

Kinetic approach to a relativistic Bose-Einstein condensate for massless and massive particles

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June 21, 2016

Outline

Introduction

- The Boltzmann-Uhlen-Uehlenbeck equation
- Our system
- About Bose-Einstein condensation

Relaxation and Equilibrium

- Equilibrium for overpopulated systems
- Equilibrium for underpopulated systems
- Equation of motion

Numerical methods

- Cash-Karp RK45-scheme

Results

- Underpopulated/Critical/Overpopulated case
- About the onset
- Massive case

Conclusion and Outlook

Backup

Non-equilibrium

- ▶ Non-equilibrium thermodynamics is a branch of thermodynamics that deals with physical systems that are not in thermodynamic equilibrium. [Wikipedia]
- ▶ Equilibrium thermodynamics ignores the time-courses of physical processes, in contrast non-equilibrium thermodynamics attempts to describe their time-courses in continuous detail. [Wikipedia]
- ▶ Thermalisation is the process of physical systems reaching thermal equilibrium through mutual interaction. [Wikipedia]

BUU-equation

Relativistic evolution equation of a bosonic (including quantum statistics by **Bose enhancement**) system in non-equilibrium

$$\begin{aligned} \frac{1}{E_1} \left(p_1^\mu \frac{\partial}{\partial x^\mu} + m \frac{\partial}{\partial p_1^\mu} K_1^\mu \right) f_1 &= \frac{1}{2E_1} \int \frac{d^3 \vec{p}_2}{2(2\pi)^3 E_2} \frac{d^3 \vec{p}_3}{2(2\pi)^3 E_3} \frac{d^3 \vec{p}_4}{2(2\pi)^3 E_4} W_{12 \leftrightarrow 34} \\ &\quad \times \left\{ f_3 f_4 (1 + f_1)(1 + f_2) - f_1 f_2 (1 + f_3)(1 + f_4) \right\} \\ &=: \mathcal{C}[2 \leftrightarrow 2]. \end{aligned}$$

$$W_{12 \leftrightarrow 34} := (2\pi)^4 \frac{|M_{12 \leftrightarrow 34}|^2}{\nu} \delta^{(4)}(P_1 + P_2 - P_3 - P_4).$$

Equilibrium f_{eq} ? Detailed balance!

$$\mathcal{C}[2 \leftrightarrow 2] \stackrel{!}{=} 0.$$

Assumptions

- ▶ isotropic system $f(t, \vec{r}, \vec{p}) \rightarrow f(t, \vec{r}, p)$
- ▶ homogeneous system $f(t, \vec{r}, p) \rightarrow f(t, p)$
- ▶ vanishing external forces $K^\mu = 0$

Detailed balance is satisfied by the Bose-Einstein distribution

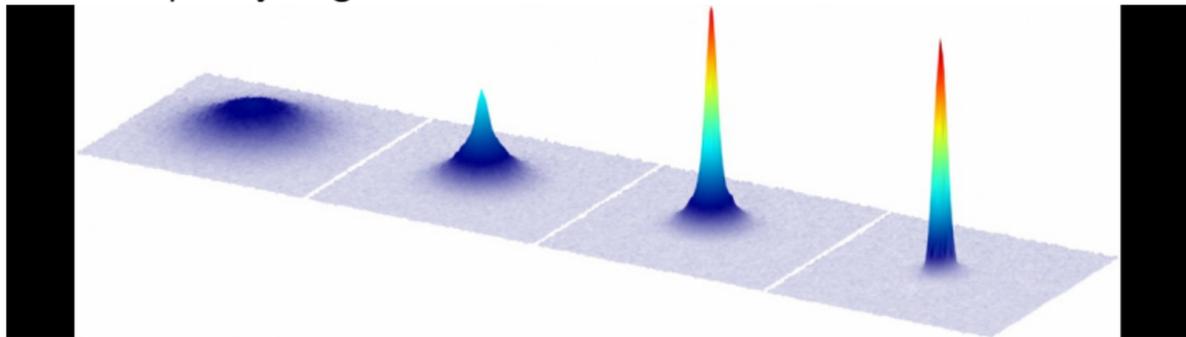
$$f_{\text{eq}}(E_i) = \frac{1}{\exp\left(\frac{E_i - \mu}{T}\right) - 1}.$$

The ground state can become macroscopically large $f_{\text{eq}}(E_0) \gg 1$.
Two cases are considered.

Decreasing the temperature

$$f(E_i) = \frac{1}{\exp\left(\frac{E_i - \mu}{T}\right) - 1} \xrightarrow{T \rightarrow 0} 0 \quad \text{for} \quad E_i > E_0 \geq \mu.$$

→ The occupation number of the ground state $f(E_0)$ becomes macroscopically large



Picture: <http://www.erbium.at/FF/wp-content/uploads/2016/01/FirstErbiumBEC-1250x350.jpg>

Increasing the particle density

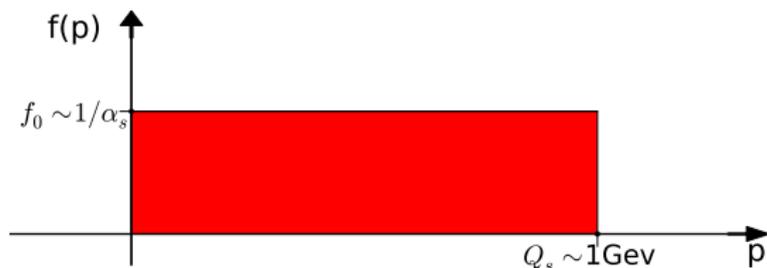
$$f(E_0) = \frac{1}{\exp\left(\frac{E_0 - \mu}{T}\right) - 1} \xrightarrow{\mu \rightarrow m} \infty \quad \text{for} \quad E_0 = m \geq \mu.$$

→ The occupation number of the ground state $f(E_0)$ becomes macroscopic large

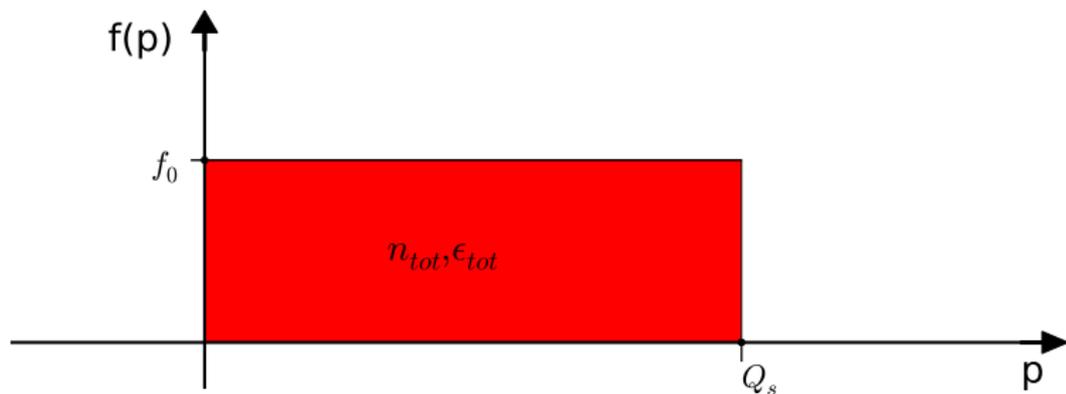
Can be applied to a very early stage of heavy ion collision:

$$f(p) \sim 1/\alpha_s \quad \text{for} \quad p < Q_s$$

$$\simeq f_0 \theta\left(1 - \frac{p}{Q_s}\right)$$



Determining the equilibrium state



$$n_{\text{tot}} = \int_0^{\infty} \frac{dp}{2\pi^2} p^2 f_{\text{init}} = \frac{f_0 Q_s^3}{6\pi^2}$$

$$\epsilon_{\text{tot}} = \int_0^{\infty} \frac{dp}{2\pi^2} p^2 E f_{\text{init}} = \frac{f_0}{16\pi^2} \left\{ Q_s E_{Q_s} (m^2 + 2Q_s^2) + m^4 \log \frac{m}{Q_s + E_{Q_s}} \right\}$$

$$E_{Q_s} := \sqrt{m^2 + Q_s^2}$$

Decompose $f(t, \mathbf{p})$ [Semikoz, Tchakev, arxiv.org/abs/hep-ph/9507306]

$$f(t, \vec{p}) = f_{\text{part}}(t, \vec{p} > 0) + n_c(t)(2\pi)^3 \delta^{(3)}(\vec{p})$$

Red known — Blue unknown — Green fixing

▶ $n_{\text{tot}} = n_{\text{part,eq}} + n_{c,\text{eq}}$ particle (density) conservation

▶ $\epsilon_{\text{tot}} = \epsilon_{\text{part,eq}} + \epsilon_{c,\text{eq}}$ energy (density) conservation

$$\text{▶ } n_{\text{part,eq}} = \int_0^\infty \frac{dp}{2\pi^2} p^2 \frac{1}{\exp\left(\frac{E - \mu_{\text{eq}}}{T_{\text{eq}}}\right) - 1}$$

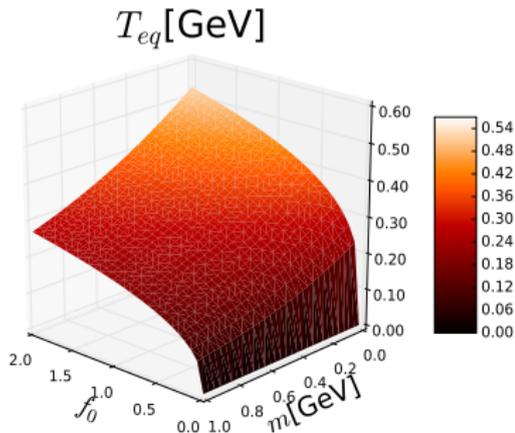
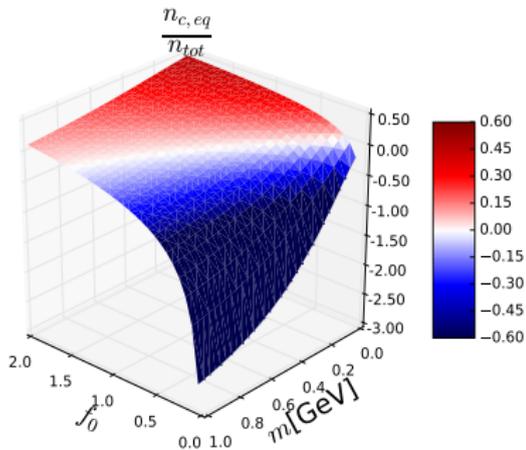
$$\text{▶ } \epsilon_{\text{part,eq}} = \int_0^\infty \frac{dp}{2\pi^2} p^2 \frac{E}{\exp\left(\frac{E - \mu_{\text{eq}}}{T_{\text{eq}}}\right) - 1}$$

▶ $\mu_{\text{eq}} = m$ and $\epsilon_c = n_c m$

Solve for T_{eq} and $n_{c,\text{eq}}$

Numerical solution for the equilibrium values

► $Q_s = 1\text{GeV}$



- The blue shaded area suggest a negative condensate density which is not physical
- **Condition** $\mu_{eq} = m$ does not apply (underpopulated case)

massless case $m = 0$

Equilibrium state is given analytically

$$\mu_{\text{eq}} = 0, \quad \epsilon_{\text{c,eq}} = 0$$

$$\epsilon_{\text{tot,eq}} = \frac{f_0 Q_s^4}{8\pi^2} \stackrel{!}{=} \frac{\pi^2 T_{\text{eq}}^4}{30} = \epsilon_{\text{part,eq}} \longrightarrow T_{\text{eq}} = \sqrt[4]{f_0 15} \frac{Q_s}{2\pi}$$

$$n_{\text{tot,eq}} = \frac{f_0 Q_s^3}{6\pi^2}, \quad n_{\text{part,eq}} = (15f_0)^{\frac{3}{4}} \frac{Q_s^3 \zeta(3)}{2\sqrt{2}\pi^5}$$

$$n_{\text{c,eq}} = \frac{f_0 Q_s^3}{6\pi^2} - (15f_0)^{\frac{3}{4}} \frac{Q_s^3 \zeta(3)}{2\sqrt{2}\pi^5}$$

$$n_{\text{c,eq}} = 0 \longrightarrow f_0 \approx 0.154 \text{ (the critical case)}$$

Equilibrium for underpopulated systems

Solve for T_{eq} and μ_{eq} :

$$n_{\text{tot}} = n_{\text{part,eq}} = \int_0^{\infty} \frac{dp}{2\pi^2} p^2 f_{\text{eq}}(\mu_{\text{eq}}, T_{\text{eq}})$$

$$\epsilon_{\text{tot}} = \epsilon_{\text{part,eq}} = \int_0^{\infty} \frac{dp}{2\pi^2} p^2 E f_{\text{eq}}(\mu_{\text{eq}}, T_{\text{eq}})$$

$$f(t, \vec{p}) \rightarrow f(t, \vec{p}) = f_{\text{part}}(t, \vec{p} > 0) + \underbrace{n_c(t)(2\pi)^3 \delta^{(3)}(\vec{p})}_{=: f_c}$$

- Set of 2 coupled first order differential equation.

Evolution equation for the BEC

$$\begin{aligned} \frac{\partial f_{c,1}}{\partial t} &= \frac{1}{2E_1} \int \frac{d^3 \vec{p}_2}{2(2\pi)^3 E_2} \frac{d^3 \vec{p}_3}{2(2\pi)^3 E_3} \frac{d^3 \vec{p}_4}{2(2\pi)^3 E_4} \\ &\quad \times (2\pi)^4 \frac{|M_{12 \leftrightarrow 34}|^2}{\nu} \delta^4(P_1 + P_2 - P_3 - P_4) \\ &\quad \times \{f_{c,1} f_3 f_4 - f_{c,1} f_2 (1 + f_3 + f_4)\} =: C[1c + 1 \leftrightarrow 2] \end{aligned}$$

Following a integration over \vec{p}_1

$$\frac{\partial n_c}{\partial t} = \int \frac{d^3 \vec{p}_1}{(2\pi)^3} C[1c + 1 \leftrightarrow 2]$$

Evolution equation for the higher modes

$$\begin{aligned}
 \frac{\partial f_1}{\partial t} = & \frac{1}{2E_1} \int \frac{d^3 \vec{p}_2}{2(2\pi)^3 E_2} \frac{d^3 \vec{p}_3}{2(2\pi)^3 E_3} \frac{d^3 \vec{p}_4}{2(2\pi)^3 E_4} \\
 & \times (2\pi)^4 \frac{|M_{12 \leftrightarrow 34}|^2}{\nu} \delta^4(P_1 + P_2 - P_3 - P_4) \\
 & \times \left[\{f_3 f_4 (f_1 + 1) (f_2 + 1) - f_1 f_2 (f_3 + 1) (f_4 + 1)\} \right. \\
 & + \{f_{c,2} f_3 f_4 - f_{c,2} f_1 (1 + f_3 + f_4)\} \\
 & + \{(1 + f_1 + f_2) f_{c,3} f_4 - f_{c,3} f_1 f_2\} \\
 & \left. + \{(1 + f_1 + f_2) f_{c,4} f_3 - f_{c,4} f_1 f_2\} \right] \\
 := & \mathcal{C}[2 \leftrightarrow 2] + \mathcal{C}[1 + 1c \leftrightarrow 2]
 \end{aligned}$$

- ▶ 9 dimensional Integrals are not practical for our numerical approach
- ▶ $\frac{|M_{12 \leftrightarrow 34}|^2}{\nu} \propto s = (P_1 + P_2)^2 \rightarrow$ integrate out every angular dependencies and also the internal momenta \tilde{p}_4

$$\frac{\partial n_c}{\partial t} = n_c \frac{9\lambda^2}{64\pi^3} \int_0^\infty dp_2 \int_0^\infty dp_3 \frac{p_2 p_3}{m_1 E_2 E_3}$$

$$\times [-1 - \epsilon(p_2 - p_3 - \tilde{p}_4) + \epsilon(p_2 + p_3 - \tilde{p}_4) + \epsilon(p_2 - p_3 + \tilde{p}_4)]$$

$$\times (m_1^2 + m_2^2 + 2m_1 E_2) \theta(\tilde{p}_4^2) [f_3 f_4 - f_2 (1 + f_3 + f_4)]$$

$$\frac{\partial f_1}{\partial t} = \underbrace{\mathcal{C}[2 \leftrightarrow 2]}_{\int_{\mathbb{R}^+ \times \mathbb{R}^+}} + \underbrace{\mathcal{C}[1 + 1c \leftrightarrow 2]}_{3 \times \int_{\mathbb{R}^+}, \propto n_c}$$

What we have so far

A **given** initial state with the mass of the particles **m**:

A final **determined** state:

$$f_{\text{init}}(p) = f_0 \theta \left(1 - \frac{p}{Q_s} \right) \xrightarrow{EoM} f_{\text{eq}}(p) = \frac{1}{\exp \left(\frac{\sqrt{p^2 + m^2} - \mu_{\text{eq}}}{T_{\text{eq}}} \right) - 1} + n_{c,\text{eq}} (2\pi)^3 \delta^{(3)}(\vec{p})$$

- ▶ 2 first order coupled integro-differential equations
- ▶ $\rightarrow \frac{|M_{12 \leftrightarrow 34}|^2}{\nu} \propto s = (P_1 + P_2)^2$
 corresponds to $\sigma_{\text{tot}} = \text{const.}$
- ▶ Inclusion of masses condensate particles is possible in contrast to $\frac{|M_{12 \leftrightarrow 34}|^2}{\nu} \propto \text{const.}$ [arXiv:1510.04552]
- ▶ Analytic solution? Researched field - [arxiv:1507.07834]
- ▶ Numerical evaluation!

$$f_{i+1}^{\text{Euler}} = f_i^{\text{Euler}} + k_i$$

$$f_{i+1}^{\text{RK4}} = f_i^{\text{RK4}} + \frac{2825}{27648} k_{i,1} + \frac{18575}{48384} k_{i,3} + \frac{13525}{55296} k_{i,4} - \frac{277}{14336} k_{i,5} + \frac{1}{4} k_{i,6}$$

$$f_{i+1}^{\text{RK5}} = f_i^{\text{RK5}} + \frac{37}{378} k_{i,1} + \frac{250}{621} k_{i,3} + \frac{125}{594} k_{i,4} - \frac{1}{5} k_{i,6}$$

[Transactions on Mathematical Software 16: 201-222, 1990. doi:10.1145/79505.79507]

- ▶ two approximations of order 4 and 5
- ▶ no additional computation time for the second approximation
- ▶ compare the approximations
- ▶ Cash-Karp method involves $h_{\text{new}} = sh_{\text{old}}$

$$s = \left| \frac{\epsilon_{\text{tol}}}{f_{i+1}^{\text{RK5}} - f_{i+1}^{\text{RK4}}} \right|^{\frac{1}{5}}$$

Applying on the Boltzmann-equation

- ▶ $f_{\text{part}}(p)$ is given on a Grid $G := \{p^0, p^1, \dots, p^i, \dots, p^N\}$ with $p^0 < p^1 < \dots < p^i < \dots < p^N$ ($N > 100$)
- ▶ then solve the Boltzmann equation independently for every Grid point (external momenta p^i) by applying the RKCK45 scheme.
- ▶ to evaluate the collision integrals we apply quadrature methods (trapezoidal, Simpson)

Condensation onset

onset:= Starting time of condensation

- ▶ the condensation process $\dot{n}_c \propto n_c$ happens only for $n_c \neq 0$
- ▶ BEC is a phenomena which arises due to fluctuations and are not included in this approach
- ▶ two Possibilities to include condensation are:
 1. initialising with a finite but negligibly small condensate seed $n_c \ll n_{\text{tot}}$
 2. inserting a small condensate seed $n_c \ll n_{\text{tot}}$ when the distribution function reaches a certain point
- ▶ extraction of 2 parameters ($\mu_{\text{eff}}, T_{\text{eff}}$) by fitting the Bose distribution to f_{part} and then inserting the seed when the condition $\mu_{\text{eff}} = m$ is given

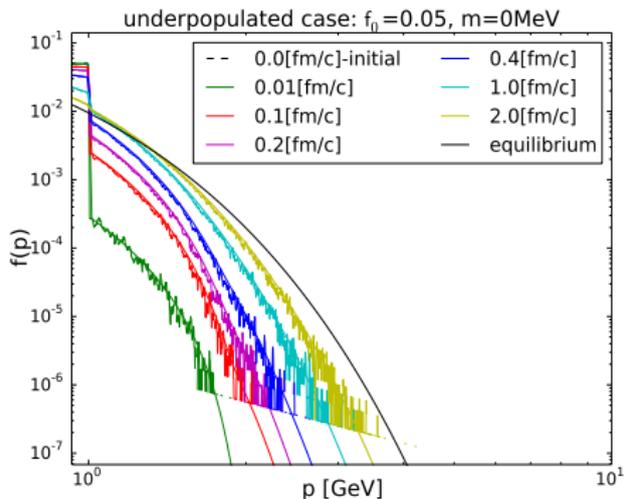
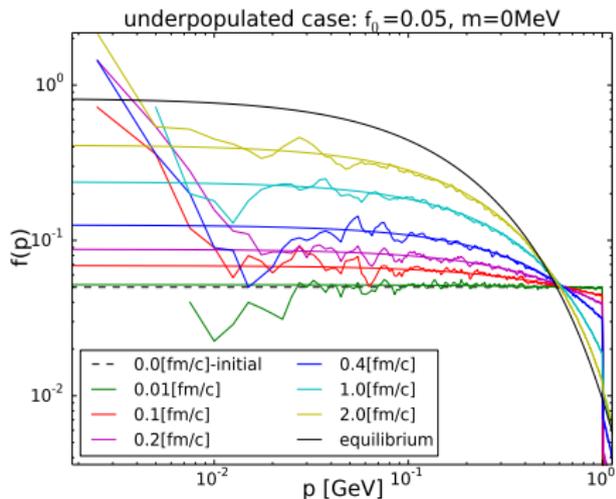
Before The result:

- ▶ Transport code BAMPS [Greiner arXiv:hep-ph/0406278]
(= Boltzmann Approach to Multi-Parton Scatterings)
- ▶ Test particle Ansatz
- ▶ Box calculation

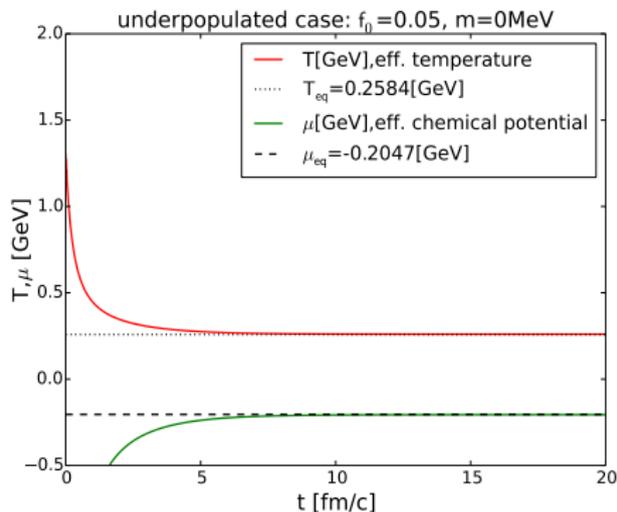
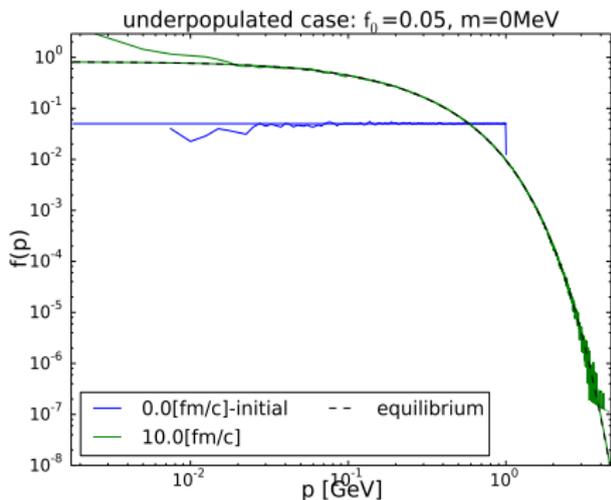
Our code: smooth lines

BAMPS: shaky lines

Underpopulated massless case

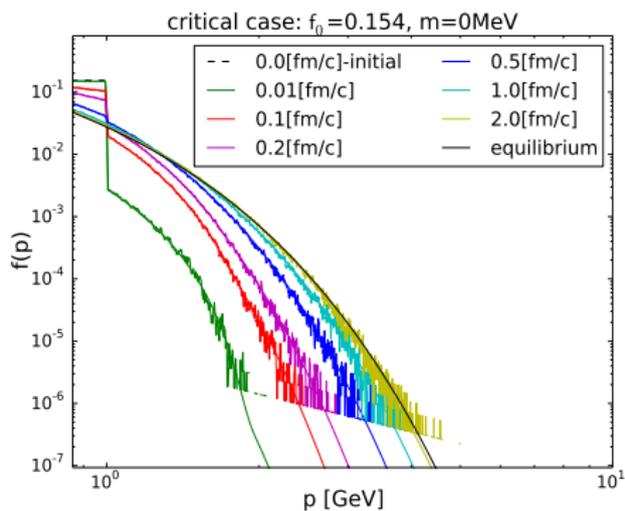
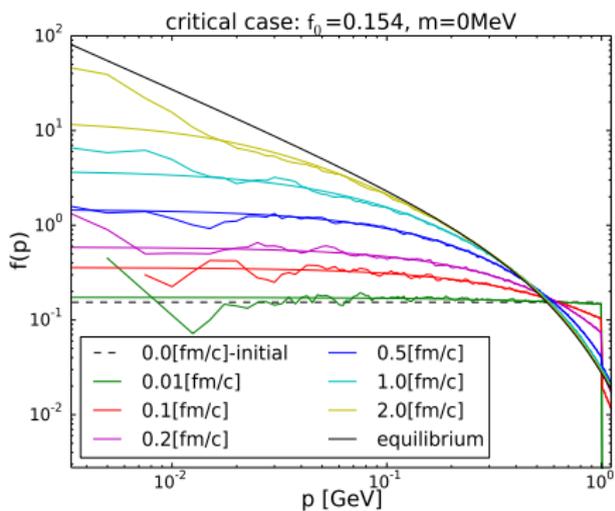


Underpopulated massless case

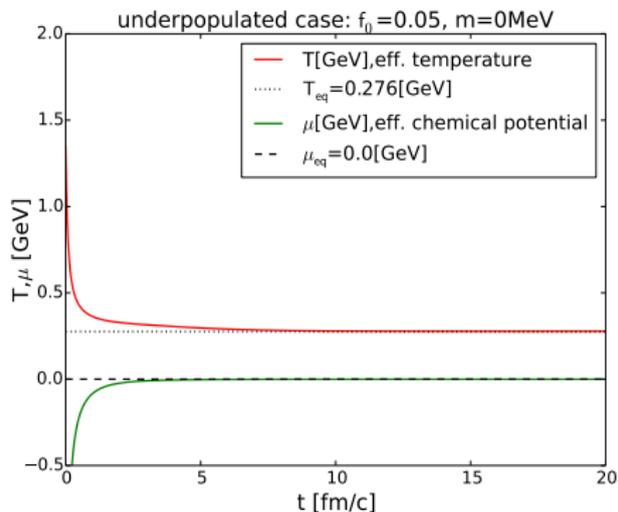
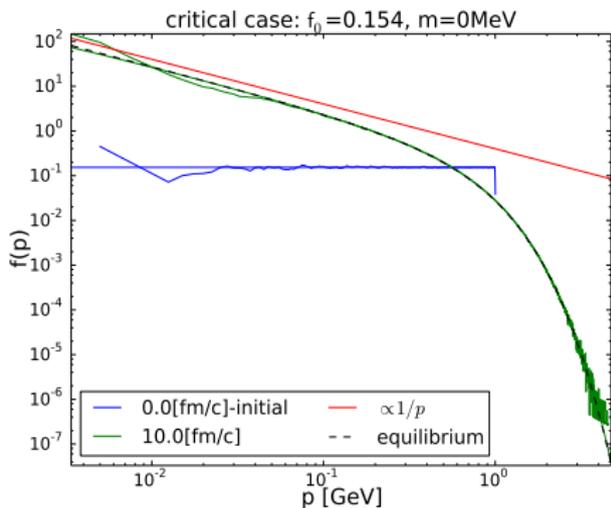


- No condensation since $\mu_{\text{eff}} < m$

Critical massless case



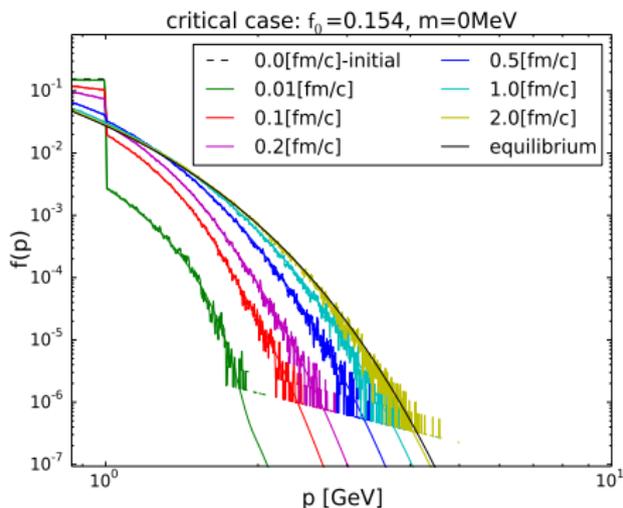
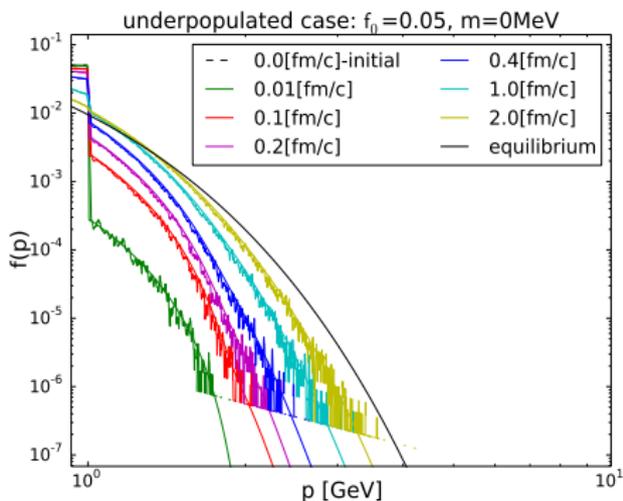
Critical massless case



- ▶ No condensation since μ_{eff} converges to the mass (0GeV)

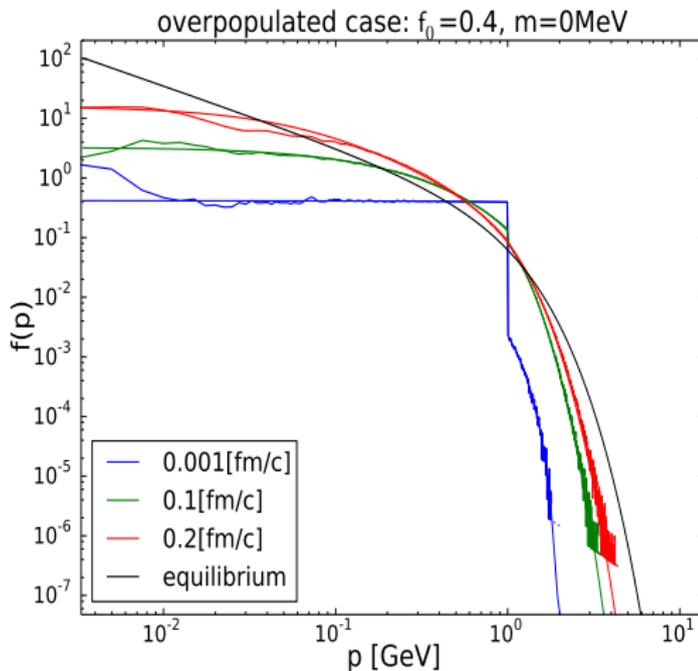
Clip

Thermalisation time



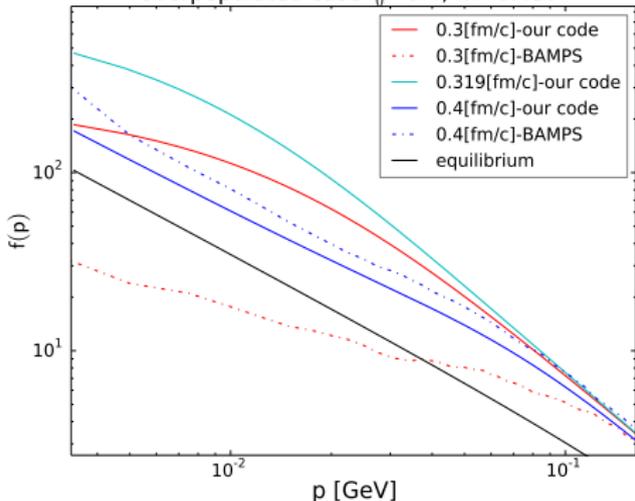
- ▶ focus $f(t = 2.0[\text{fm}/c], p)$ XXXXX
- ▶ Increasing the total particle density leads to a faster thermalisation-consistent

Overpopulated massless case

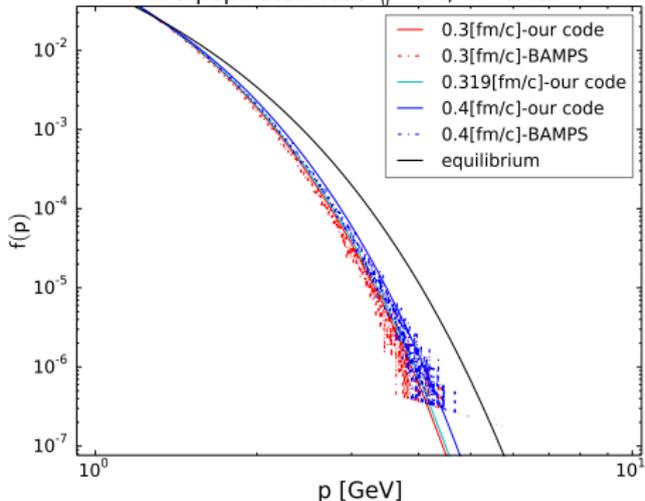


Overpopulated massless case

overpopulated case $f_0=0.4$, $m=0\text{MeV}$

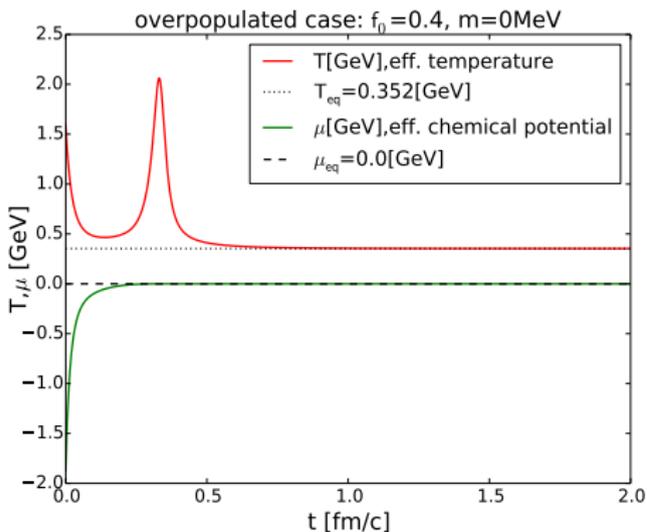
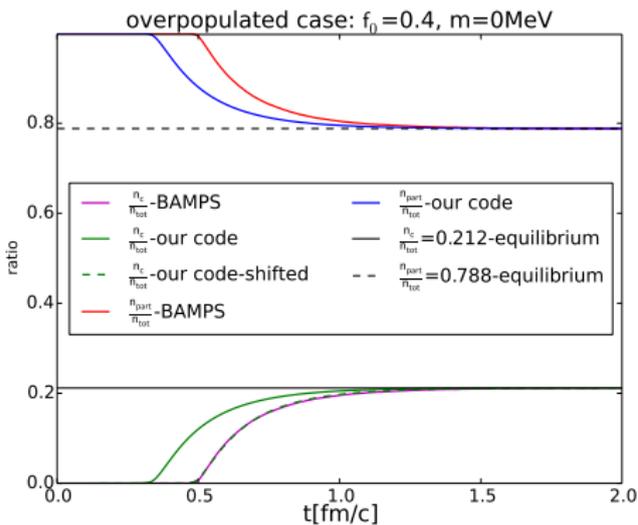


overpopulated case $f_0=0.4$, $m=0\text{MeV}$

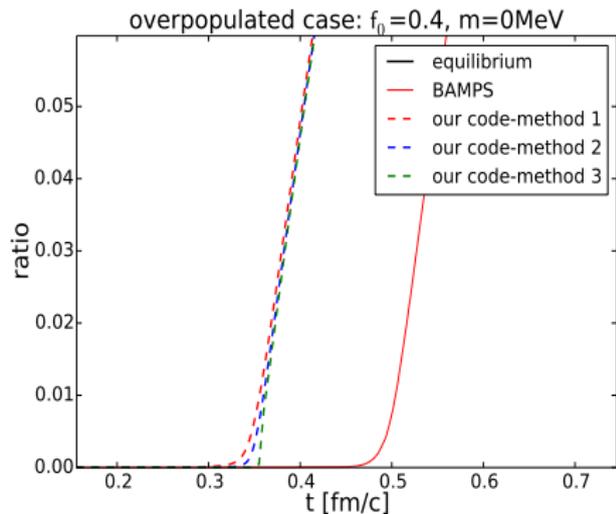
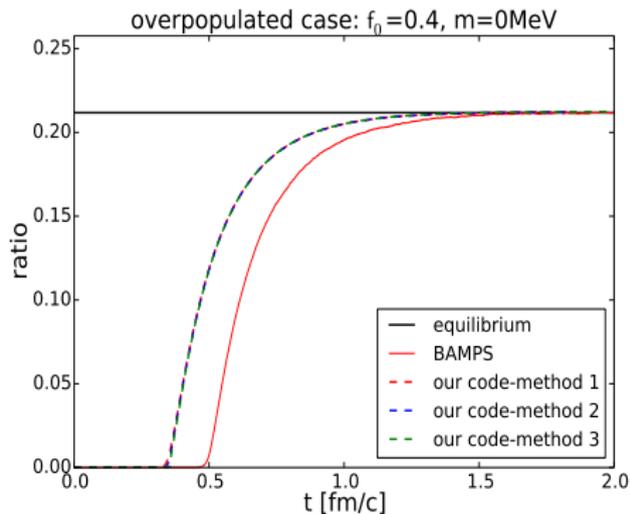


- ▶ Maybe just a minor bug? A time-shift?

Condensate evolution



► About the onset (onset:= Starting time of condensation)



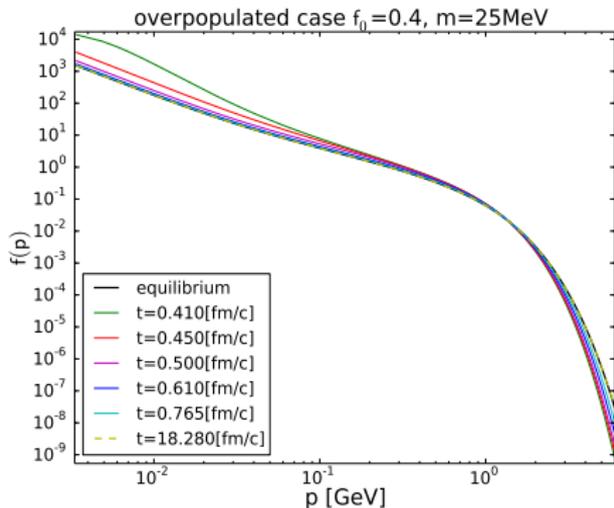
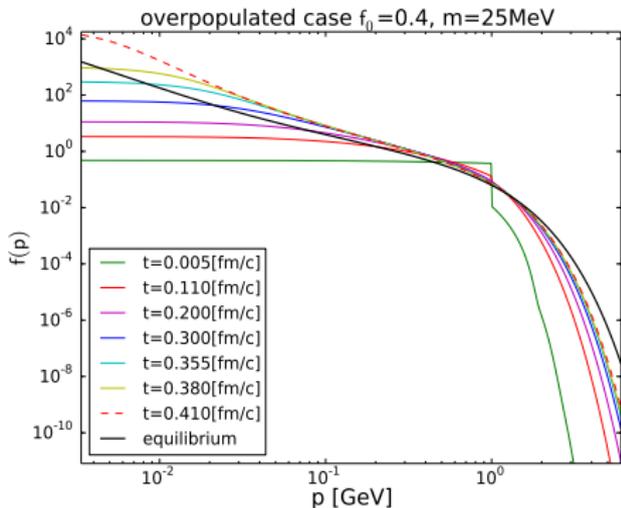
- **method 1**: Starting with condensate seed
- **method 2**: Inserting a seed when $\mu_{\text{eff}} = 0$
- **method 3**: Any time

Clip

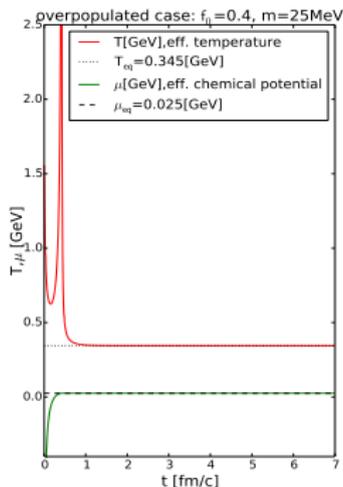
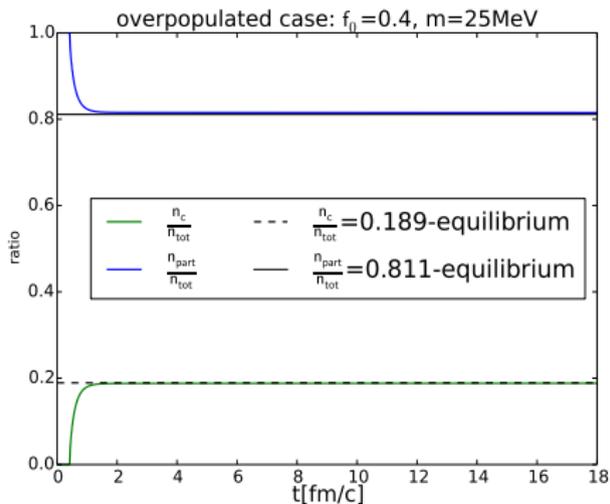
Overpopulated massive $m = 25\text{MeV}$ case

Before the onset

After the onset



Overpopulated massive $m = 25\text{MeV}$ case



Conclusion and Outlook

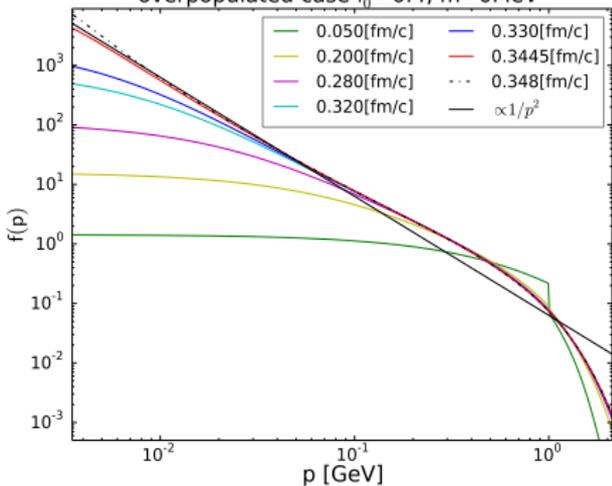
- ▶ All our simulations thermalize into equilibrium.
- ▶ The evolution of our system without condensate $\dot{f}_1 = C[2 \leftrightarrow 2]$ is in a good agreement with BAMPS.
- ▶ Overpopulated systems differs by a time shift later while Equilibration is still given
- ▶ The different onset methods are equivalent.

- ▶ A new BAMPS run is going on to set the onset manually at a earlier time.
- ▶ Further numerical tests have to include a comparison with the analytic solution for a classical system in [\[arxiv:1507.07834\]](https://arxiv.org/abs/1507.07834):
- ▶ Adapting this scheme for longitudinal expanding (anisotropic) systems (Bjorken coordinates).

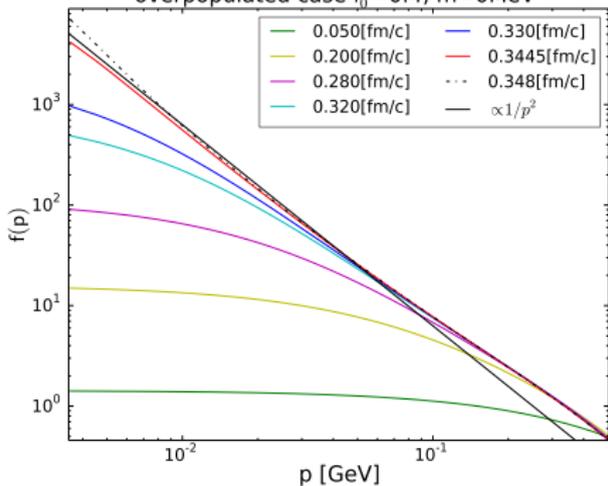
Thank You!



overpopulated case $f_0=0.4$, $m=0\text{MeV}$



overpopulated case $f_0=0.4$, $m=0\text{MeV}$



Eulers method

$$\frac{df}{dt} = C(t, f) \quad , \quad f(t_0) = f_0$$

$$\left. \frac{df}{dt} \right|_{t=t_i} = C(t_i, f_i)$$

$$f_t = f_0 + C(t_0, f_0)(t - t_0)$$

$$f_{i+1} = f_i + C(t_i, f_i)(t_{i+1} - t_i)$$

Using a uniform step size $h := t_{i+1} - t_i = \text{const.}$ and substituting $k_i = hC(t_i, f_i)$ one ends up with:

$$f_{i+1}^{\text{Euler}} = f_i + k_i$$

Cash-Karp RK45 -scheme

starting point: 2 approximations whereby both approximations need the evaluation of the following six values

$$k_{i,1} = hC(t_i, f_i)$$

$$k_{i,2} = hC\left(t_i + \frac{1}{5}h, f_i + \frac{1}{5}k_{i,1}\right)$$

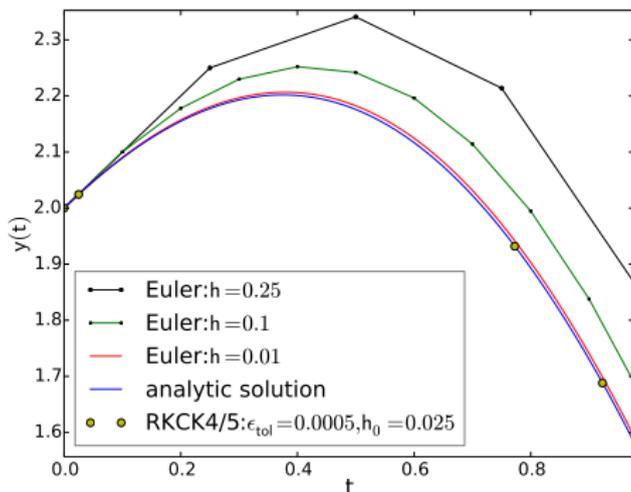
$$k_{i,3} = hC\left(t_i + \frac{3}{10}h, f_i + \frac{3}{40}k_{i,1} + \frac{9}{40}k_{i,2}\right)$$

$$k_{i,4} = hC\left(t_i + \frac{3}{5}h, f_i + \frac{3}{10}k_{i,1} - \frac{9}{10}k_{i,2} + \frac{6}{5}k_{i,3}\right)$$

$$k_{i,5} = hC\left(t_i + h, f_i - \frac{11}{54}k_{i,1} + \frac{5}{2}k_{i,2} - \frac{70}{27}k_{i,3} + \frac{35}{27}k_{i,4}\right)$$

$$k_{i,6} = hC\left(t_i + \frac{7}{8}h, f_i + \frac{1631}{55296}k_{i,1} + \frac{175}{512}k_{i,2} + \frac{575}{13824}k_{i,3} + \frac{44275}{110592}k_{i,4} - \frac{253}{4096}k_{i,5}\right)$$

illustration



Given an first order ODE with the following IVP

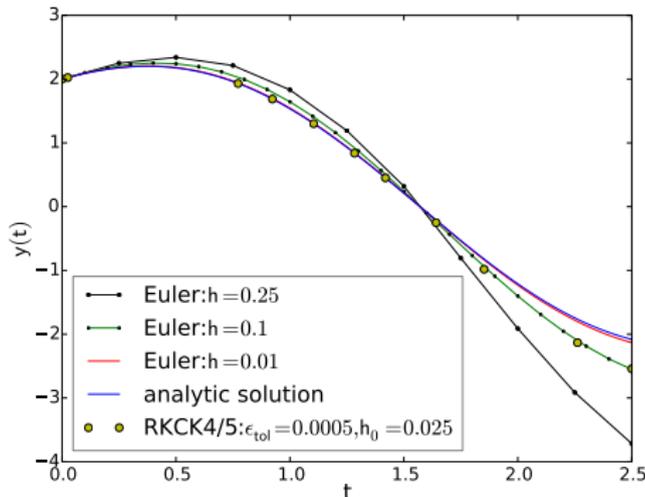
$$\frac{df}{dt} = \cos^2(t) + \tan(t)f$$

$$f(0) = 2$$

and its analytic solution

$$f(t) = [\sin(t)+2] \cos(t)$$

illustration



Given an first order ODE with the following IVP

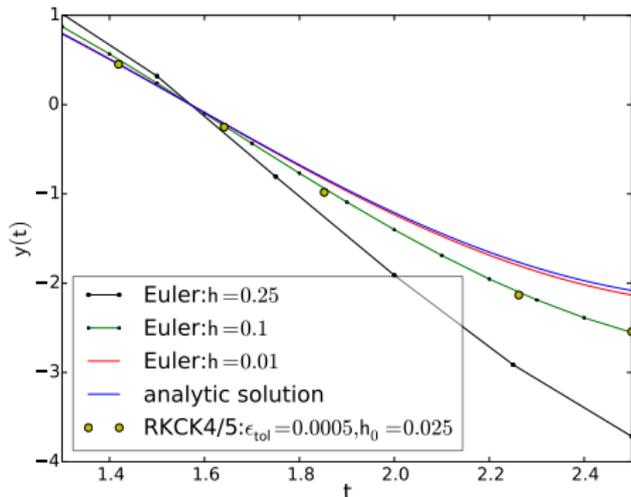
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illustration



Given an first order ODE with the following IVP

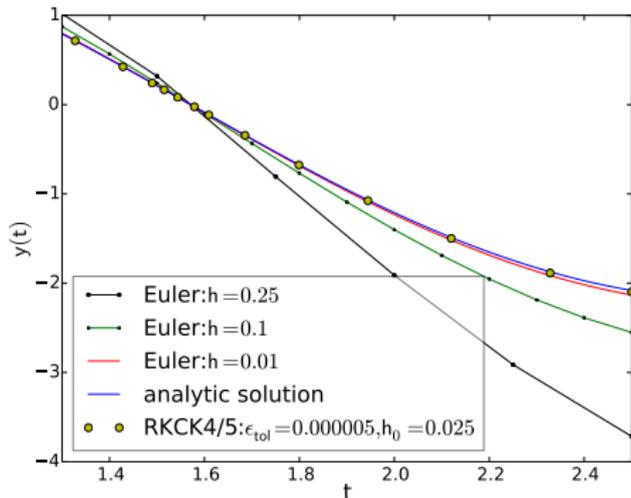
$$\frac{df}{dt} = \cos^2(t) + \tan(t)f$$

$$f(0) = 2$$

and its analytic solution

$$f(t) = [\sin(t)+2] \cos(t)$$

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Overpopulated massless case $f_0 = 0.4$

- ▶ our code initialised with a seed and in BAMPS the condensate evolution is prohibited

