A Theory of Modern ART M. Hanauske¹, A. Bhattacharyya^{1,2}, L.M. Satarov^{1,3}, I.N. Mishustin^{1,3,4}, J. Schaffner-Bielich^{1,5}, D. Zschiesche¹, H. Stöcker¹, and W. Greiner¹ ¹Institut für Theoretische Physik, J. W. Goethe-Universität, D-60054 Frankfurt, Germany ² Department of Physics, Scottish Church College, 1 & 3 Urquhart Square, Calcutta - 700 006, India ³The Kurchatov Institute, 123182 Moscow, Russia ⁴The Niels Bohr Institute, DK-2100 Copenhagen O, Denmark ⁵Department of Physics, Columbia University, 538 West 120th Street, New York, NY 10027, USA Yang-Mills Theory ⇔ Allgemeine Relativitäts Theorie Allgemeine Relativitäts Theorie (General Relativity) QuantumCromoDynamic The QuantumCromoDynamic (QCD) is a $SU(3)_{(c)}$ -Yang-Mills Gauge The field equations of the gravitational force (Einstein-Equation) follow Theory. This force acts between particles which have a Color-Charge. from the variational principle: Four Interactions are presently known: The interaction particle is a massless Gluon which has itself a $\delta_{q_{\mu\nu}} \int (R + \kappa \mathcal{L}_{Matter}) \sqrt{-g} d^4x = 0$ • The Electrodynamic Force Color-Charge. This Force acts between charged or magnetized Particles and is des- \mathcal{L}_{Matter} : Lagrangian density of all matter and fields (The gra-Barionic particles are composed of three valence quarks. cribed by a U(1)–Yang-Mills Gauge Theory. vitational field itself is the only exception; it is not Red Quark • The Week Nuclear Force included here). : This invariant is called the Ricci-Scalar. It follows is responsible for the β -decay of nuclei. It is described by a SU(2)– from the Ricci-Tensor $R_{\mu\nu}$ by summation over the Yang-Mills Gauge Theory. spacetime indices. It describes roughly how strong • The Strong Nuclear Force Green Quark the spacetime is curved. This Force is the strongest Force, it binds the hadrons inside a nuclei : Square root of the determinant of the spacetime meand is also responsible for the confinement of Quarks. The fundamental description of this Force is known as the Quantum Cromo Dyna-Gluons : Variation of the metric $q_{\mu\nu}$. mic which is a SU(3)–Yang-Mills Gauge Theory. Blue Quark Einstein–Equation ↓ Gravity The gravitational Force is the weakest Force. In comparison to all the other Forces it is not possible to describe this Interaction by a $R_{\mu\nu} - \frac{1}{2}R\,g_{\mu\nu}$ All of the fundamental particles (quarks and Gluons) inside this barion Yang-Mills Gauge Theory. The gravitational Force was reformulated have color charges. The most important property of the by Einstein as a curvature of SpaceTime and he described it within QCD-Interaction is the fact, that the gluonic interaction is so strong for The Einstein–Equation means the following: the Allgemeine Relativitäts Theorie (Genearal Relativity). low momentum that the quarks are confined inside the barion. The property of CONFINEMENT curves the color-space in such a strong It is mathematically possible to unify the Electrodynamic, the Week and How does the spacetime cur- if we have somewhere an way that no color charges can be found by an outside observer. This the Strong Force within the Grand Unified Theory. This GUT theory is a amount of energy-momentum. means that the different color charges of the fundamental particles SU(5)– Yang-Mills Gauge Theory. How do the gravitational fields if we have somewhere an The important expressions of Yang-Mills Gauge Theories and General Relativity (ART) have similarities and differences: cancel each other, so that the resulting barion seems to be white! amount of energy-momentum. look like $(g_{\mu\nu}, \Gamma^{\mu}_{\alpha\beta})$: SU(3)–Color indices A, BYang-Mills Theory | ART (General Relativity) $G_{\beta A}{}^{B}$: Gluons $D_{\beta}v^{\alpha} = \partial_{\beta}v^{\alpha} + \Gamma^{\alpha}_{\sigma\beta}v^{\sigma}$ = : The covariant derivative contains two color in- $D_{\beta a}{}^{b} = \partial_{\beta} 1_{a}{}^{b} + ig A_{\beta a}{}^{b}$ $R^{\delta}_{\ \mulphaeta}\,v^{\mu}=\left[D_{lpha},D_{eta} ight]v^{\delta}$ $F_{\alpha\beta a}^{b} = \frac{1}{ia} \left[D_{\alpha a}^{c}, D_{\beta c}^{c} \right]$ dices (A, B) and one spacetime index β $R^{\delta}{}_{\mu\alpha\beta} = \Gamma^{\delta}{}_{\mu\alpha|\beta} - \Gamma^{\delta}{}_{\mu\beta|\alpha} + \Gamma^{\delta}{}_{\nu\beta}\Gamma^{\nu}{}_{\mu\alpha} + \Gamma^{\delta}{}_{\nu\alpha}\Gamma^{\nu}{}_{\mu\beta}$ $= A_{\beta a}^{b}{}_{|\alpha} - A_{\alpha a}^{b}{}_{|\beta}$: The Quarks are described by a Dirac-Spinor having one color and one flavour index $\mathcal{L}_{YM} = \frac{1}{4} F_{\mu\nu a}{}^{b} F^{\mu\nu}{}_{a}{}^{b} \mathcal{L}_{G} = R + (c_{1} R_{\mu\nu} R^{\mu\nu} + ...$ ≡0 for ART QCD-motivated Models of the Strong Interaction ONLINE- Create your own Compact Star -ONLINE Spacetime inside a Compact Star Calculating properties of compact stars is now easy: For high Temperatures (T > 170 MeV) or high densities $(\rho_B > 3 \rho_0)$ We have to solve the full Einstein Equation: All you need is to choose an EoS from the left hand side and a hadronic matter is expected to be deconfined. In such a deconfined $R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = 8\pi \kappa T_{\mu\nu}$ We restrict the system to be static and spherically symmetric. Moreover computer code which binds the star by gravity (curvature). matter (Quark-Gluon Plasma) the quarks and gluons are quite free; the star should not rotate and should not have a charge. they are not restricted to their own hadron. Tollman - Oppenheimer Equation of State To calculate the properties of compact stars one needs the Equation of - Volkoff Equation We can make the following Ansatz of the metric: State (EoS) of hadronic matter at low temperatures (T < 1 MeV) and high densities. Until now it is not possible to extract an EoS strait forward from QCD; therefore we took the following effective QCD-motivated models: Models for Hadronic Matter **Ansatz of the Energy-Momentum Tensor:** • Relativistic Mean Field Models (RMF) Create your own neutron TM1, TM1YY tar: (Online Calculation !!) • static and spherical symmetric : $T^{\mu\nu}(x^{\alpha}) \to T^{\mu\nu}(r)$ NLZ, NLZYY • no sheer forces: $T^{\mu\nu} \equiv 0$, if $\mu \neq \nu$ • SU(3) Chiral Hadron Model (CHM) Equation of states: • EMT of an ideal fluid **General properties:** Models for Quark Matter Chirales_Model $T^{\mu\nu} = -p g^{\mu\nu} + (p+\epsilon) u^{\mu} u^{\nu}$ $u^{\mu} = \frac{dx^{\mu}}{d\tau}$: local 4-velocity MIT-Bag Model • Energydensity-profiles inside the star Central pressure • Pressure-profiles inside the star [Mev/fm^3]: • Nambu-Jona-Lasinio Model (NJL) Mass of the star $= (\frac{1}{\sqrt{g_{tt}}}, 0, 0, 0)$: fluid with no motion 50 • Massive Quasi Particle Model (MQM) Models for Hybrid Matter Inserting $T^{\mu\nu}$ and $g_{\mu\nu}$ in the Einstein Equation one results in a system of The Quark-Hadron Phase Transition differential equations called the Tollman - Oppenheimer - Volkoff Equation • All Models where the Gibbs Construction is possible $g_{rr} = -\left(1 - \frac{2m(r)}{r}\right)^{-1}$ **Curvature of spacetime inside** Models for Matter with and around the star: other types of Phase Transitions $m(r) := 4\pi \int_{-\infty}^{\infty} \epsilon(r')r'^2 dr', \quad m(R) = M$ • Phase Transition to Hyperon Matter Metric component g00 Metric component g11 $rac{d u}{dr} \;=\; rac{m(r)+4\pi\,r^3p(r)}{r(r-2m(r))}, \quad u(R)=\ln\left(1-rac{2M}{r} ight)$ http://www.th.physik.uni-frankfurt.de/~hanauske/index.html $\frac{dp}{dr} \ = \ -(p(r) + \epsilon(r)) \frac{d\nu}{dr}, \quad p(R) = 0$ **Neutron Stars Hybrid Stars Quark Stars Black Holes** Neutron Star Neutron Star Quark Star Between 2 M and Normal Nuclear Mixed Phase Huge core of Crust Crust Crust 3 M only exotic Matter deconfined $\epsilon > 3 \epsilon_0$ with a Density: 106 Density: 10 matter can be $\simeq \epsilon_0 := 2.51 \cdot 10^{14}$ thin nuclear Quark-Gluon $g/cm^3 < \epsilon \ll \epsilon_0$ $g/cm^3 < \epsilon \ll \epsilon_0$ found. g/cm³ Plasma with up matter mantel Mainly down and Schwarzschild composed out of strange quarks Singularity at $R_S = 2 M$ Hyperd little fractions of Possibly a Matte protons (p), core c Real Singularity electrons (e) and deconfine Quark-Gluo composed out myons (μ) $p,n,e,\mu,\Lambda,\Sigma,\Xi.$ Plasma $\epsilon > 5 \epsilon_0$ Properties of Strange Quark Stars Properties of Hybrid Stars Properties of Neutron Stars Properties of Black Holes within the NJL Model The Schwarzschild Black Hole Metric component g_{00} **Density Profiles** Particle Cocktail Gibbs Construction Restrictions: Spherical symmetry, no rotation, no char-**Equation of State** Particle Cocktail - M = 1.63 M₂ ge and $T_{\mu\nu}=0$ ---· M = 1.60 M NLZYY-B180 ·· M = 1.23 M The Einstein Equation reduces under these restricti- $---\cdot \xi = 0.5$ ons to $R_{\mu\nu}=0$ and one can find a coordinate system --- HCM with the following general metric: 4 6 8 10 12 14 Radius [km] Particle Cocktail within the Hadronic Chiral Model Inserting this metric into the Einstein Equation one $\rho_{\rm B}/\rho_{\rm 0}$ $ho_{\mathsf{B}}/ ho_{\mathsf{0}}$ (HCM): can proof the following results: Mass - Radius Properties (NLZYY-Bag Model): $\lambda(r,t)$ and $\nu(r,t)$ are not time dependent (Birkhoff) ; Density Profiles within the NJL Model: $\lambda(r) = -\nu(r), \ \ e^{\nu(r)} = 1 - \frac{2\,M}{r}$ (M is the mass of the black hole); Schwarzschild singularity $R_S := 2 M$; Real (in----- HCM trinsic) singularity at r=0 \longrightarrow NJL_{SU(3)} 15 Spacetime Diagram of the Schwarzschild metric 13 13 ⁴=178 MeV ^t=180 MeV 1600 B^{1/4}>190 MeV 0.4 0.6 0.8 1.0 1.2 1.4 1.6 1.8 M/M_{\odot} 10 12 Radius (km) - Particle energy E_i for Σ^i Twin Star Properties (NLZYY-Bag180 Model): Mass - Radius Properties within the NJL Model: Total number of Barions Twin star Profiles At the Schwarzschild singularity space and time chan-Mass - Radius Properties within the HCM: ges into each other. The result is that all fields and matquark star collapse to a black hole 1.78e+57 ter with r < 2 M need to fall into the real singularity Neutron Stars 14.5 at r=0; no way to get out! 1.76e + 57Neutron, Hybrid and Quark Stars with masses greater 14.0 than the maximum mass M_{max} are unstable and need to collapse to Twin Stars or Black Holes. During the Strange Quark Stars ∉ twin star collapse fast collapse more and more matter in the stars interior deconfines. unstable Unfortunately it looks like that during the stage when 12.0 1.35 1.37 1.38 ---- HCM the quarks mean to get really free and deconfined Radius [km] $\mathsf{M}/\mathsf{M}_{\odot}$ - NJL_{SU(3)} they have crossed the event horizon, they can not 11.5 show their color charge to the outer world – the Black 0.6 1.0 1.4In the interior of Hybrid Stars occurs a Quark-Hadron $\mathsf{M}/\mathsf{M}_{\scriptscriptstyle{G}}$ 0.2 0.4 0.6 0.8 1.0 1.2 1.4 1.6 1.8 Hole has macroscopically confined the Quark-Gluonphase transition. Within a small parameter range Plasma A Black Hole with a typical mass of a Neutron or Twin Stars are possible. A Twin Star collapse could Quark Stars are much smaller than Neutron Stars Neutron Stars have radii of $R \approx 11 - 14$ km emit enough energy to explain Gamma-Ray burst; and Masses M $< 1.65 M_{\odot}$ Quark Star has a Schwarzschild radius of $R \approx 4$ km. $R \approx 8 \text{ km}$ Gravitational wave emission is also possible. and their limiting Mass is smaller M $< 1.3 M_{\odot}$ Astrophysical Observables for the Quark-Gluon Plasma Quark Stars are smaller, the observation of a Twin Star would be the proof of having a phase transition, the observation of a sub millisecond Pulsar, the 'Spin-Up' effect, the observed 'Spin-Down' Effect in LMXB, Gamma-Ray Burst and gravitational wave emission from a Twin Star collapse, Supernova–Hypernova Observed Astrophysical Data Pulsar Glitches **Pulsar Periods** Other Quark Star Candidates Ouark Star obseved !?! Neutron Star masses X-Ray I 1975 Glitch of the Crab

