{\partial \alpha} + \frac{1}{h} \frac{\partial \mathcal{R}_{-1,0}^{(2)}}{\partial \beta} \right).$$

$$\delta_{\pi\pi} = \tau_{\pi} \left( \frac{4}{3} + \frac{m_0^2}{3} \mathcal{R}_{-2,0}^{(2)} \right), \quad \tau_{\pi\pi} = \tau_{\pi} \left( \frac{10}{7} + \frac{4m_0^2}{7} \mathcal{R}_{-2,0}^{(2)} \right), \quad \lambda_{\pi\Pi} = \tau_{\pi} \left( \frac{6}{5} + \frac{2m_0^2}{5} \mathcal{R}_{-2,0}^{(0)} \right),$$

$$\tau_{\pi V} = -\tau_{\pi} \frac{2m_0^2}{5} \frac{\partial \mathcal{R}_{-1,0}^{(1)}}{\partial \ln \beta}, \quad \ell_{\pi V} = -\tau_{\pi} \frac{2m_0^2}{5} \mathcal{R}_{-1,0}^{(1)}, \quad \lambda_{\pi V} = -\tau_{\pi} \frac{2m_0^2}{5} \left( \frac{\partial \mathcal{R}_{-1,0}^{(1)}}{\partial \alpha} + \frac{1}{h} \frac{\partial \mathcal{R}_{-1,0}^{(1)}}{\partial \beta} \right).$$

## Second-order fluid dynamics III.

Coefficients in the massless limit

All other coefficients agree in all approximations; i.e.,  $\gamma_{r0}$  vs.  $\Gamma_{r0}$  vs.  $\mathcal{R}_{-r,0}$ .

# Second-order fluid dynamics IV.

## With a binary collision integral

For a gas of ultrarelativistic hard spheres with constant cross-section, the 14-dynamical moment approximation for truncation orders  $N_0 - 2 = N_1 - 1 = N_2 = 1000$ , corresponding to  $N_0 + 3N_1 + 5N_2 + 9 = 9014$  moments included in the basis.

## Coefficients of the diffusion equation

Method	$\kappa$	$\tau_V[\lambda_{\text{mfp}}]$	$\delta_{VV}[\tau_V]$	$\ell_{V\pi}[\tau_V] = \tau_{V\pi}[\tau_V]$	$\lambda_{VV}[\tau_V]$	$\lambda_{V\pi}[\tau_V]$
14M	$1/12\sigma$	$9/4$	1	$\beta/20$	$3/5$	$\beta/20$
IReD	$0.15892/\sigma$	2.0838	1	$0.028371\beta$	0.89862	$0.069273\beta$
DNMR	$0.15892/\sigma$	2.5721	1	$0.11921\beta$	0.92095	$0.051709\beta$
cDNMR	$0.15892/\sigma$	2.5721	1	$0.098534\beta$	0.92095	$0.056878\beta$

## Coefficients of the shear-stress equation

Method	$\eta$	$\tau_\pi[\lambda_{\text{mfp}}]$	$\delta_{\pi\pi}[\tau_\pi]$	$\ell_{\pi V}[\tau_\pi]$	$\tau_{\pi V}[\tau_\pi]$	$\tau_{\pi\pi}[\tau_\pi]$	$\lambda_{\pi V}[\tau_\pi]$
14M	$4/(3\sigma\beta)$	$5/3$	$4/3$	0	0	$10/7$	0
IReD	$1.2676/(\sigma\beta)$	1.6557	$4/3$	$-0.56960/\beta$	$-2.2784/\beta$	1.6945	$0.20503/\beta$
c&DNMR	$1.2676/(\sigma\beta)$	2	$4/3$	$-0.68317/\beta$	$-2.7327/\beta$	1.6888	$0.24188/\beta$

D. Wagner, V. E. Ambrus and E. Molnar, "Analytical structure of the binary collision integral and the ultrarelativistic limit of transport coefficients of an ideal gas," Phys. Rev. D **109**, no.5, 056018 (2024)

# Applications I.

## Relativistic ideal gas of massive particles

## Relativistic ideal gas I.

## Thermodynamic integrals

$$\begin{aligned}
 I_{rq}(\alpha, \beta) &\equiv \frac{(-1)^r}{(2q+1)!!} \int dK E_{\mathbf{k}}^{r-2q} (\Delta^{\mu\nu} k_{\mu} k_{\nu})^q f_{0\mathbf{k}} \\
 &= \frac{g e^{\alpha}}{2\pi^2} \frac{m_0^{r+2}}{(2q+1)!!} \int_1^{\infty} dx x^{r-2q} (x^2 - 1)^{q+\frac{1}{2}} e^{-zx}, \quad (52)
 \end{aligned}$$

where  $z = m_0/T$  and  $f_{0\mathbf{k}} = e^{\alpha - \beta E_{\mathbf{k}}}$ , while

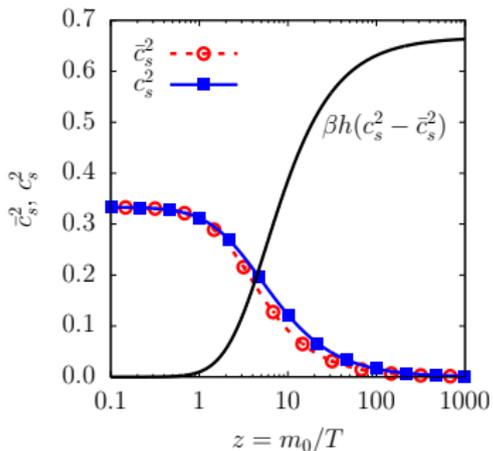
## Bessel functions of second-kind

$$K_q(z) \equiv \frac{z^q}{(2q-1)!!} \int_1^{\infty} dx (x^2 - 1)^{q-\frac{1}{2}} e^{-zx}. \quad (53)$$

## Primary thermodynamic quantities

$$n \equiv I_{10} = \frac{g e^{\alpha}}{2\pi^2} T^3 z^2 K_2(z), \quad P \equiv I_{21} = nT, \quad e \equiv I_{20} = P \left[ 3 + z \frac{K_1(z)}{K_2(z)} \right]. \quad (54)$$

## Relativistic ideal gas II.



The speed of sound - with  $N^\mu$  conservation

$$c_s^2 \equiv \left( \frac{\partial P}{\partial e} \right)_n + \frac{1}{h} \left( \frac{\partial P}{\partial n} \right)_e = \frac{c_p P}{c_v (e + P)}, \quad (55)$$

The speed of sound - withOUT  $N^\mu$  conservation

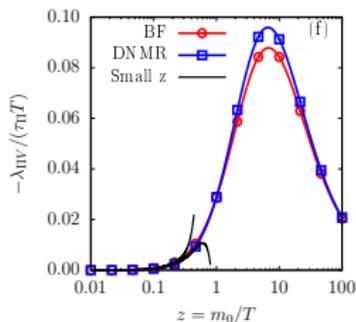
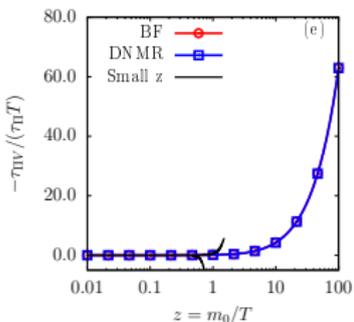
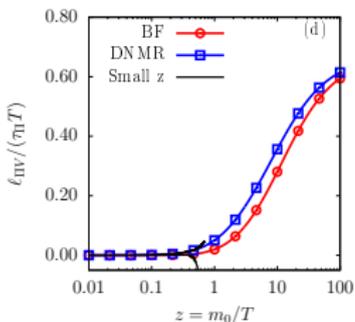
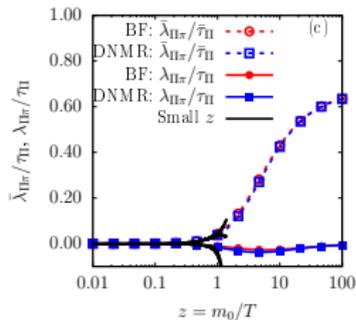
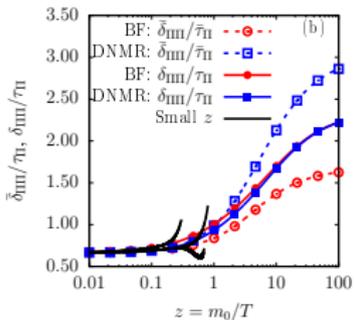
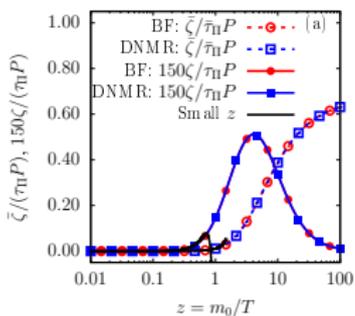
$$\bar{c}_s^2 \equiv \left( \frac{\partial P}{\partial e} \right)_\mu = \frac{l_{31}}{l_{30}} = \frac{P(e + P)}{c_v P^2 + e^2}, \quad (56)$$

where  $c_v$  and  $c_p = c_v + 1$  are heat capacities at constant volume or pressure.

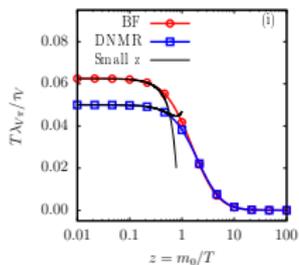
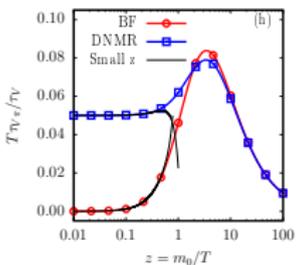
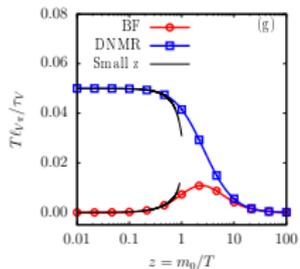
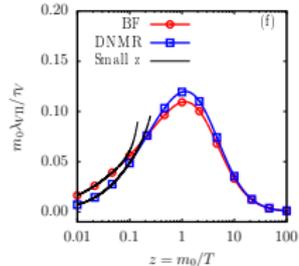
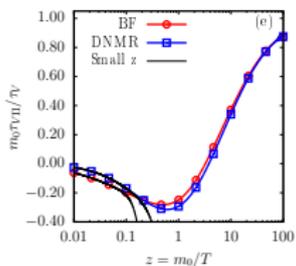
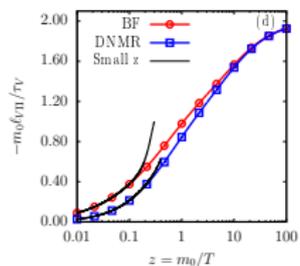
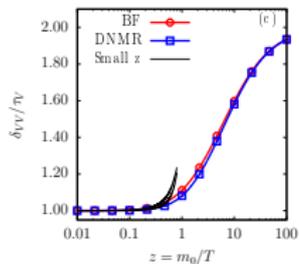
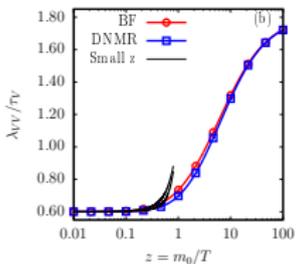
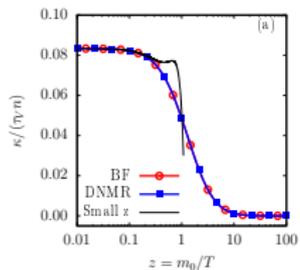
The coefficient(s) of bulk viscosity

$$\frac{\zeta}{\tau_\Pi} = \frac{5}{3} \frac{\eta}{\tau_\pi} - c_s^2 (e + P), \quad \frac{\bar{\zeta}}{\tau_\Pi} = \frac{5}{3} \frac{\eta}{\tau_\pi} - \bar{c}_s^2 (e + P).$$

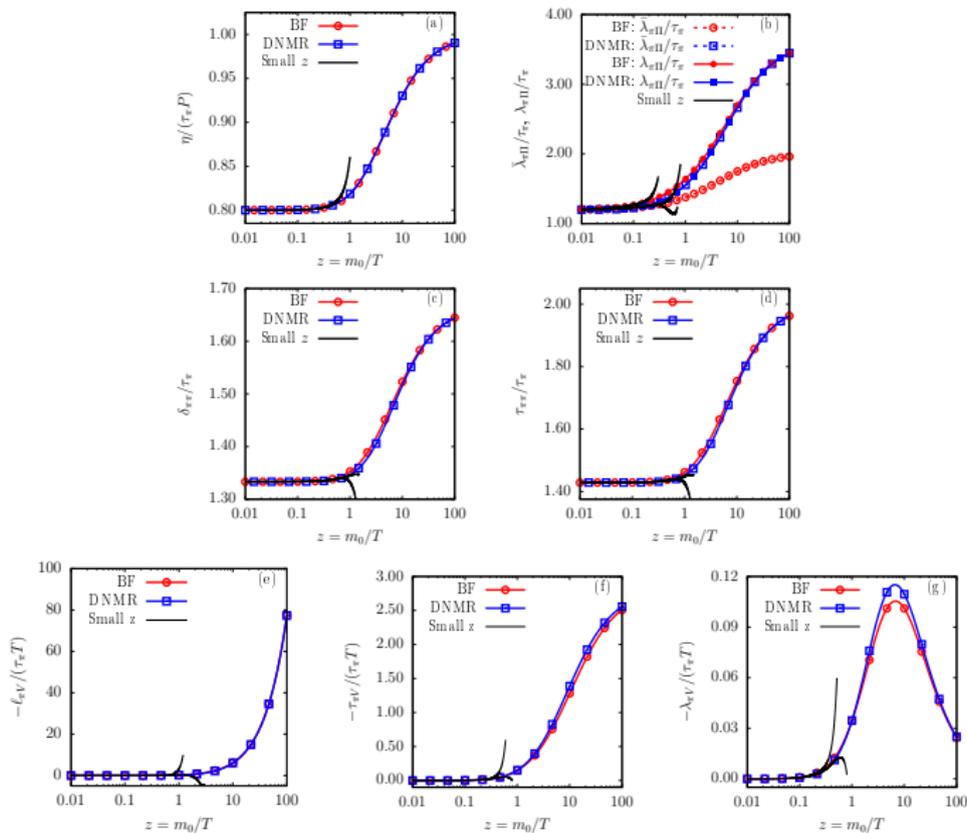
## Coefficients of the bulk equation



## Coefficients of the diffusion equation



## Coefficients of the shear equation



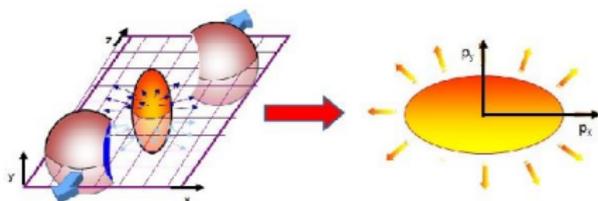
# Applications II.

## Bjorken expansion

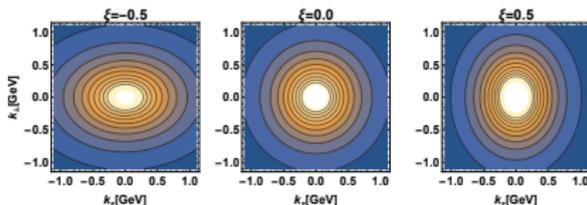
$$u^\mu \equiv \left( \frac{t}{\tau}, 0, 0, \frac{z}{\tau} \right) = (\cosh \eta_s, 0, 0, \sinh \eta_s)$$

$$l^\mu \equiv \left( \frac{z}{\tau}, 0, 0, \frac{t}{\tau} \right) = (\sinh \eta_s, 0, 0, \cosh \eta_s)$$

## Applications II.



- Elliptic flow - momentum space anisotropy of particle emission in non-central heavy-ion collisions.



$$\hat{f}_{RS} \equiv \left[ \exp \left( -\beta \sqrt{k_{\perp}^2 + (1 + \xi)k_z^2} \right) \right]$$

Spheroidal momentum distribution function.  
Prolate (left) - ISOtropic - Oblate (right)

## Anisotropic fluid dynamics

$$\hat{N}^{\mu} = \hat{n}u^{\mu} + \hat{n}_l l^{\mu} \quad (57)$$

$$\hat{T}^{\mu\nu} = \hat{e}u^{\mu}u^{\nu} + 2\hat{M}u^{(\mu}l^{\nu)} + \hat{P}_l l^{\mu}l^{\nu} - \hat{P}_{\perp}\Xi^{\mu\nu} \quad (58)$$

where  $\Xi^{\mu\nu} \equiv g^{\mu\nu} - u^{\mu}u^{\nu} + l^{\mu}l^{\nu} = \Delta^{\mu\nu} + l^{\mu}l^{\nu}$  such that  $\Xi^{\mu\nu}u_n u = \Xi^{\mu\nu}l_{\nu} = 0$ .

## Boltzmann equation for the boost invariant expansion

$$\frac{\partial f_{\mathbf{k}}}{\partial \tau} - \frac{v^z}{\tau} (1 - v_z^2) \frac{\partial f_{\mathbf{k}}}{\partial v_z} = -\frac{1}{\tau_R} (f_{\mathbf{k}} - f_{0\mathbf{k}}), \quad v^z \equiv \tanh(y - \eta_s)$$

## Second-order fluid dynamics - 4 diff equations

$$Dn + \frac{n}{\tau} = 0, \quad De + \frac{1}{\tau} (e + P_l) = 0, \quad P_l \equiv P + \Pi - \pi, \quad P_{\perp} \equiv P + \Pi + \frac{\pi}{2},$$

$$\tau_R D\Pi + \Pi = -\frac{\zeta}{\tau} - \delta_{\Pi\Pi} \frac{\Pi}{\tau} + \lambda_{\Pi\pi} \frac{\pi}{\tau}, \quad \tau_R D\pi + \pi = \frac{4\eta}{3\tau} - \delta_{\pi\pi} \frac{\pi}{\tau} - \tau_{\pi\pi} \frac{\pi}{3\tau} + \lambda_{\pi\Pi} \frac{2\Pi}{3\tau}.$$

## Anisotropic fluid dynamics - 3 diff equations

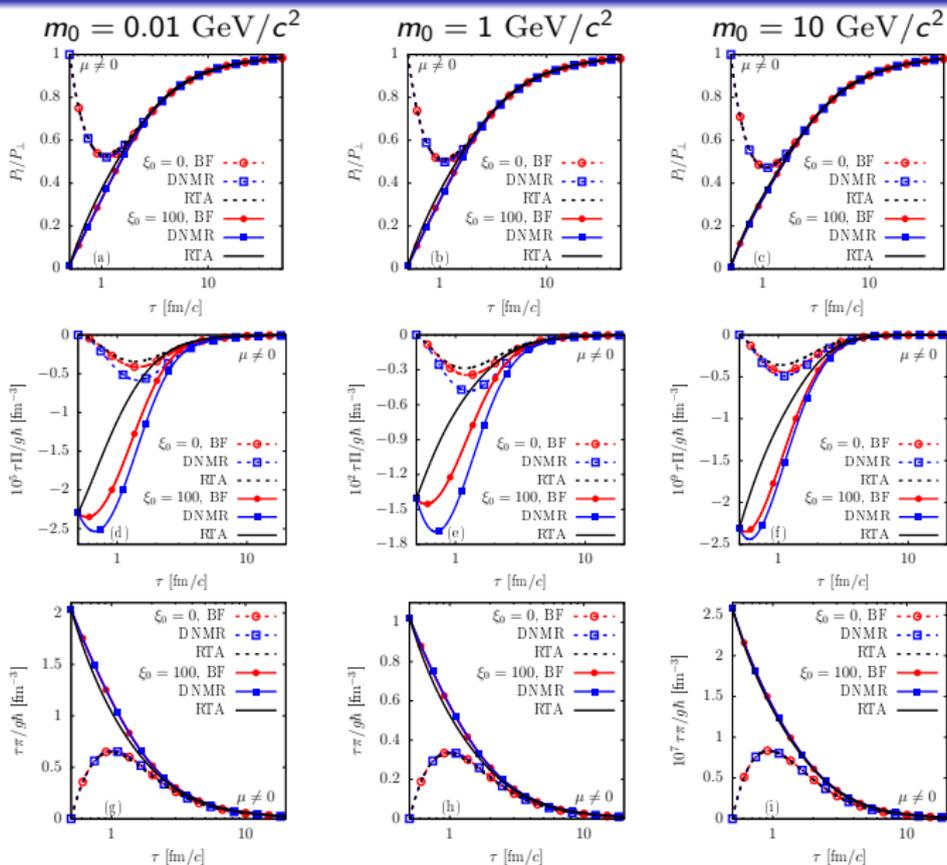
$$\hat{f}_{RS} = \exp\left(\hat{\alpha} - \frac{k^T}{\Lambda} \sqrt{1 + \xi v_z^2}\right), \quad \hat{I}_{000}^{RS} = \int dK \hat{f}_{RS}, \quad \hat{I}_{240}^{RS} = \int dK E_{\mathbf{k}u}^{-2} E_{\mathbf{k}l}^4 \hat{f}_{RS},$$

$$D\hat{n} + \frac{\hat{n}}{\tau} = -\frac{1}{\tau_R} (\hat{n} - n), \quad D\hat{e} + \frac{1}{\tau} (\hat{e} + \hat{P}_l) = -\frac{1}{\tau_R} (\hat{e} - e), \quad \hat{P}_{\perp} \equiv \hat{I}_{201}^{RS} = \frac{1}{2} (e - \hat{P}_l - m_0^2 \hat{I}_{000}^{RS})$$

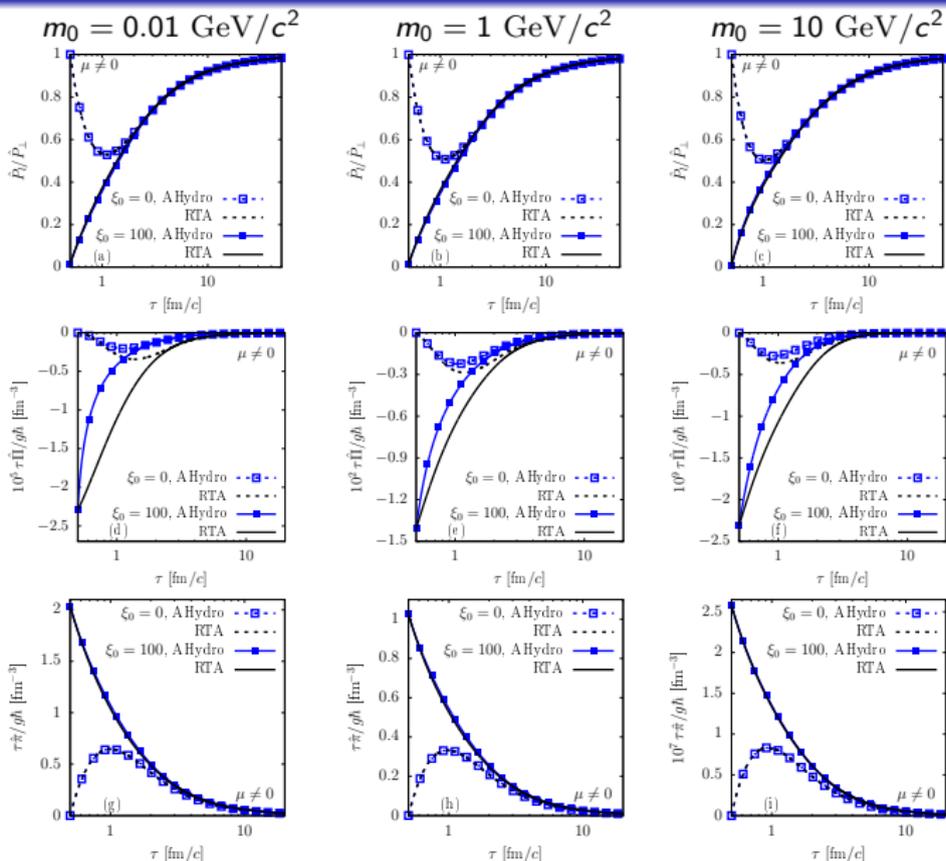
$$D\hat{P}_l + \frac{1}{\tau} (3\hat{P}_l - \hat{I}_{240}^{RS}) = -\frac{1}{\tau_R} (\hat{P}_l - P), \quad \hat{\Pi} = \frac{1}{3} (\hat{P}_l + 2\hat{P}_{\perp}) - P, \quad \hat{\pi} = \frac{2}{3} (\hat{P}_{\perp} - \hat{P}_l).$$

There are no free parameters, everything above is a function of  $z = m_0/T$ .

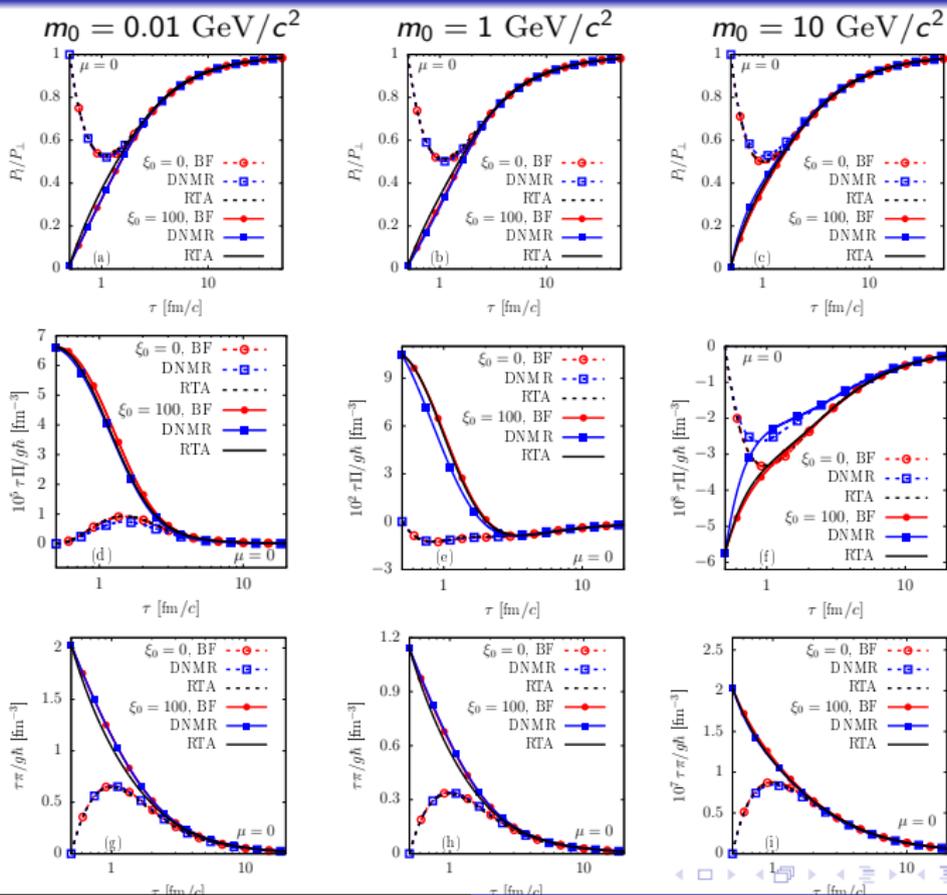
## Kinetic theory vs. second-order fluid dynamics - WITH conserved particles



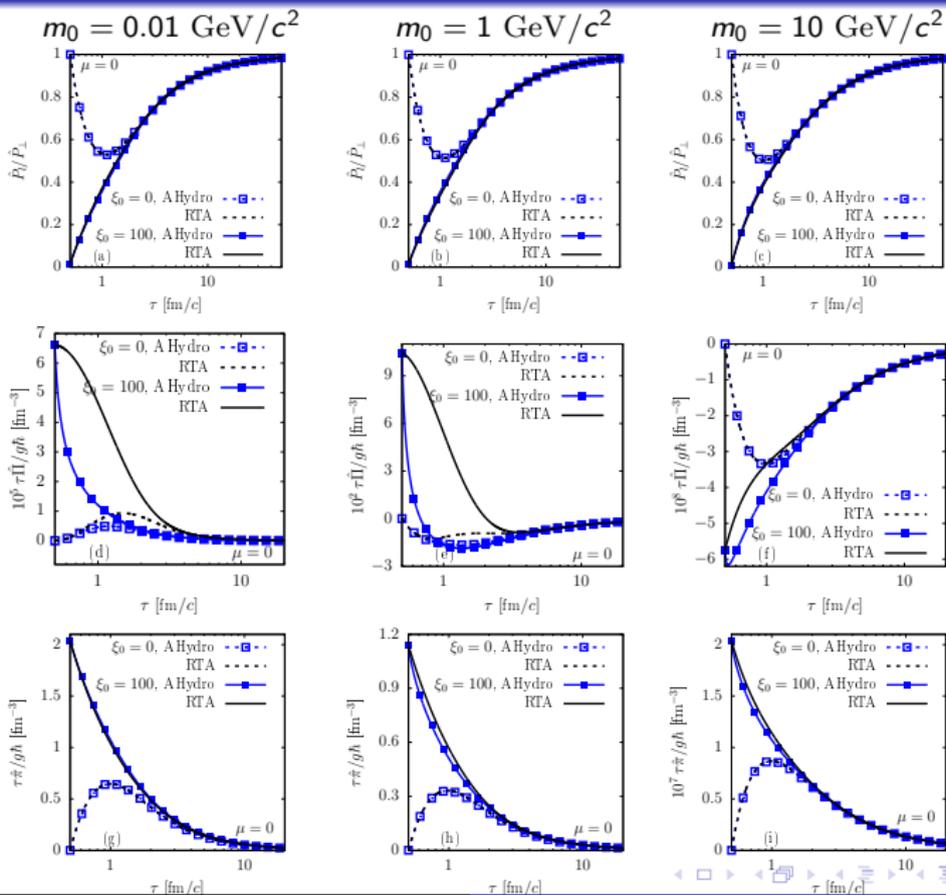
## Kinetic theory vs. anisotropic fluid dynamics - WITH conserved particles



## Kinetic theory vs. second-order fluid dynamics - NO conserved particles



## Kinetic theory vs. anisotropic fluid dynamics - NO conserved particles



# The End