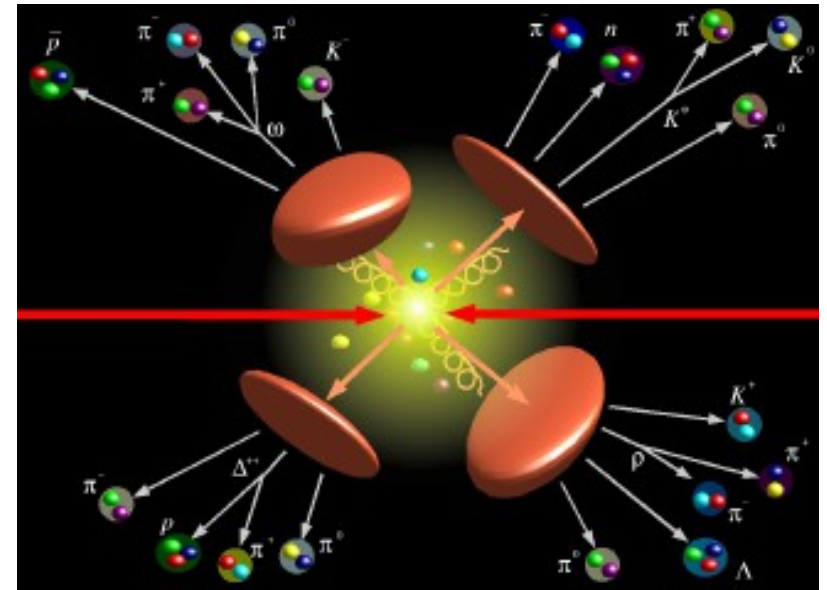
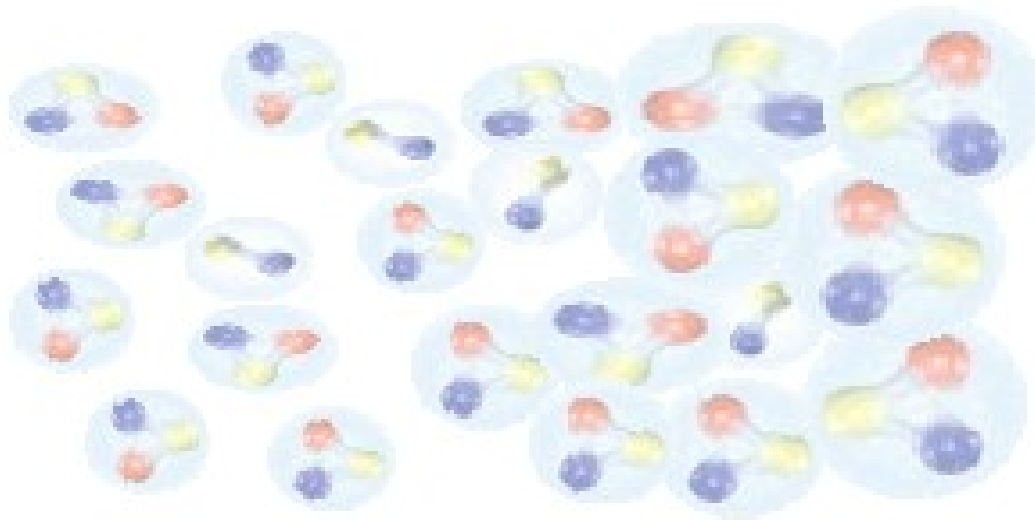


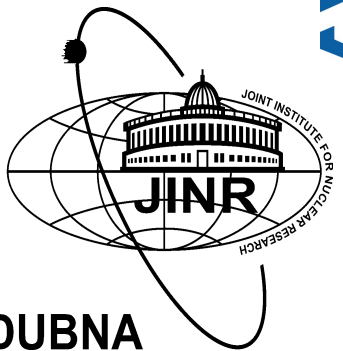
# Baryon stopping signal for a Phase Transition in Heavy-Ion Collisions

David Blaschke

University of Wroclaw, Poland & JINR Dubna, Russia



Transport Theory Seminar, FIAS Frankfurt, November 11, 2015



# Baryon stopping signal for a Phase Transition in Heavy-Ion Collisions

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## 1. Introduction:

The baryon stopping signal for a 1<sup>st</sup> order phase transition

## 2. Robustness: The dependence on experimental cuts

Based on arxiv:1504.03992 [with Yu. Ivanov]

## 3. Further developments:

Particization, Detector response, UrQMD, new class EoS

High-Density Matter Seminar, BLTP JINR Dubna, June 24, 2015



# Strategy towards event simulations testing PT signal

## Two alternative approaches:

### I) Direct approach based on transport codes:

Particle trajectories are followed;

Properties of the medium are encoded in propagators and cross sections

→ UrQMD (Aichelin et al.),

→ PHSD (Bratkovskaya, Cassing, et al.),

→ PHSD + SACA (Bratkovskaya, Aichelin, LeFevre, et al.)

### II) Hybrid approach:

Joins hydrodynamic evolution of a (multi-)fluid system described by an **EoS** with

Particle transport via a procedure called “**particlization**” (Karpenko)

Particularly suitable for studying effects of a strong phase transition in model EoS

a) Sandwich: UrQMD + hydro + hadronic cascade (H. Petersen et al.)

→ PT in hydro stage only

b) **3-fluid hydro**dynamics (Ivanov) + particlization (Karpenko)

→ PT in baryon stopping regime already!

(main difference to sandwich; appropriate for energy range of NICA / CBM)

Both approaches provide the inputs for the simulation of the **detector response**  
(GEANT-MPD: Rogachevsky, Voronyuk, Batyuk, Wielanek, et al.)

# Hydrodynamic modelling for NICA / FAIR

**More complicated for lower energies:**

- baryon stopping effects,
- finite baryon chemical potential,
- EoS unknown from first principles

We want to simulate the effects of, and ultimately discriminate different EoS/PT types  
The model has to be coupled to a detector response code to simulate detector events



taken from: MADAI.us

Initial  
state



3-fluid hydro,  
(Yu. Ivanov)

hydrodynamic  
evolution



adapt the procedure  
from existing hybrid model  
(Iu. Karpenko)

particlization

hadronic  
corona

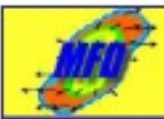


(optionally) cascade:  
PHSD, UrQMD, etc  
(E. Bratkovskaya,  
H. Petersen)

detector  
response



GEANT  
MPD, BM @N  
(O. Rogachevsky,  
V. Voronyuk, et al.)



# 3-Fluid Dynamics

Baryon Stopping

JINR, 24.08.10

Model

Rapidity Density

Fit

Reduced curvature

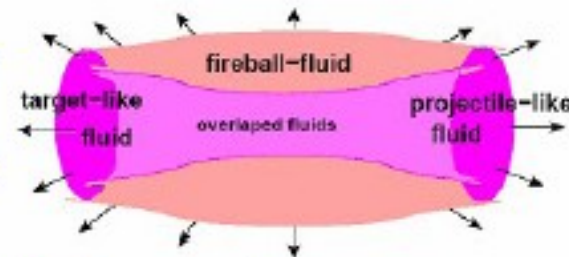
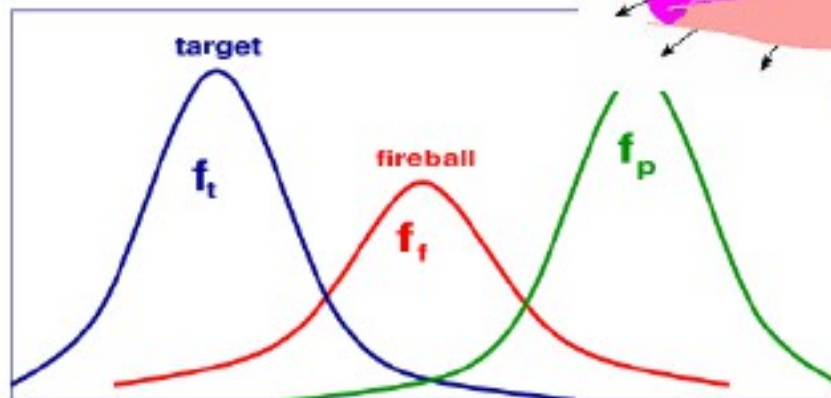
Trajectories

Crossover

Summary

Produced particles populate mid-rapidity  
 $\Rightarrow$  **fireball** fluid

distribution function



momentum along beam

**Target-like fluid:**

$$\partial_\mu J_t^\mu = 0$$

Leading particles carry bar. charge

$$\partial_\mu T_t^{\mu\nu} = -F_{tp}^\nu + F_{ft}^\nu$$

exchange/emission

**Projectile-like fluid:**

$$\partial_\mu J_p^\mu = 0,$$

$$\partial_\mu T_p^{\mu\nu} = -F_{pt}^\nu + F_{fp}^\nu$$

**Fireball fluid:**

$$J_f^\mu = 0,$$

Baryon-free fluid

$$\partial_\mu T_f^{\mu\nu} = F_{pt}^\nu + F_{tp}^\nu - F_{fp}^\nu - F_{ft}^\nu$$

Source term      Exchange

The **source term** is delayed due to a formation time  $\tau \sim 1 \text{ fm/c}$

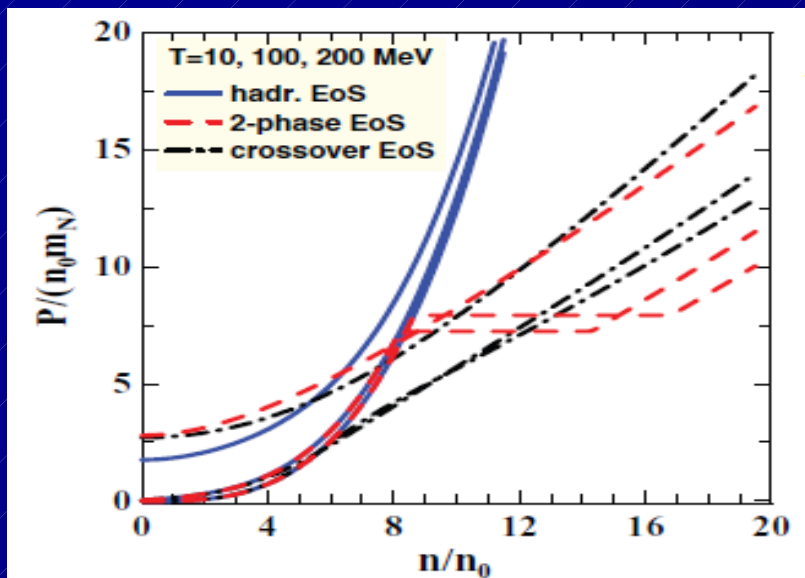
**Total energy-momentum conservation:**

$$\partial_\mu (T_p^{\mu\nu} + T_t^{\mu\nu} + T_f^{\mu\nu}) = 0$$

<http://theory.gsi.de/~ivanov/mfd/>

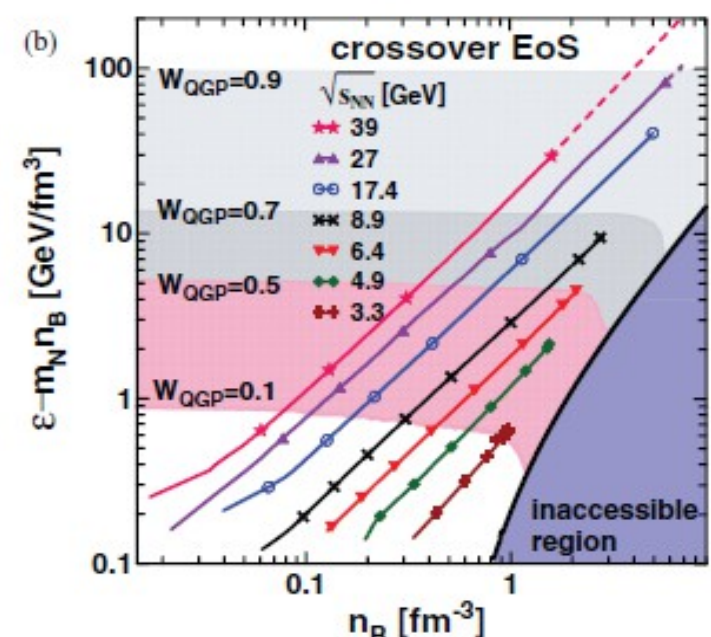
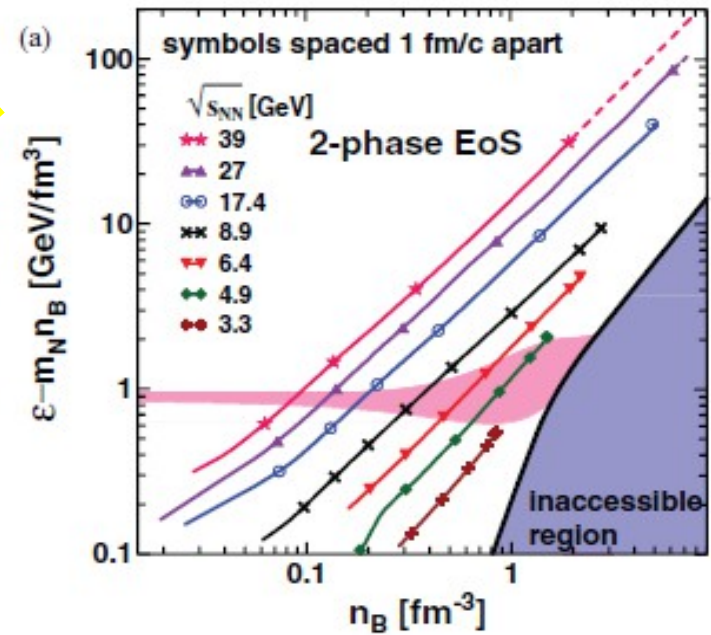
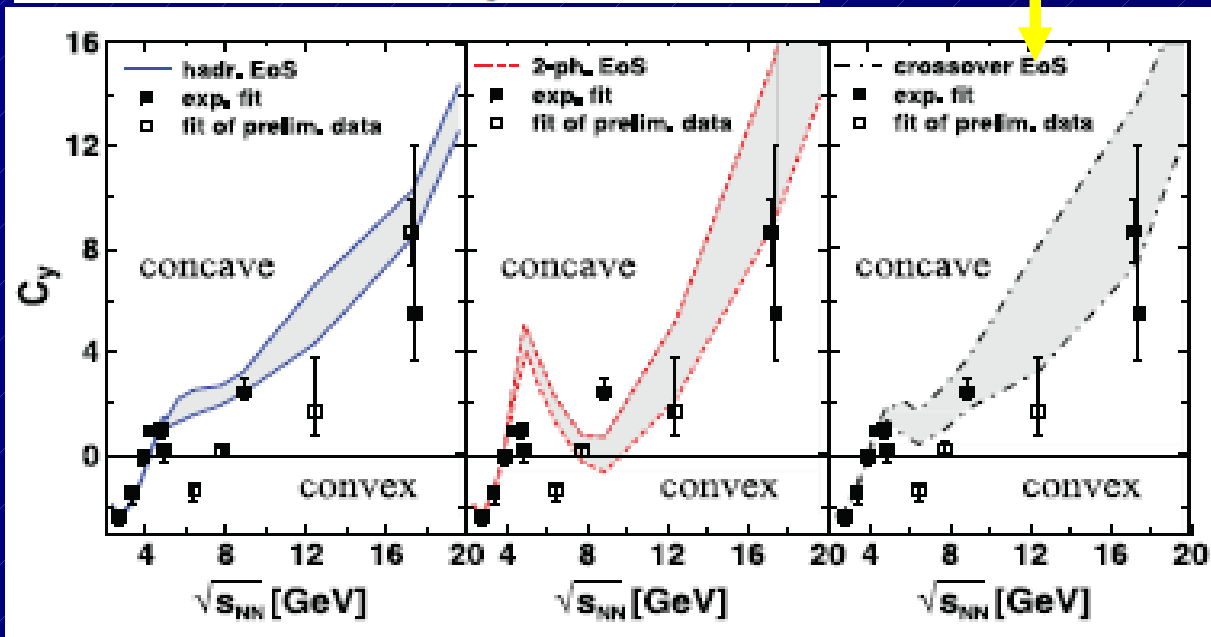
# Net proton rapidity distribution – test case for a 1<sup>st</sup> order PT signal

**Theory:** Yu.B. Ivanov, Phys. Rev. C 87, 064904 (2013)



**EoS**

3-fluid hydro  
Evolution ...

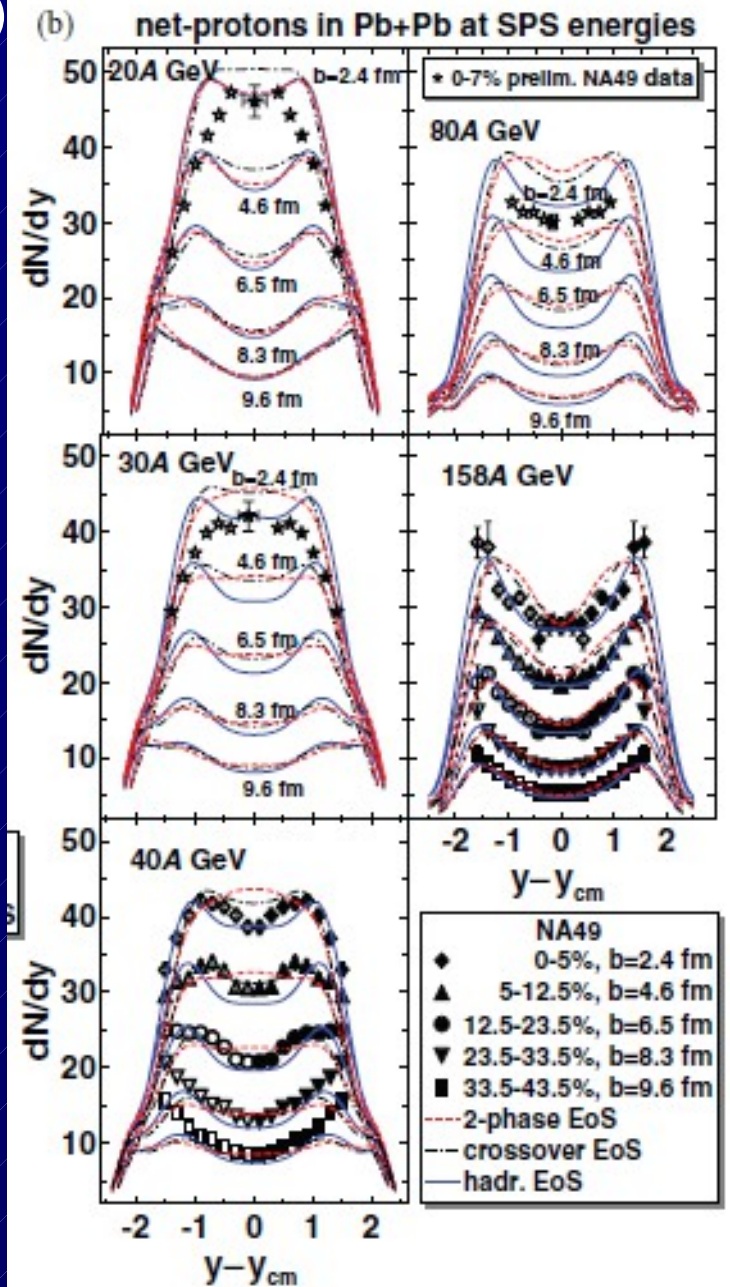
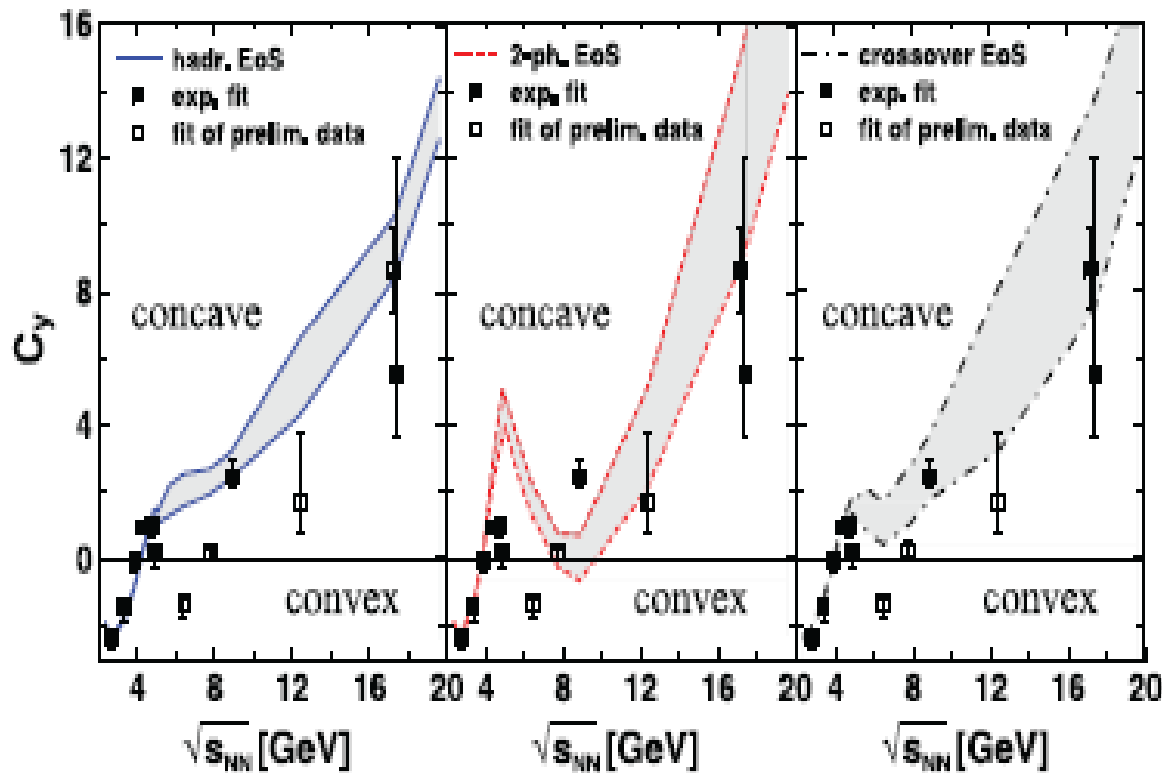


# Net proton rapidity distribution – test case for a 1<sup>st</sup> order PT signal

**Theory:** Yu.B. Ivanov, Phys. Rev. C 87, 064904 (2013)

$$C_y = \left( y_{c.m.}^3 \frac{d^3 N}{dy^3} \right)_{y=y_{c.m.}} / \left( y_{c.m.} \frac{dN}{dy} \right)_{y=y_{c.m.}}$$

$$= (y_{c.m.}/w_s)^2 (\sinh^2 y_s - w_s \cosh y_s).$$



# Net proton rapidity distribution – test case for a 1<sup>st</sup> order PT signal

## Event set:

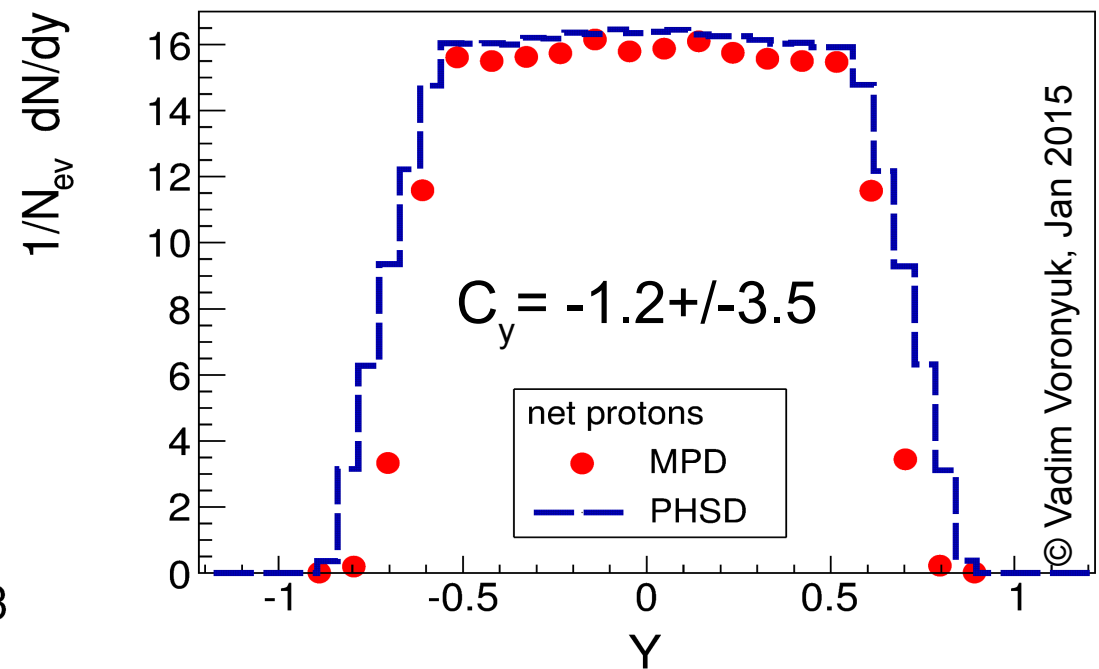
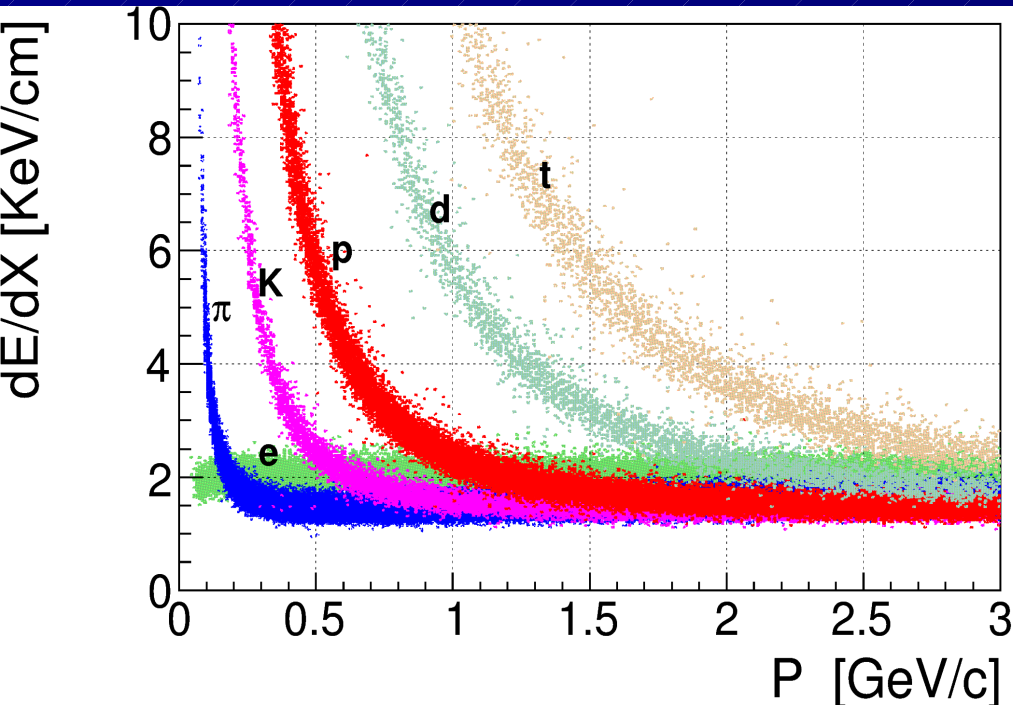
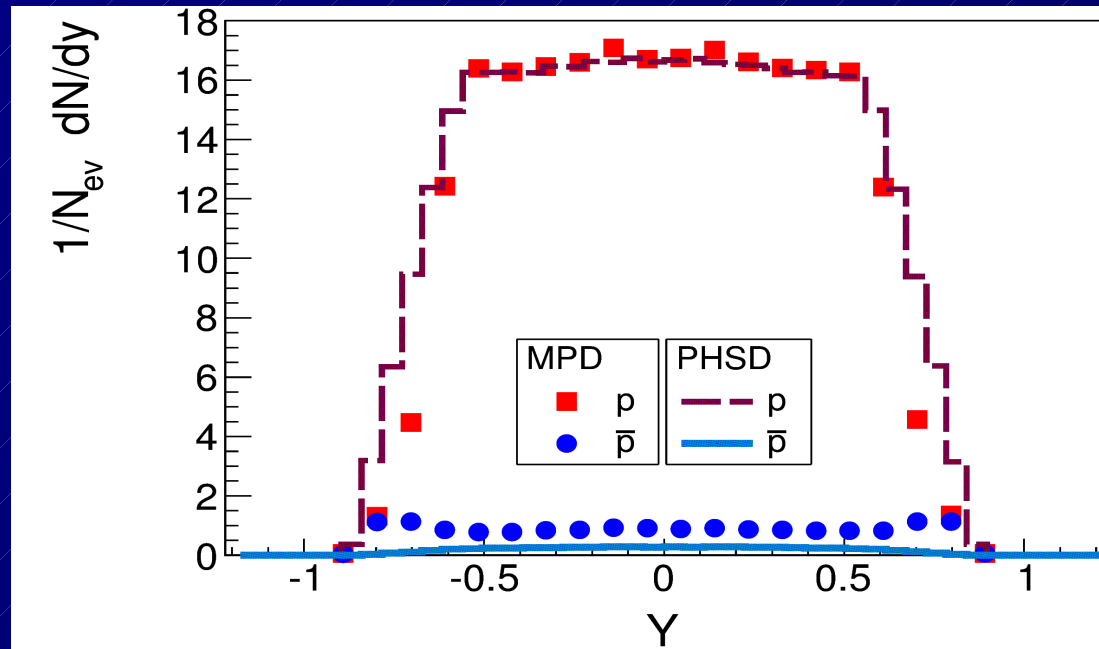
40k AuAu @  $\sqrt{s}_{NN} = 9$  GeV [0-5%]  
 The most reliable region  
 $|\eta| < 1.2$  ;  $0.4 < p_t [\text{GeV}/c] < 0.8$

## Result:

PHSD input  $\rightarrow$  GEANT+MPD  
 detector reproduces the rapidity distribution !  
 (previous concerns not confirmed !!)

## Signal:

$$C_y = \left( y_{cm}^3 \frac{d^3 N}{dy^3} \right)_{y=0} / \left( y_{cm} \frac{dN}{dy} \right)_{y=0}$$





# Net proton rapidity distribution – test case for a 1<sup>st</sup> order PT signal

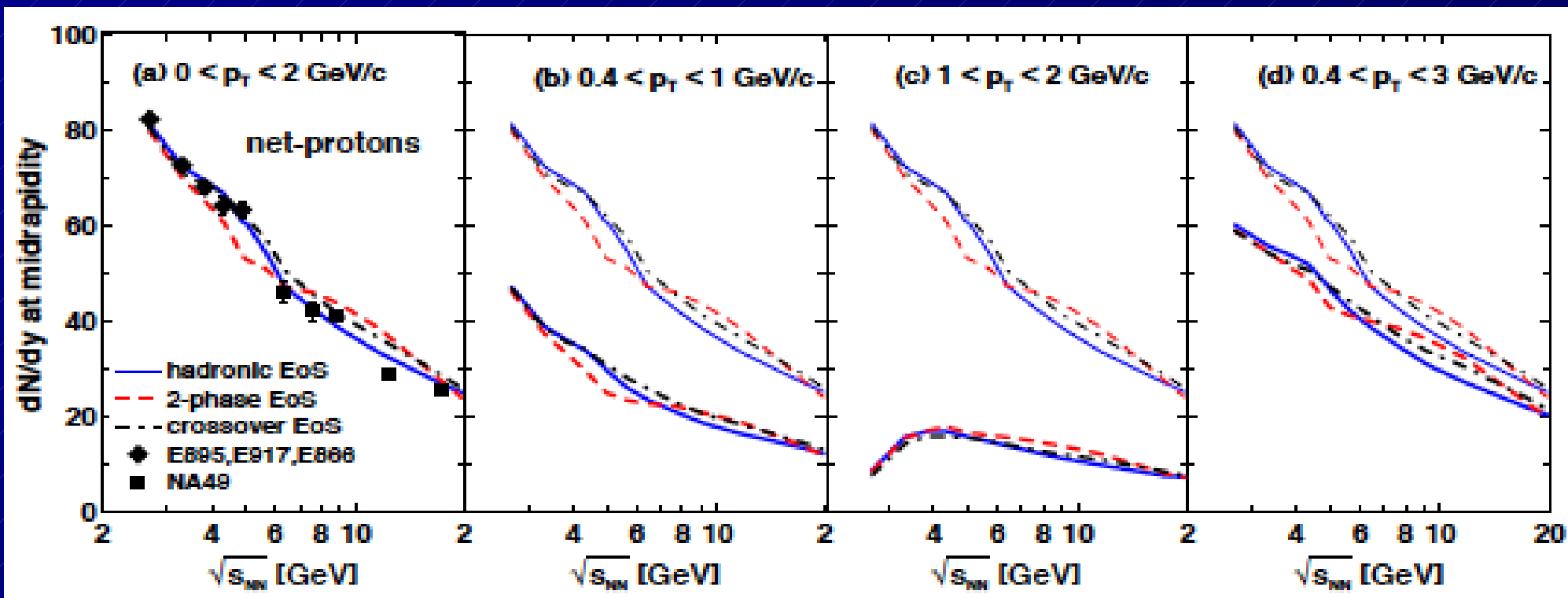
## Investigation of $p_T$ cuts:

Yu. Ivanov & D. Blaschke, arxiv:1504.03992

$$C_y = \left( y_{\text{beam}}^3 \frac{d^3 N}{dy^3} \right)_{y=0} / \left( y_{\text{beam}} \frac{dN}{dy} \right)_{y=0}$$

$$= (y_{\text{beam}}/w_s)^2 (\sinh^2 y_s - w_s \cosh y_s).$$

- i.  $0 < p_T < 2 \text{ GeV}/c$  and a very unrestrictive constraint to the rapidity range  $|y| < 0.7 y_{\text{beam}}$ , where  $y_{\text{beam}}$  is the beam rapidity in the collider mode, which is practically equivalent to the full acceptance;
- ii.  $0.4 < p_T < 1 \text{ GeV}/c$  and  $|y| < 0.5$ , the expected MPD acceptance [17];
- iii.  $1 < p_T < 2 \text{ GeV}/c$  and  $|y| < 0.5$ , an acceptance range where low-momentum particles witnessing collective behaviour are largely eliminated;
- iv.  $0.4 < p_T < 3 \text{ GeV}/c$  and  $|y| < 0.5$ , the range of the STAR acceptance [18].



# Net proton rapidity distribution – test case for a 1<sup>st</sup> order PT signal

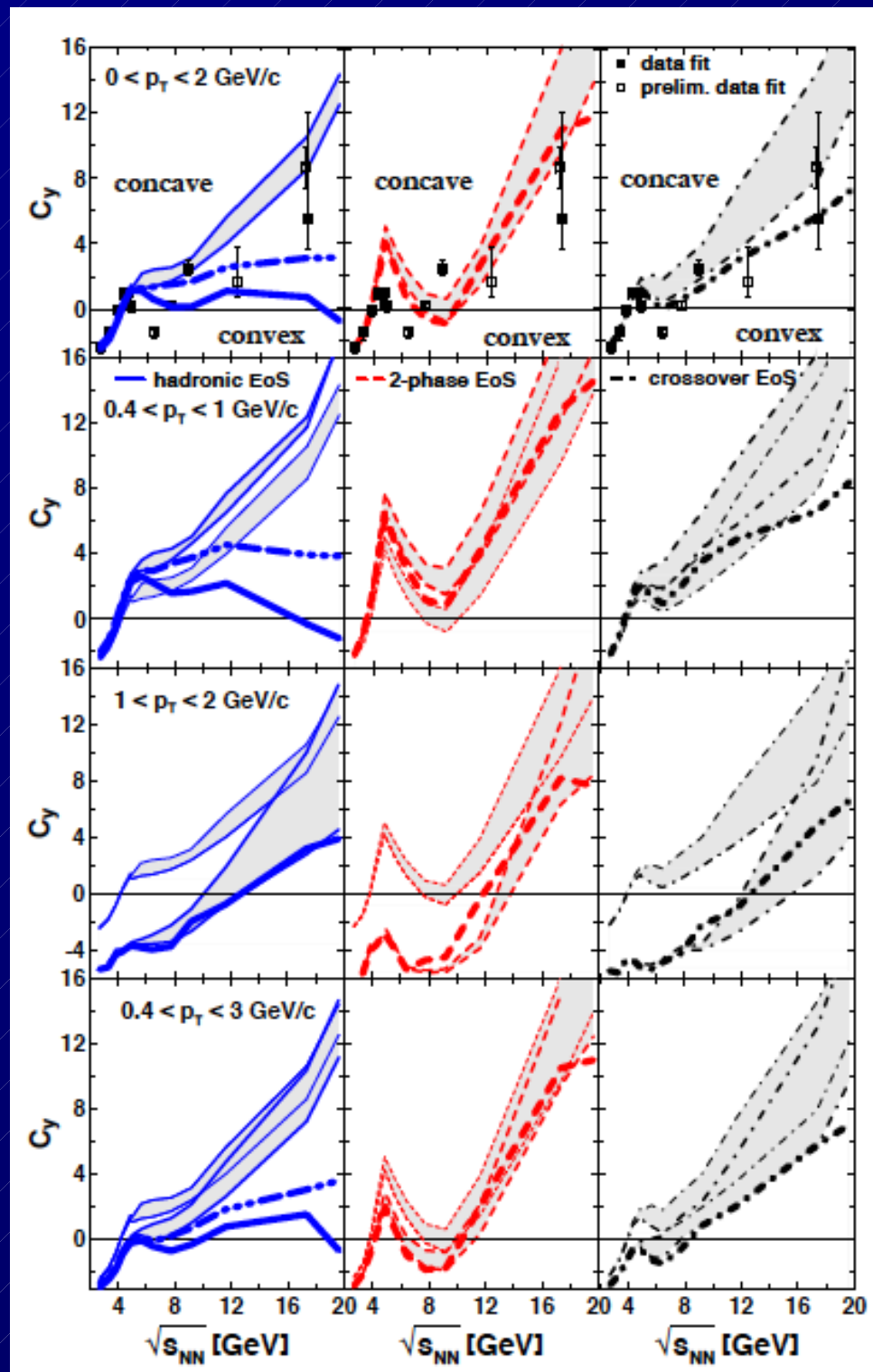
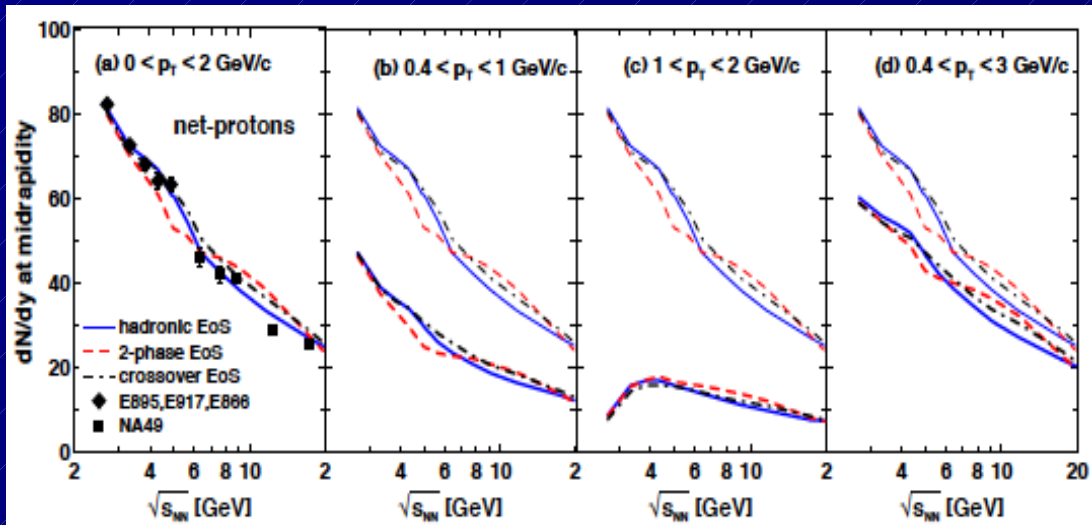
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Yu. Ivanov & D. B., PRC 92, 024916 (2015)

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$$= (y_{\text{beam}}/w_s)^2 (\sinh^2 y_s - w_s \cosh y_s).$$

- “wiggle” formed in the nonequilibrium compression stage of the collision, where  $p_T$  **only in 3FH**
- robust against serious  $p_T$  cuts
- at high  $p_T$  (1 - 2 GeV/c) in convex region
- at low  $p_T$  (0.2 - 1 GeV/c) in concave region
- required accuracy in  $C_y$  determination:  $\Delta C_y < 2$



# Net proton rapidity distribution – test case for a 1<sup>st</sup> order PT signal

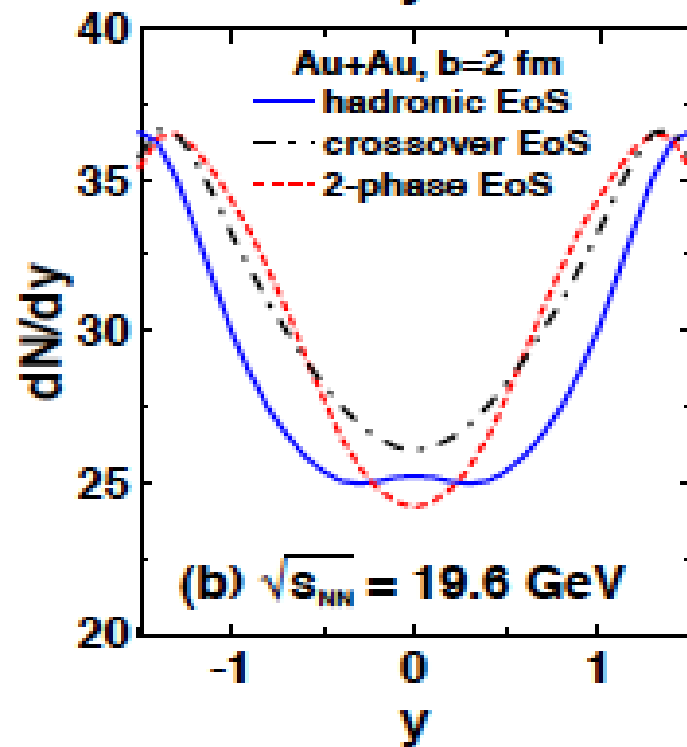
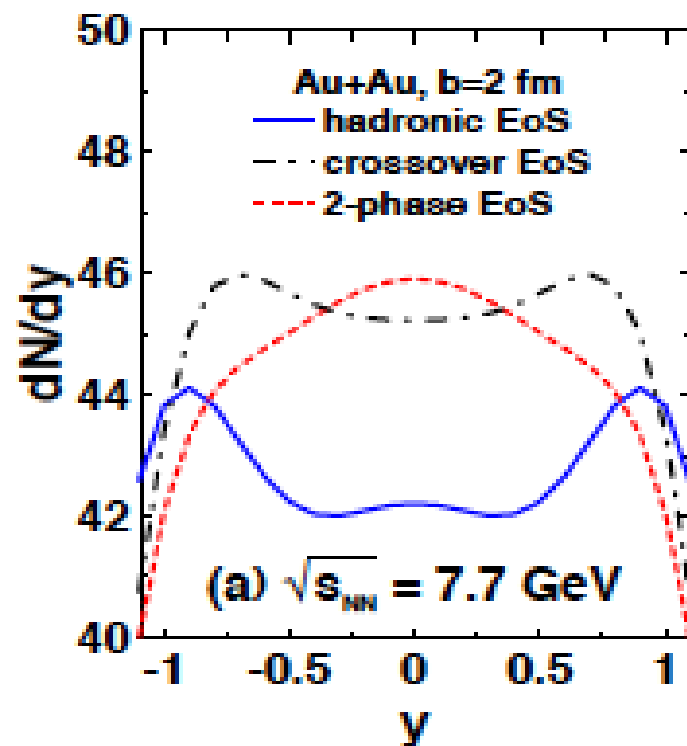
## Investigation of $p_T$ cuts:

Yu. Ivanov & D. Blaschke, arxiv:1504.03992

$$\frac{dN}{dy} = a \left( \exp \left\{ -\left(1/w_s\right) \cosh(y - y_s) \right\} + \exp \left\{ -\left(1/w_s\right) \cosh(y + y_s) \right\} \right),$$

$$C_y = \left( y_{\text{beam}}^3 \frac{d^3 N}{dy^3} \right)_{y=0} / \left( y_{\text{beam}} \frac{dN}{dy} \right)_{y=0} = (y_{\text{beam}}/w_s)^2 (\sinh^2 y_s - w_s \cosh y_s).$$

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# 3+1D viscous hydro-cascade model (Yu. Karpenko, FIAS)

3+1D viscous hydro+cascade model was applied for A+A collisions at RHIC Beam Energy Scan energies ( $\sqrt{s} = 7.7 - 39$  GeV), and for SPS energy points

Cascade-hydro-cascade approach:

Initial state: UrQMD cascade

S.A. Bass et al., Prog. Part. Nucl. Phys. 41 255-369, 1998

Hydrodynamic phase: numerical 3+1D hydro solution via original relativistic viscous hydro code

Iu. Karpenko, P. Huovinen, M. Bleicher, arXiv:1312.4160

Hydro starts at  $\tau = \sqrt{t^2 - z^2} = \tau_0$  (red curve):

$$\tau_0 = \frac{2R}{\gamma v_z}$$

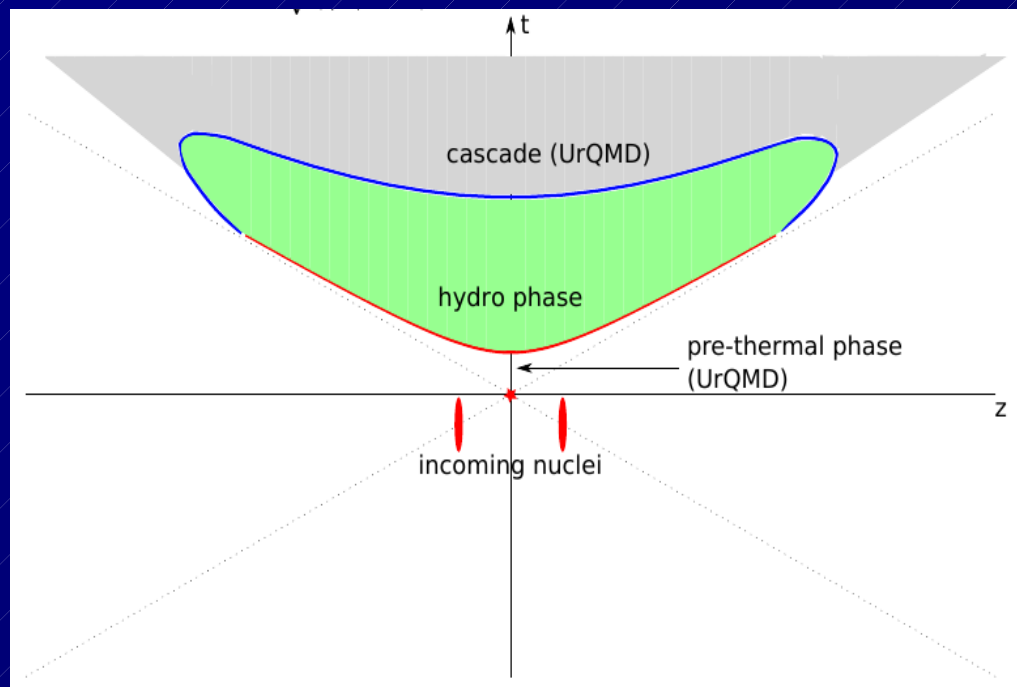
$\{T^{0\mu}, N_b^0, N_q^0\}$  of fluid = averaged  $\{T^{0\mu}, N_b^0, N_q^0\}$  of particles

Fluid  $\rightarrow$  particle transition

$\varepsilon = \varepsilon_{SW} = 0.5 \text{ GeV/fm}^3$  (blue curve):

$\{T^{0\mu}, N_b^0, N_q^0\}$  of hadron-resonance gas =  $\{T^{0\mu}, N_b^0, N_q^0\}$  of fluid

Hadronic cascade: UrQMD



## Equations of state for hydrodynamic phase

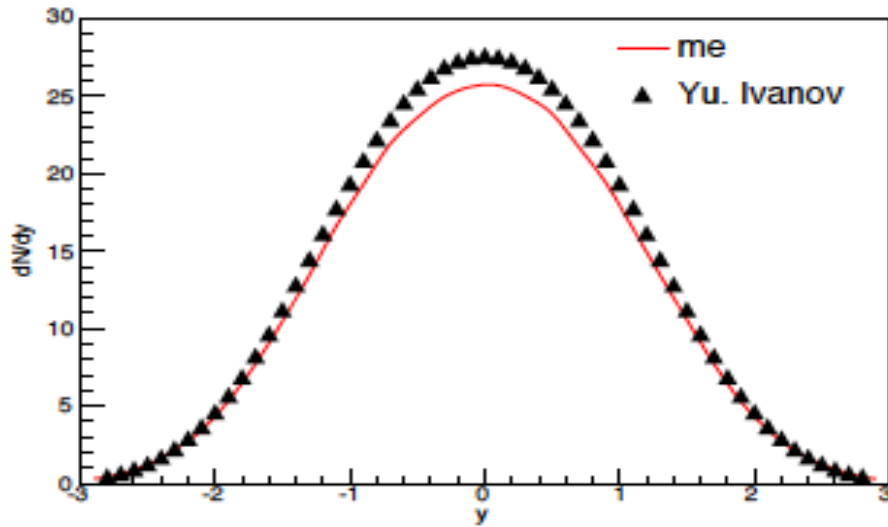
- Chiral model
  - ▶ coupled to Polyakov loop to include the deconfinement phase transition
  - ▶ good agreement with lattice QCD data at  $\mu_B = 0$ , also applicable at finite baryon densities
  - ▶ (current version) has **crossover type PT** between hadron and quark-gluon phase at all  $\mu_B$
- Hadron resonance gas + Bag Model (a.k.a. EoS Q)
  - ▶ hadron resonance gas made of  $u, d$  quarks including repulsive meanfield
  - ▶ the phases matched via Maxwell construction, resulting in **1<sup>st</sup> order PT**

# Preview: Particlization of 3-fluid Hydrodynamics model

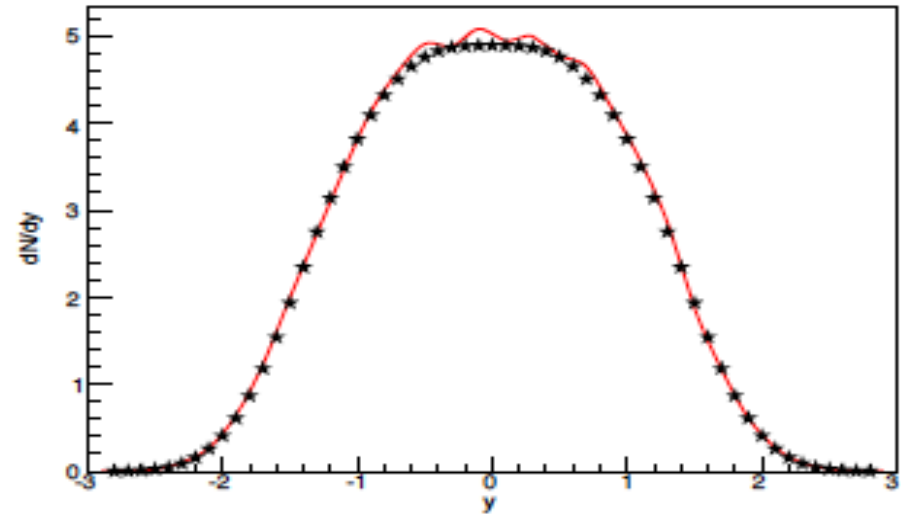
Yu. Karpenko & Yu. Ivanov: May 2014 - May 2015

Rapidity distribution

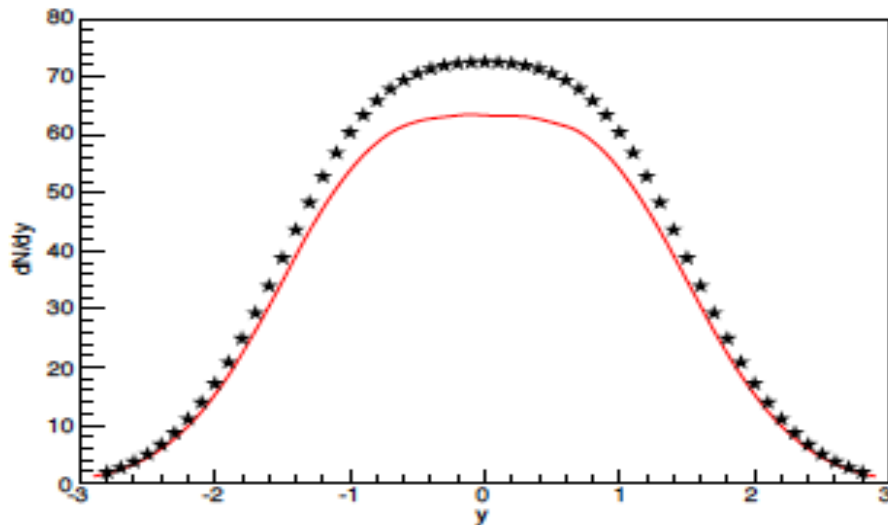
pions, fireball



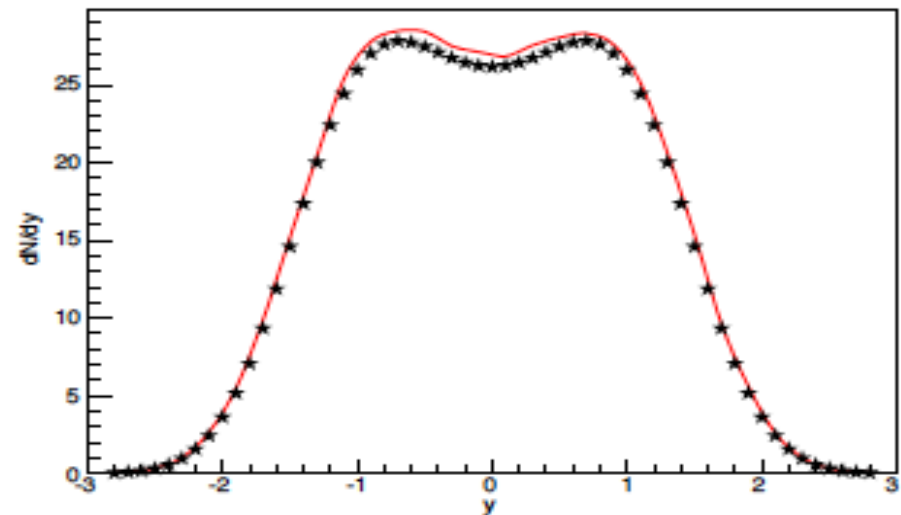
kaons, fireball



pions, baryon-rich



kaons, baryon-rich

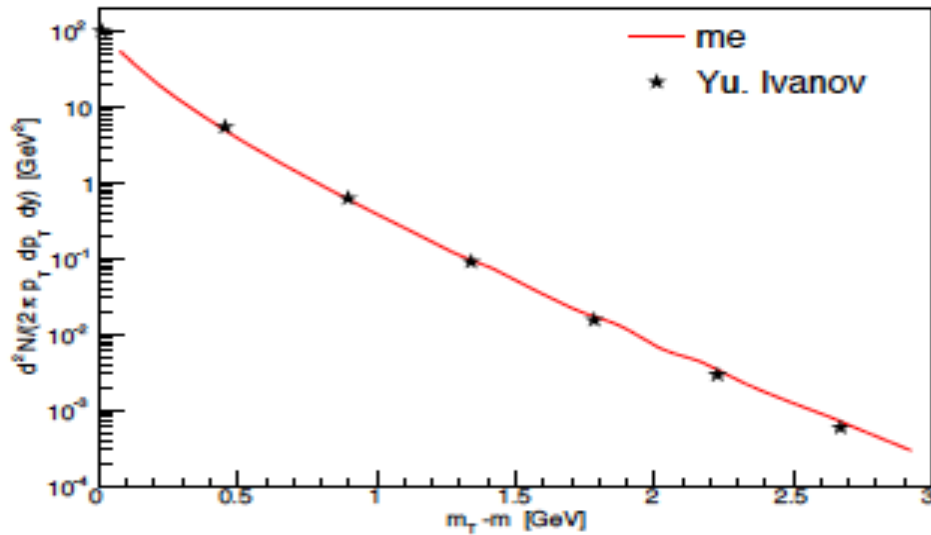


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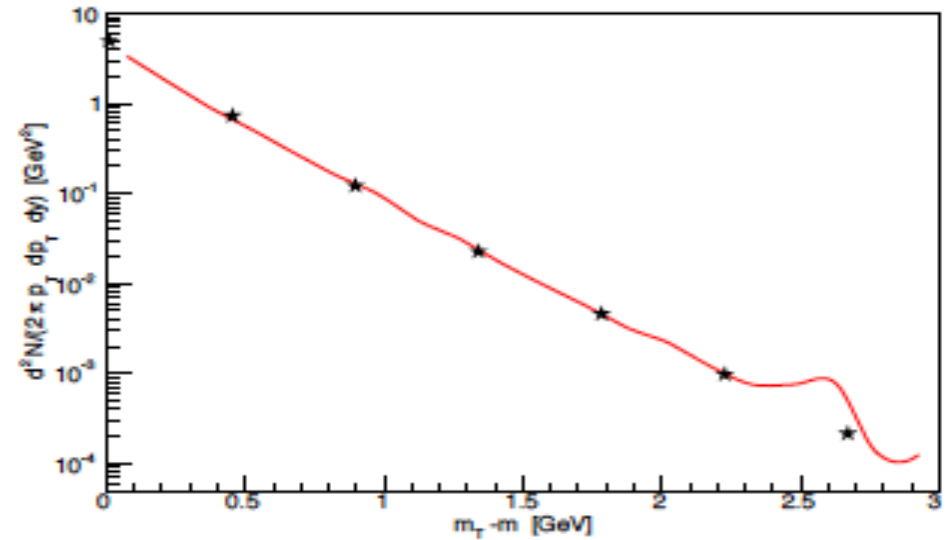
Yu. Karpenko & Yu. Ivanov: May 2014 - May 2015

$P_T / m_T$  distribution

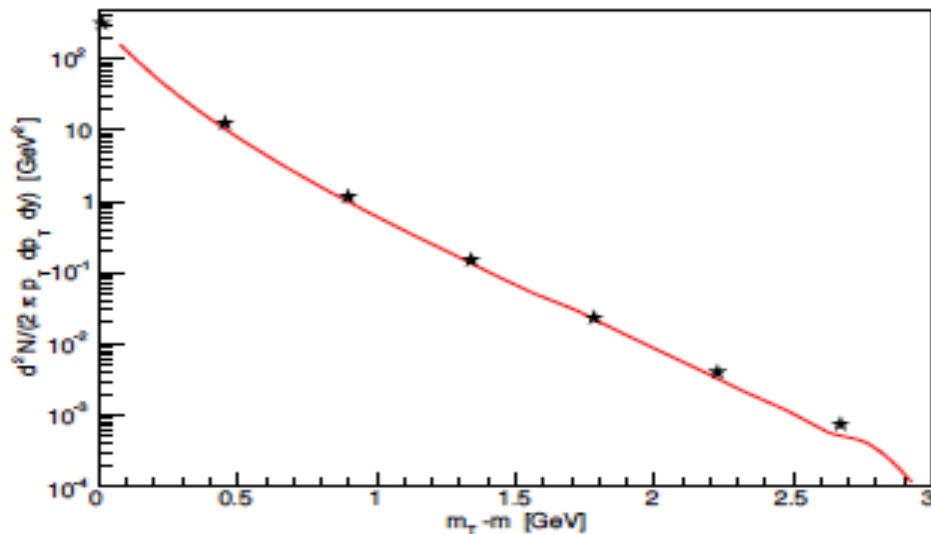
pions, fireball



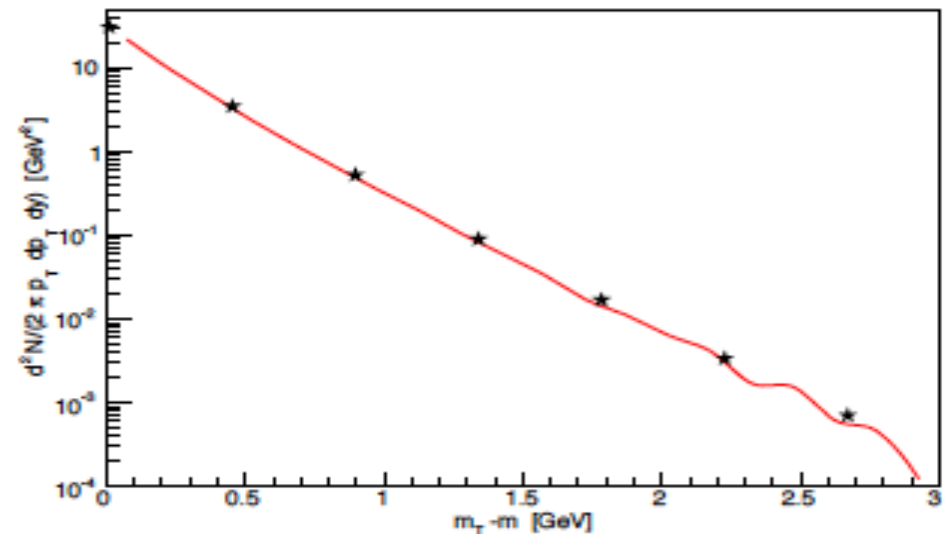
kaons, fireball



pions, baryon-rich



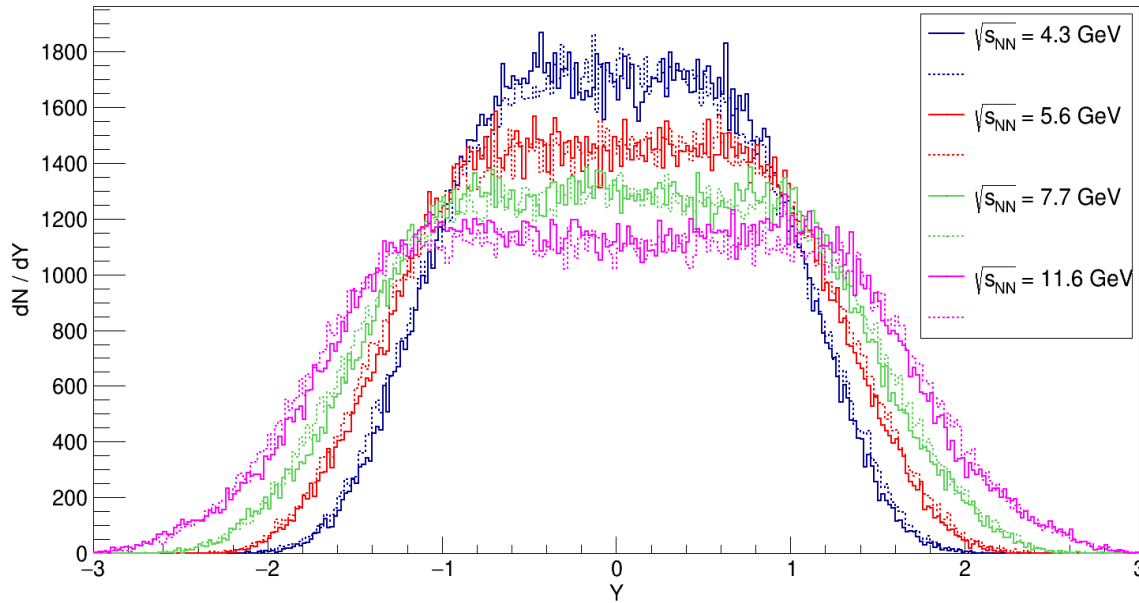
kaons, baryon-rich



**Baryon stopping signal for first order phase transition ?**

# Baryon stopping signal for first order phase transition ?

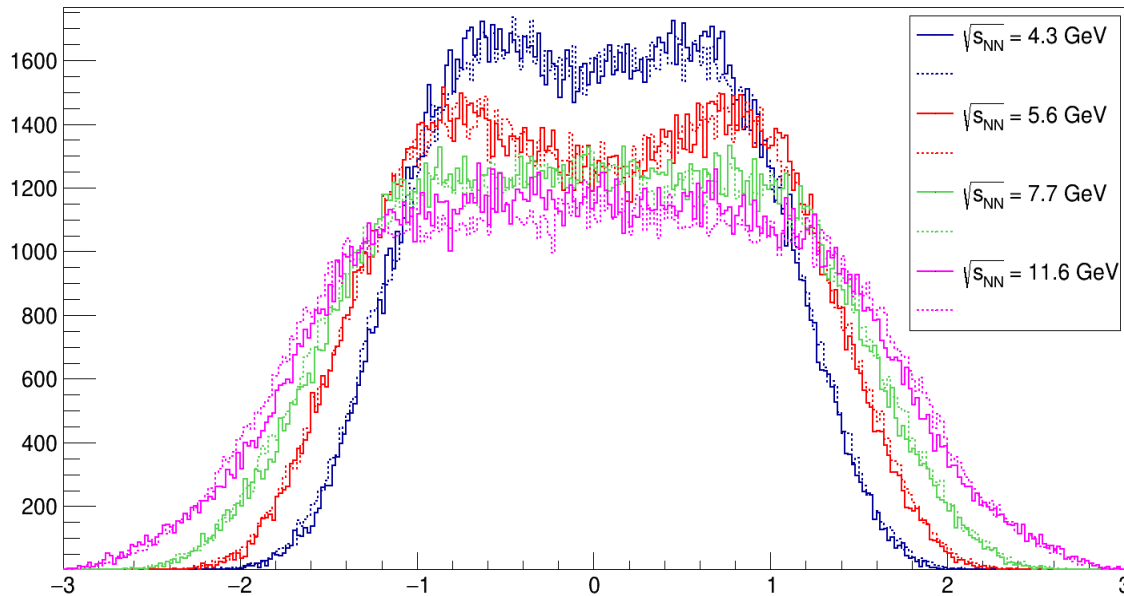
i1, EoS - crossover, solid line - 3FD, dashed line - 3FD + UrQMD



EoS: Crossover

Impact parameter:  $b=2$  fm

i1, EoS - 1PT, solid line - 3FD, dashed line = 3FD + UrQMD

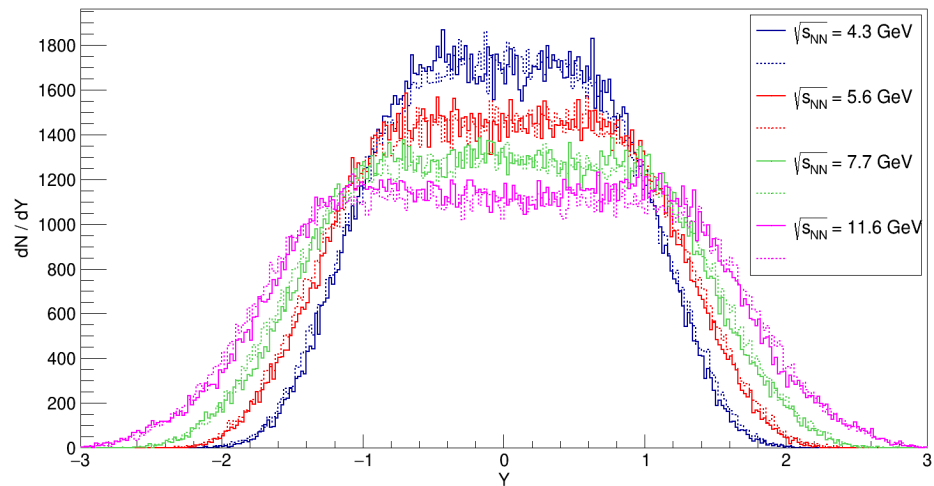


EoS: First order PT

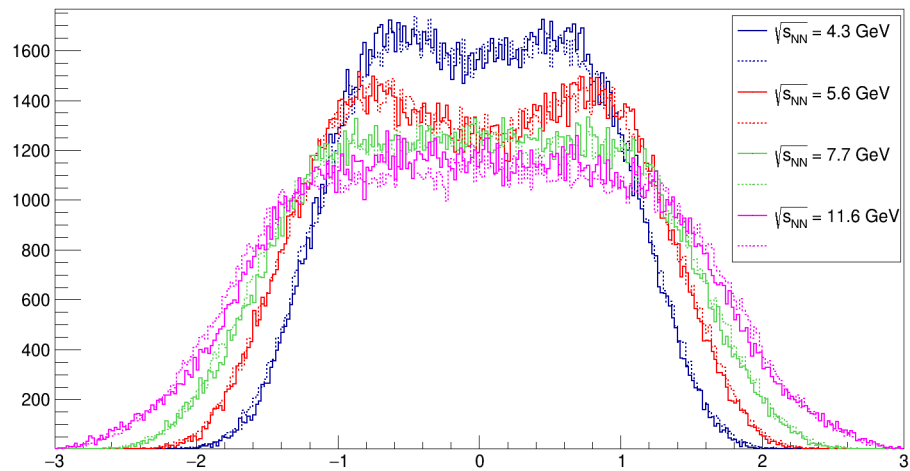
Impact parameter:  $b=2$  fm



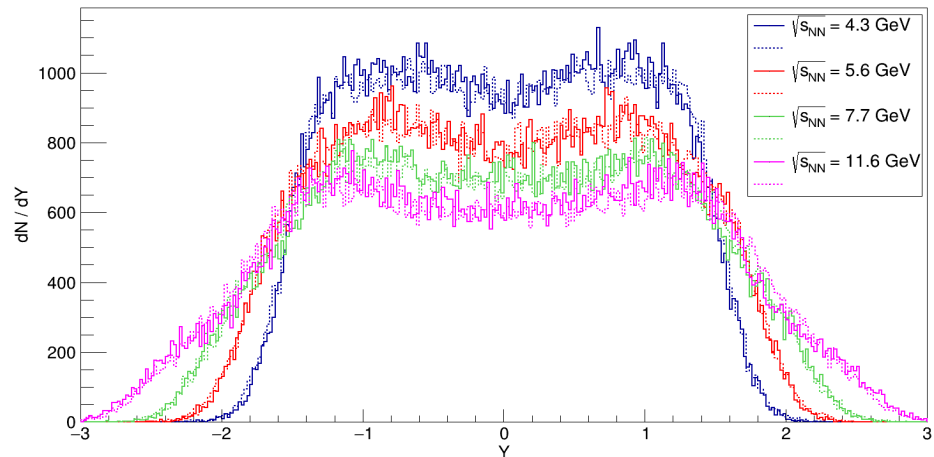
i1, EoS - crossover, solid line - 3FD, dashed line - 3FD + UrQMD



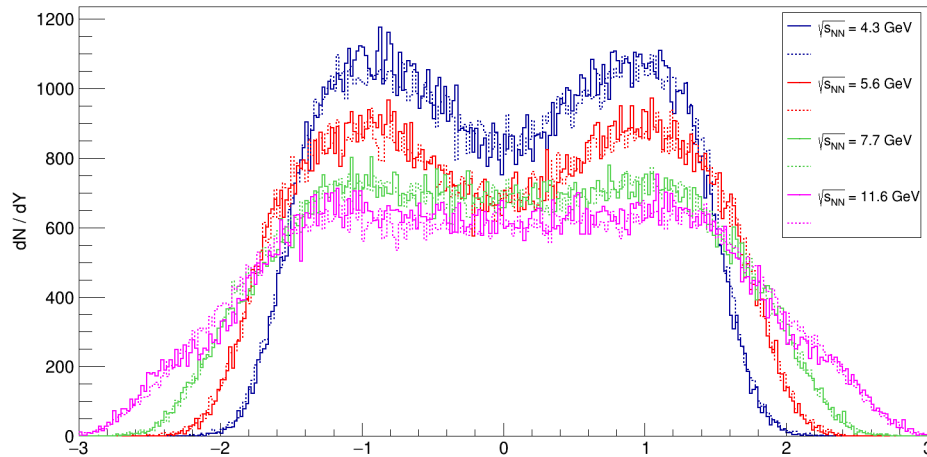
i1, EoS - 1PT, solid line - 3FD, dashed line = 3FD + UrQMD



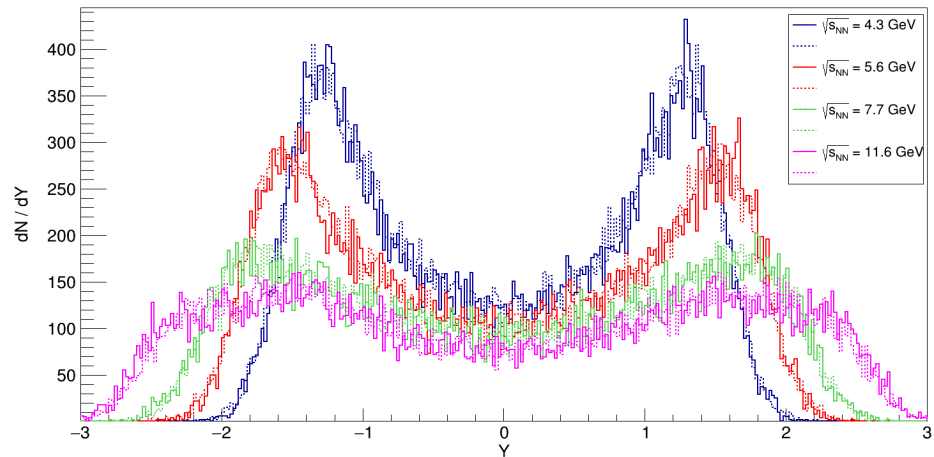
i3, EoS - crossover, solid line - 3FD, dashed line - 3FD + UrQMD



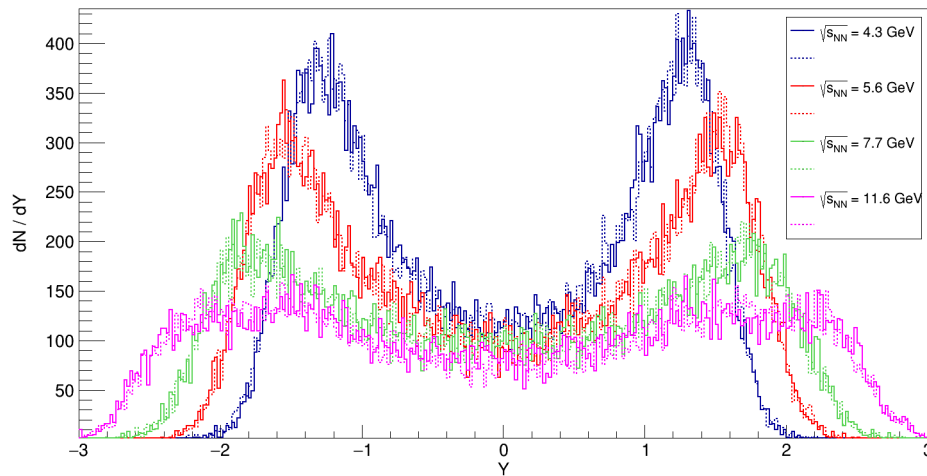
i3, EoS- 1PT, solid line - 3FD, dashed line - 3FD + UrQMD



i5, EoS - crossover, solid line - 3FD, dashed line - 3FD + UrQMD



i5, EoS - 1PT, solid line - 3FD, dashed line - 3FD + UrQMD



## **Further developments:**

- Hadron rescattering in UrQMD (Hannah Petersen)**
- MPD Detector simulation (Oleg Rogachevsky et al.)**
- New 2-phase EoS (David Blaschke et al.)**

**A new class of 2-phase EoS:  
Motivation from Astrophysics**

# Modern 2-phase EoS in Astrophysics of Compact Stars

## 1. Pauli blocking effect → Excluded volume

Well known from modeling dissociation of clusters in the supernova EoS:

- excluded volume: Lattimer-Swesty (1991), Shen-Toki-Oyematsu-Sumiyoshi (1996), ...
- Pauli blocking: Roepke-Grigo-Sumiyoshi-Shen (2003), Typel et al. PRC 81 (2010)
- excl. Vol. vs. Pauli blocking: Hempel, Schaffner-Bielich, Typel, Roepke PRC 84 (2011)

**Here:** nucleons as quark clusters with finite size --> excluded volume effect !

Available volume fraction:  $\Phi = V_{\text{av}}/V = 1 - v \sum_{i=n,p} n_i$ ,  $v = \frac{1}{2} \frac{4\pi}{3} (2r_{\text{nuc}})^3 = 4V_{\text{nuc}}$

Equations of state for T=0 nuclear matter:  $p_{\text{tot}}(\mu_n, \mu_p) = \frac{1}{\Phi} \sum_{i=n,p} p_i + p_{\text{mes}}$ ,

$$p_i = \frac{1}{4} (E_i n_i - m_i^* n_i^{(s)}),$$

$$\varepsilon_{\text{tot}}(\mu_n, \mu_p) = -p_{\text{tot}} + \sum_{i=n,p} \mu_i n_i,$$

$$n_i = \frac{\Phi}{3\pi^3} k_i^3,$$

$$n_i^{(s)} = \frac{\Phi m_i^*}{2\pi^2} \left[ E_i k_i - (m_i^*)^2 \ln \frac{k_i + E_i}{m_i^*} \right],$$

Effective mass:  $m_i^* = m_i - S_i$ .

Scalar meanfield:  $S_i \sim n_i^{(s)}$

$$E_i = \sqrt{k_i^2 + (m_i^*)^2} = \mu_i - V_i - \frac{v}{\Phi} \sum_{j=p,n} p_j,$$

Vector meanfield:  $V_i \sim n_i$

# Modern 2-phase EoS in Astrophysics of Compact Stars

## 2. Stiff quark matter at high densities

S. Benic, Eur. Phys. J. A 50, 111 (2014)

$$\mathcal{L} = \bar{q}(i\cancel{\partial} - m)q + \mu_q \bar{q}\gamma^0 q + \mathcal{L}_4 + \mathcal{L}_8, \quad \mathcal{L}_4 = \frac{g_{20}}{\Lambda^2} [(\bar{q}q)^2 + (\bar{q}i\gamma_5\tau q)^2] - \frac{g_{02}}{\Lambda^2} (\bar{q}\gamma_\mu q)^2,$$
$$\mathcal{L}_8 = \frac{g_{40}}{\Lambda^8} [(\bar{q}q)^2 + (\bar{q}i\gamma_5\tau q)^2]^2 - \frac{g_{04}}{\Lambda^8} (\bar{q}\gamma_\mu q)^4 - \frac{g_{22}}{\Lambda^8} (\bar{q}\gamma_\mu q)^2 [(\bar{q}q)^2 + (\bar{q}i\gamma_5\tau q)^2]$$

**Meanfield approximation:**  $\mathcal{L}_{\text{MF}} = \bar{q}(i\cancel{\partial} - M)q + \tilde{\mu}_q \bar{q}\gamma^0 q - U,$

$$M = m + 2\frac{g_{20}}{\Lambda^2} \langle \bar{q}q \rangle + 4\frac{g_{40}}{\Lambda^8} \langle \bar{q}q \rangle^3 - 2\frac{g_{22}}{\Lambda^8} \langle \bar{q}q \rangle \langle q^\dagger q \rangle^2,$$

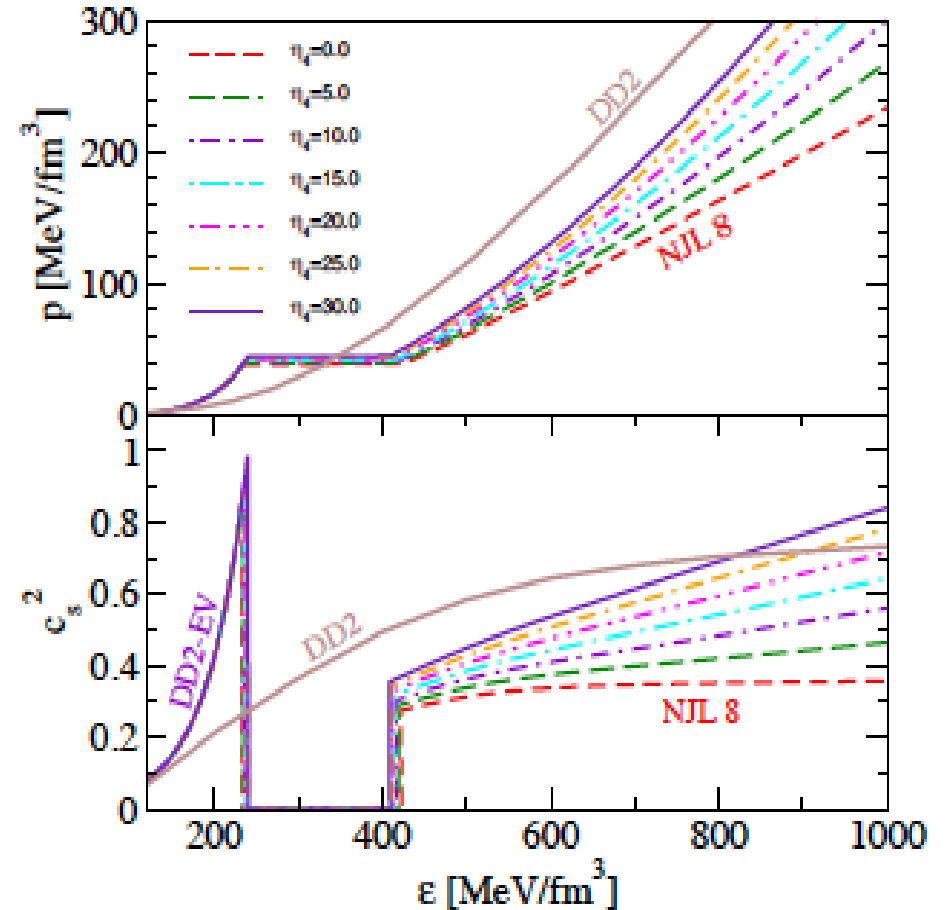
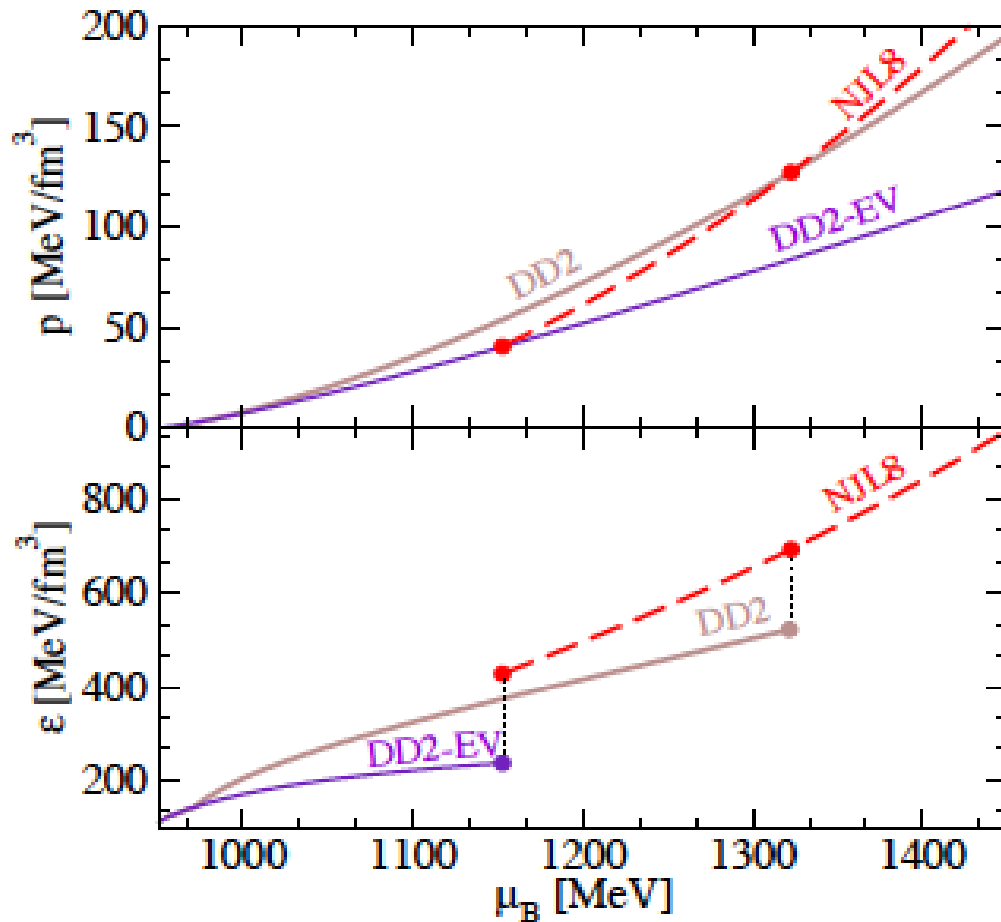
$$\tilde{\mu}_q = \mu_q - 2\frac{g_{02}}{\Lambda^2} \langle q^\dagger q \rangle - 4\frac{g_{04}}{\Lambda^8} \langle q^\dagger q \rangle^3 - 2\frac{g_{22}}{\Lambda^8} \langle \bar{q}q \rangle^2 \langle q^\dagger q \rangle,$$

$$U = \frac{g_{20}}{\Lambda^2} \langle \bar{q}q \rangle^2 + 3\frac{g_{40}}{\Lambda^8} \langle \bar{q}q \rangle^4 - 3\frac{g_{22}}{\Lambda^8} \langle \bar{q}q \rangle^2 \langle q^\dagger q \rangle^2 - \frac{g_{02}}{\Lambda^2} \langle q^\dagger q \rangle^2 - 3\frac{g_{04}}{\Lambda^8} \langle q^\dagger q \rangle^4.$$

**Thermodynamic Potential:**

$$\Omega = U - 2N_f N_c \int \frac{d^3 p}{(2\pi)^3} \left\{ E + T \log[1 + e^{-\beta(E - \tilde{\mu}_q)}] + T \log[1 + e^{-\beta(E + \tilde{\mu}_q)}] \right\} + \Omega_0$$

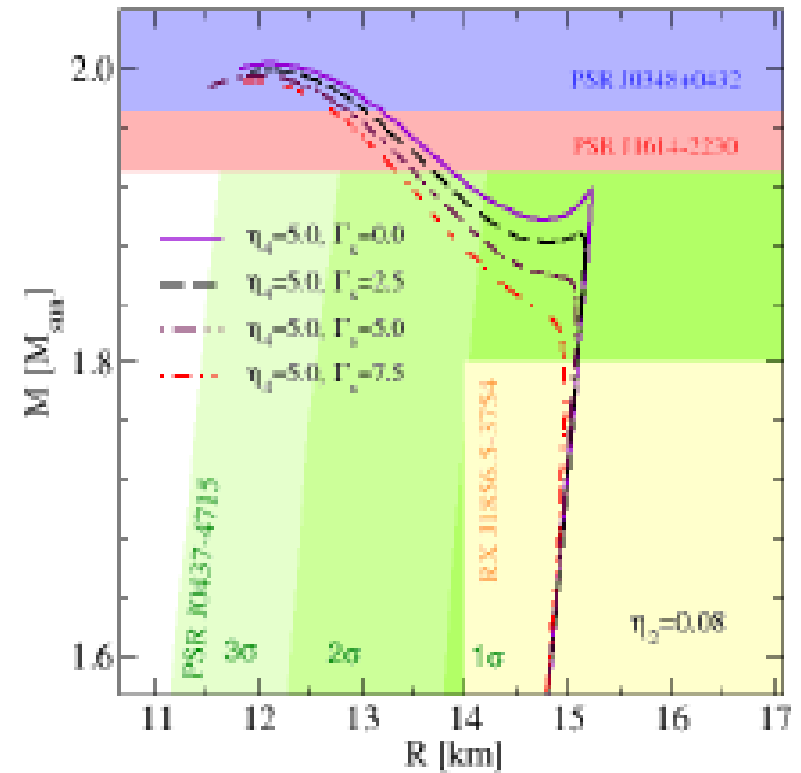
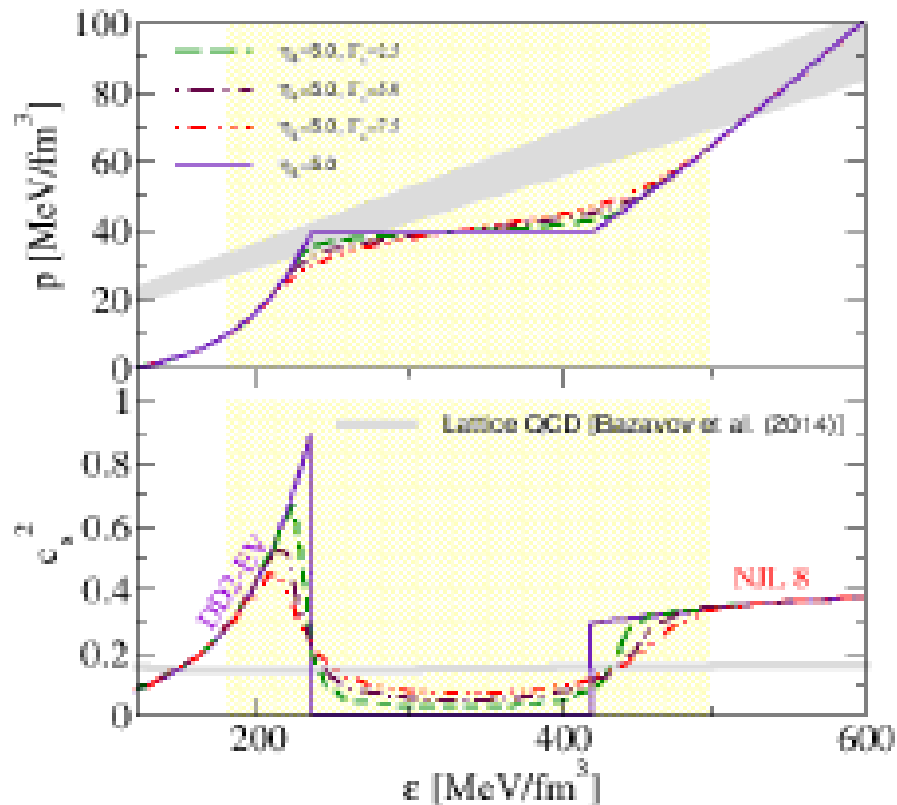
# Modern 2-phase EoS in Astrophysics of Compact Stars



Here: Stiffening of dense hadronic matter by excluded volume in density-dependent RMF  
 Stiffening of dense quark matter by higher order quark vector current interactions ( $\eta_4$ )

# Modern 2-phase EoS in Astrophysics of Compact Stars

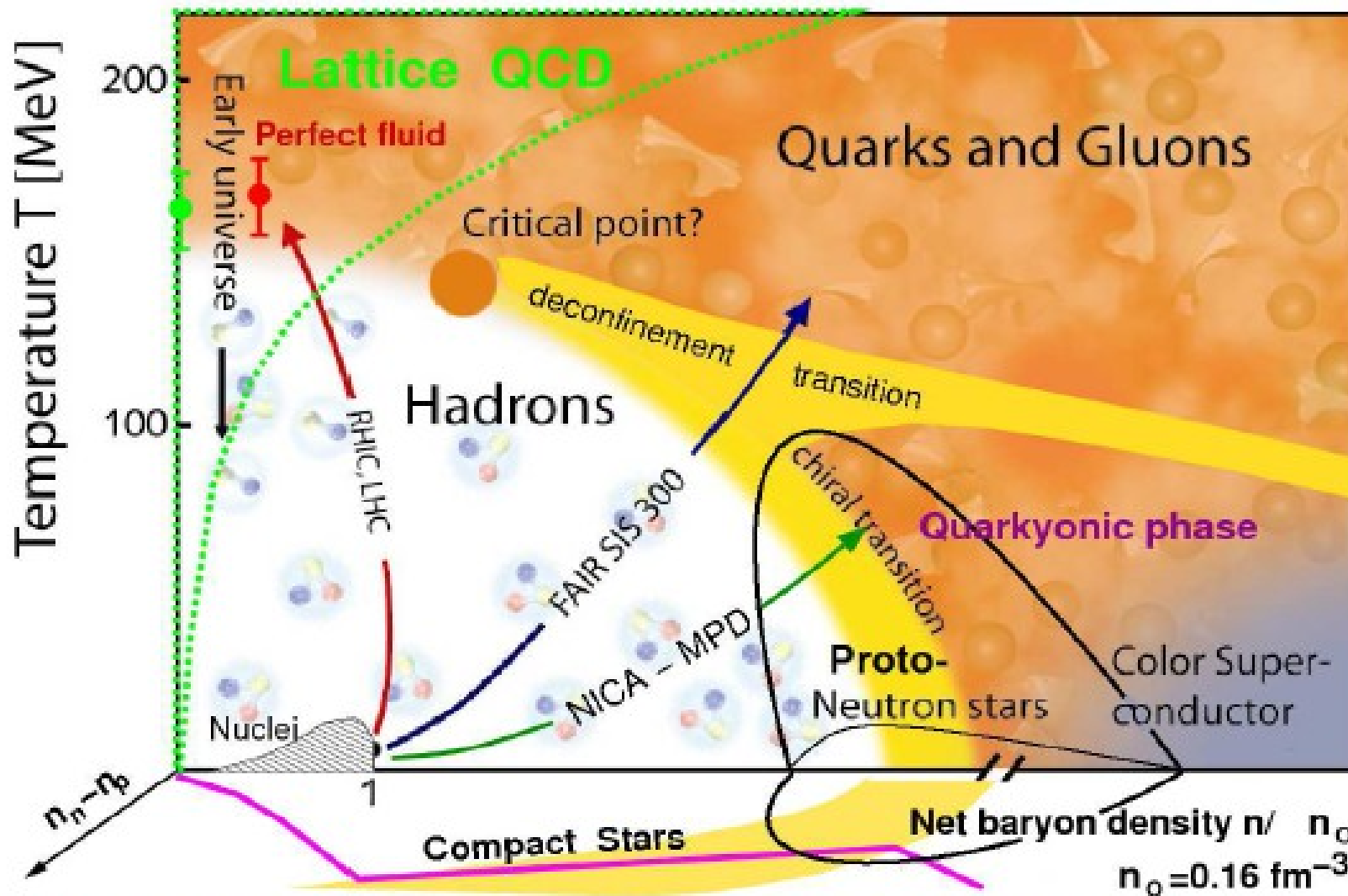
Estimate effects of structures in the phase transition region (“pasta”)



High-mass Twins relatively robust against “smoothing” the Maxwell transition construction

D. Alvarez-Castillo, D.B., [arxiv:1412.8463](https://arxiv.org/abs/1412.8463)

# Support a CEP in QCD phase diagram with Astrophysics?

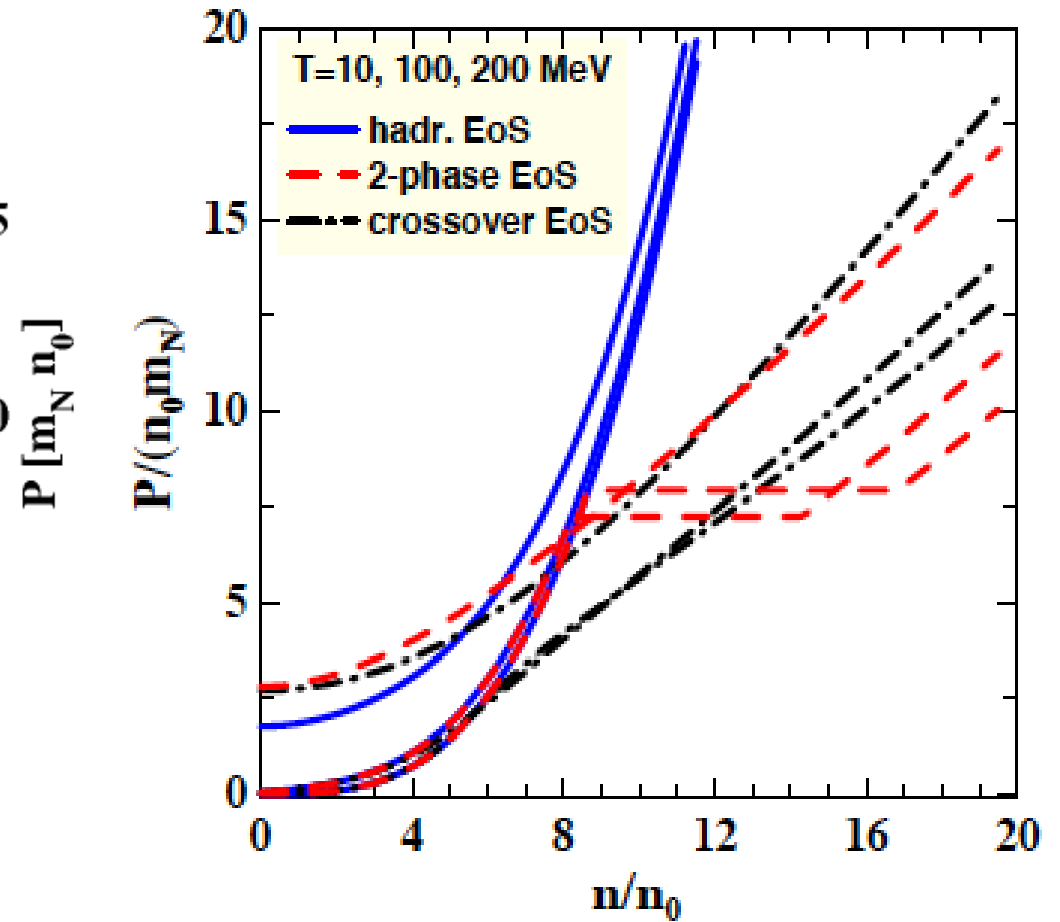
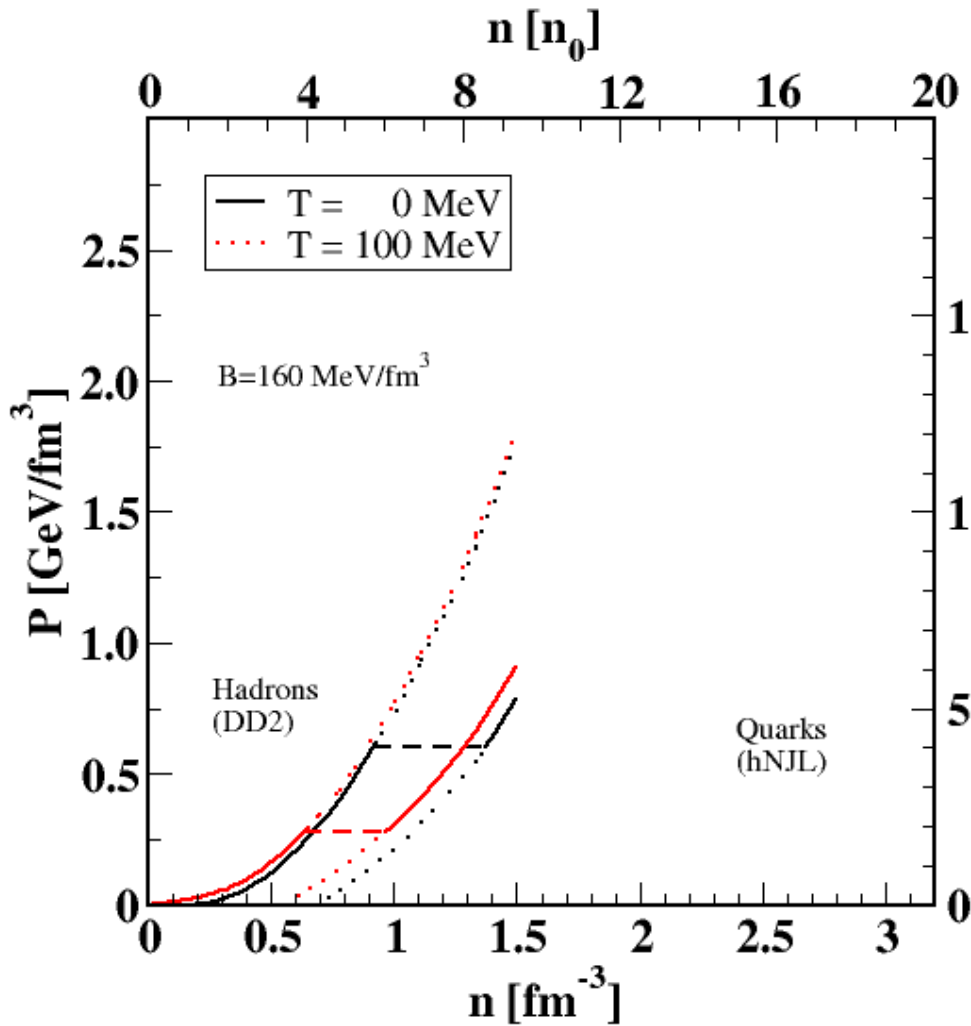


NICA White Paper, <http://theor.jinr.ru/twiki-cgi/view/NICA/WebHome>

Crossover at finite  $T$  (Lattice QCD) + First order at zero  $T$  (Astrophysics) = Critical endpoint exists!



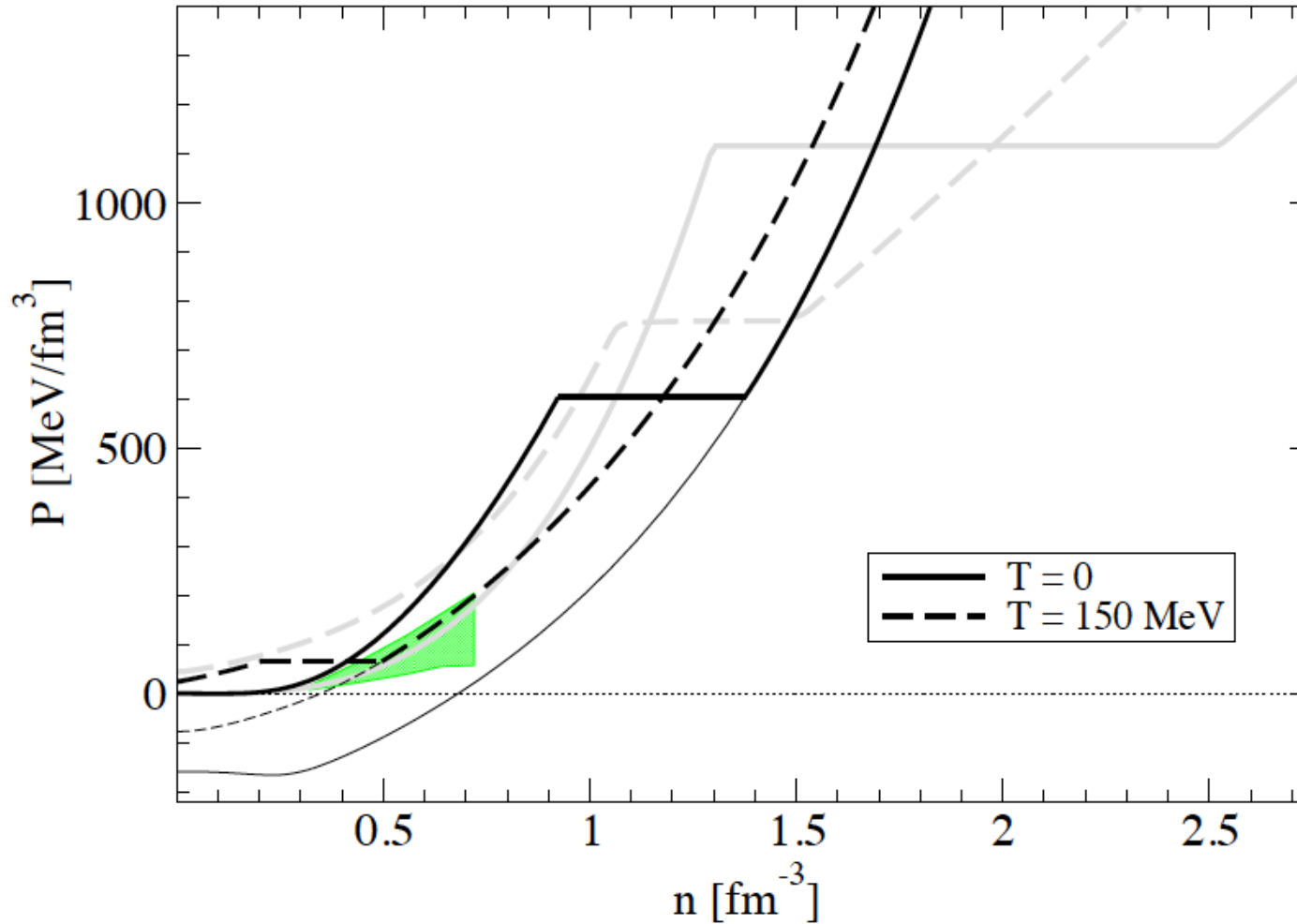
# Comparison 2-phase EoS



N.-U. Bastian, D. Blaschke (S. Benic, S. Typel),  
 In progress (2015)

A. Khvorostukhin et al. EPJC 48 (2006) 531  
 Yu. Ivanov, D. Blaschke, arxiv:1504.03992

# Comparison 2-phase EoS



Grey: Ivanov (2010)

Black: DD2 vs. HNJLb

$B=160 \text{ MeV}/\text{fm}^3$

# Summary / Outlook:

- Baryon stopping signal (“wiggle”) remains a robust signal for 1<sup>st</sup> order PT also under severe cuts in transverse momentum !
- Discrimination between hadronic phase and crossover transition ambiguous
- Position of the “wiggle” in the beam energy scan is EoS dependent – new EoS ?!
- Particlization of 3-Fluid Hydrodynamics model works !
- UrQMD “afterburner” works too !
  
- Detector simulation in progress
- Systematic study of modern 2-phase EoS (Bayesian analysis) in progress

# Critical Point and Onset of Deconfinement

## CPOD 2016

University of Wrocław, Poland

May 30<sup>th</sup> – June 3<sup>rd</sup>, 2016

### Scientific Topics:

- Critical Point
- Phase transitions in hot and dense matter
- Deconfinement and chiral symmetry restoration
- Hadronization and chemical freeze-out
- Compact Stars
- Future facilities, detectors and methods

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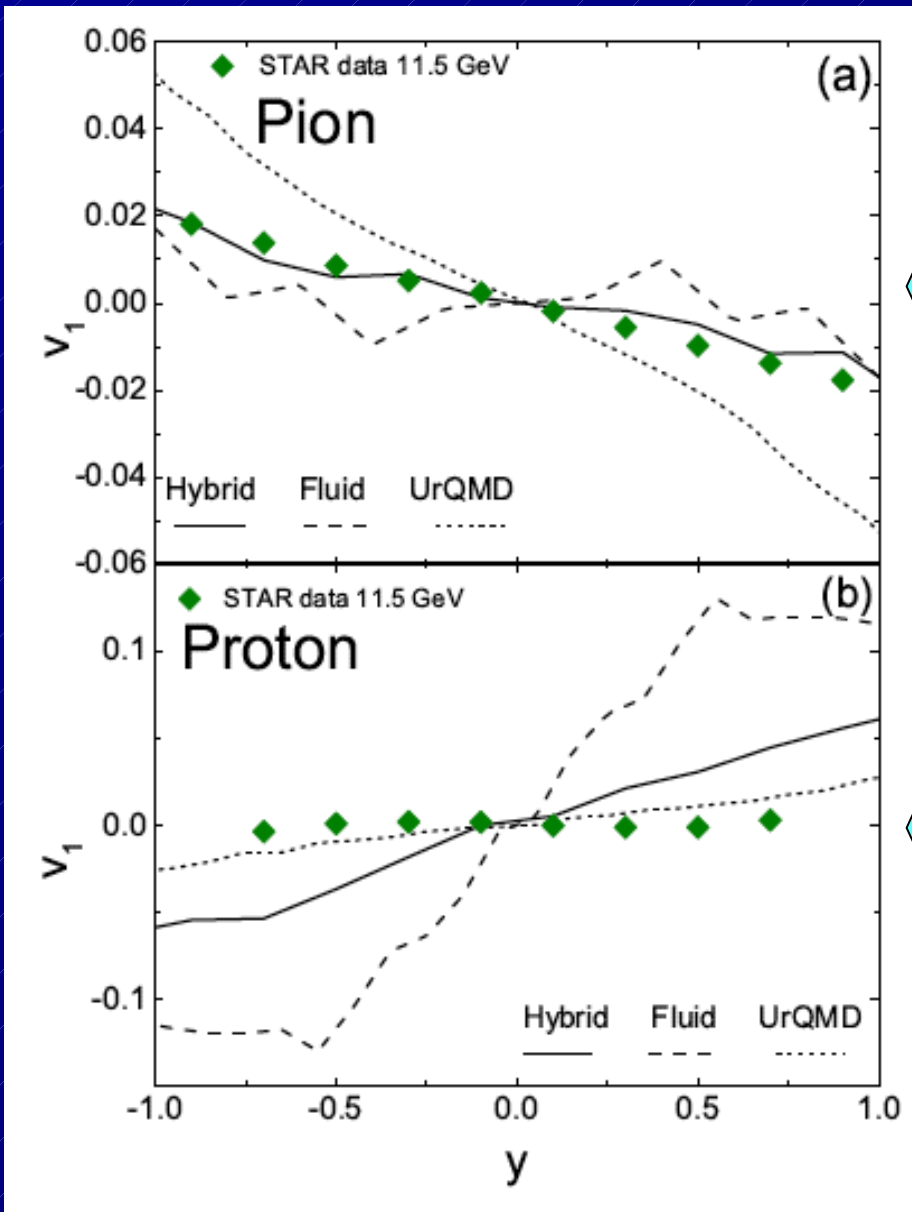
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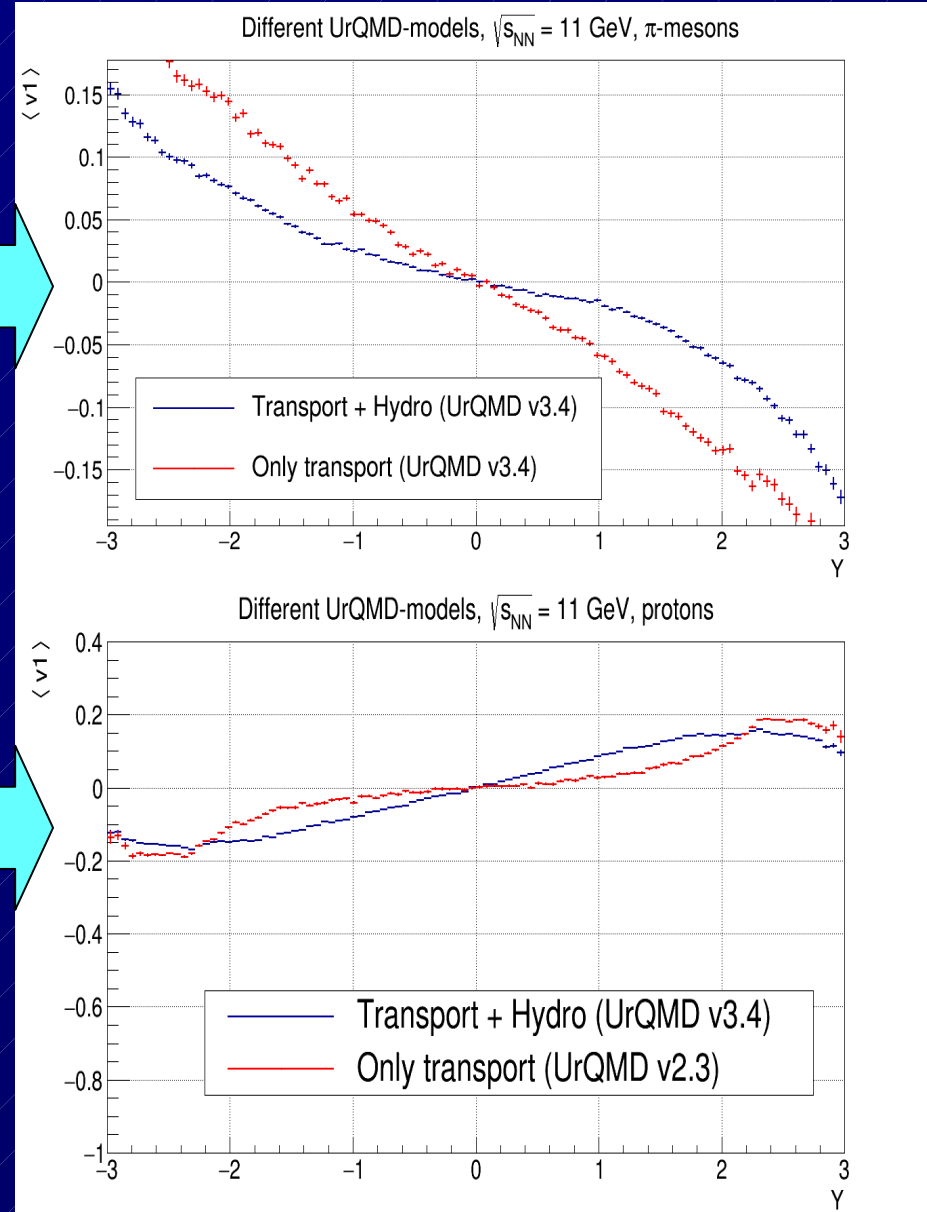
**Additional Slides**

# Further results of test simulations – flow observables

J. Steinheimer et al., Phys. Rev. C89 (2014) 054913



Reproduced!



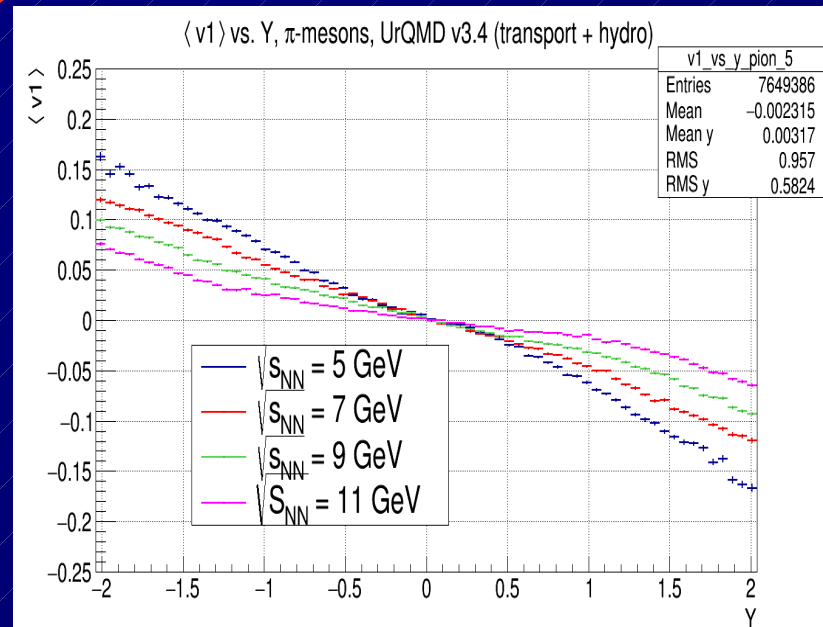
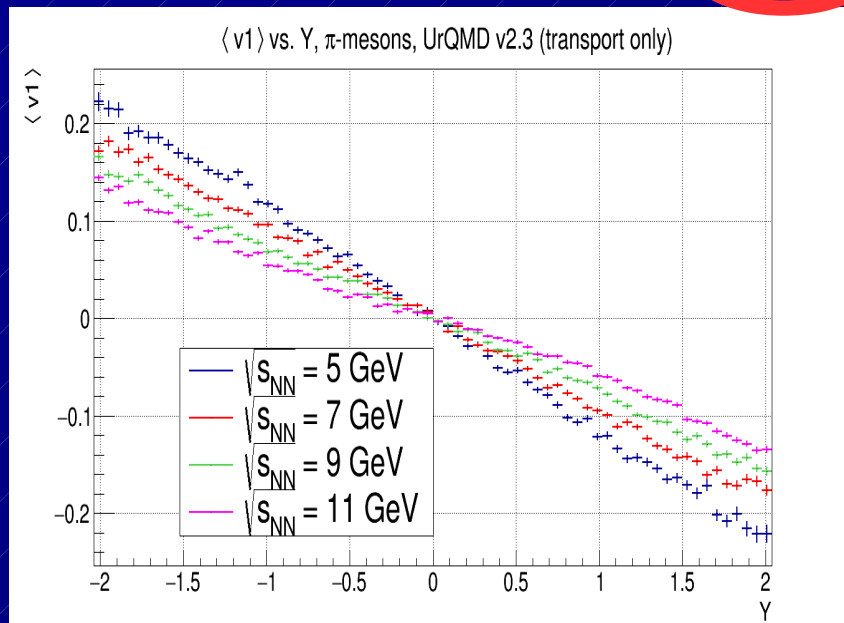
# Further results of test simulations – flow observables

NICA energy scan: **UrQMD**

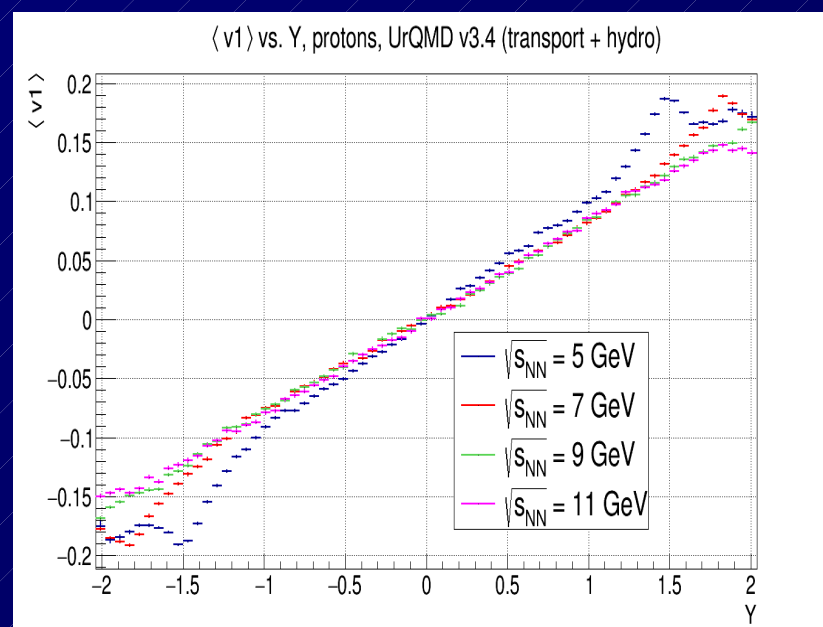
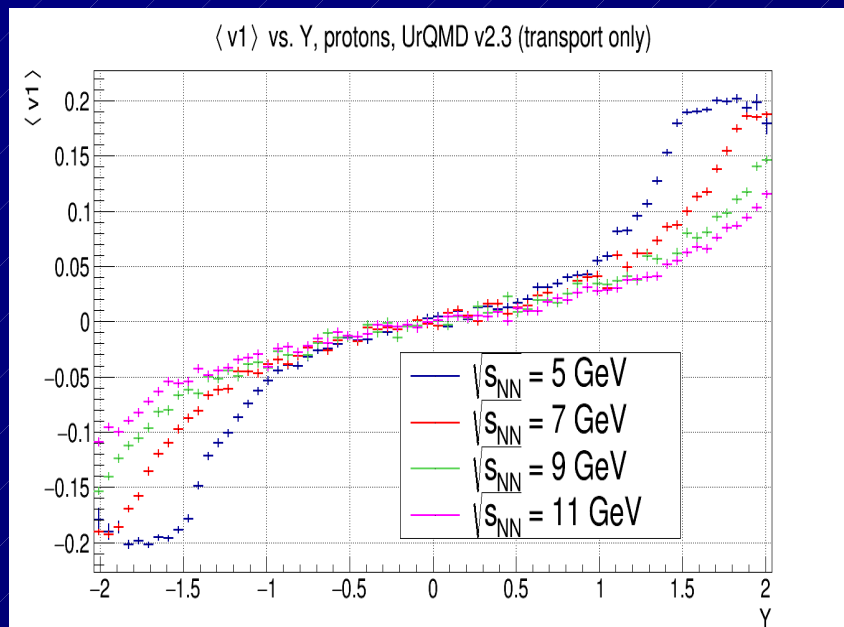
$\langle v_1 \rangle$

UrQMD + hydro (1<sup>st</sup> order PT)

Pions



Protons



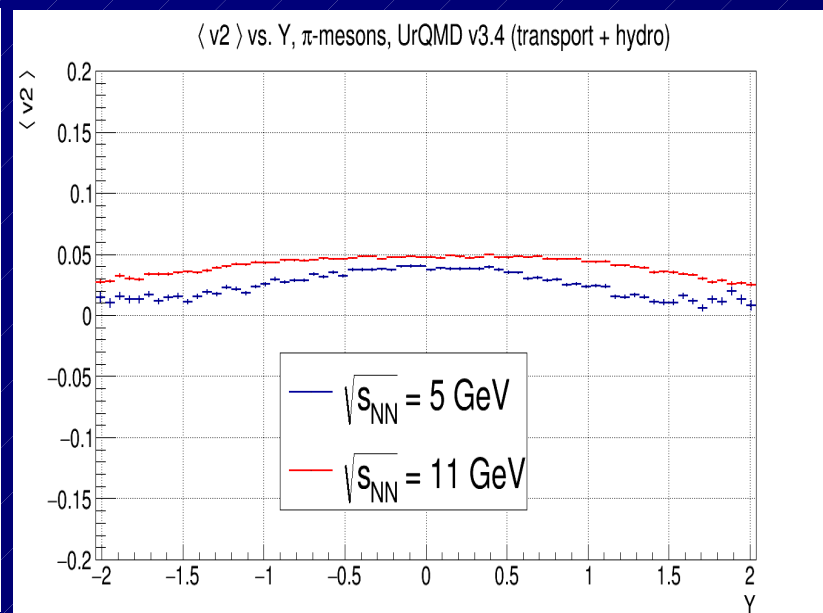
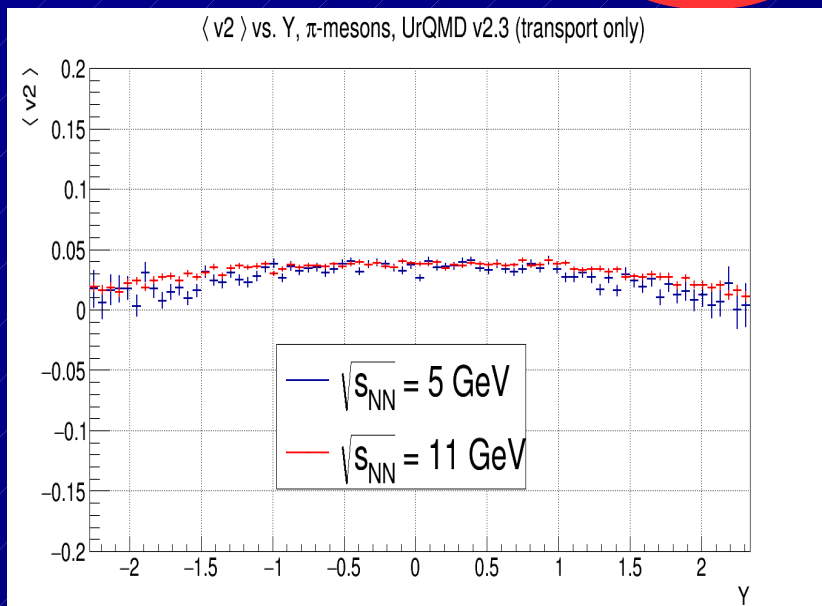
# Further results of test simulations – flow observables

NICA energy scan: UrQMD

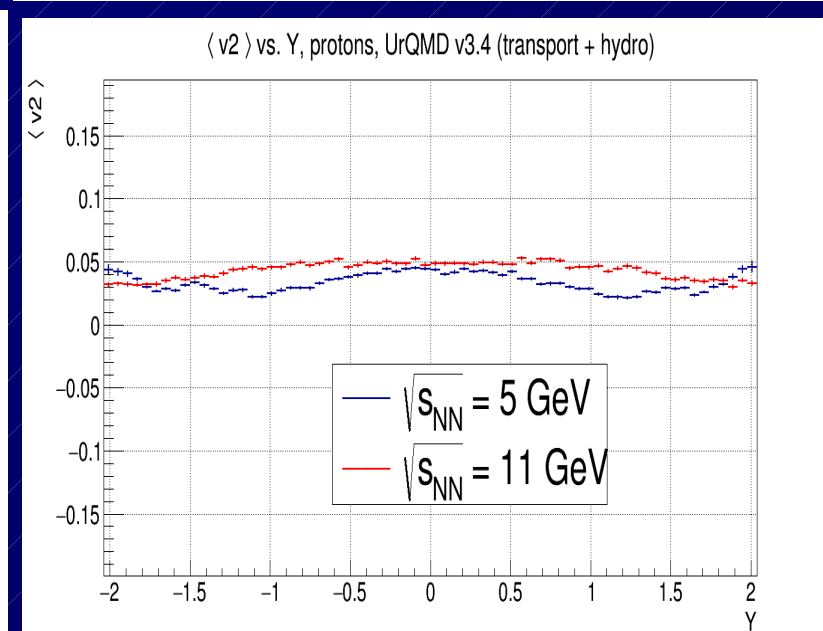
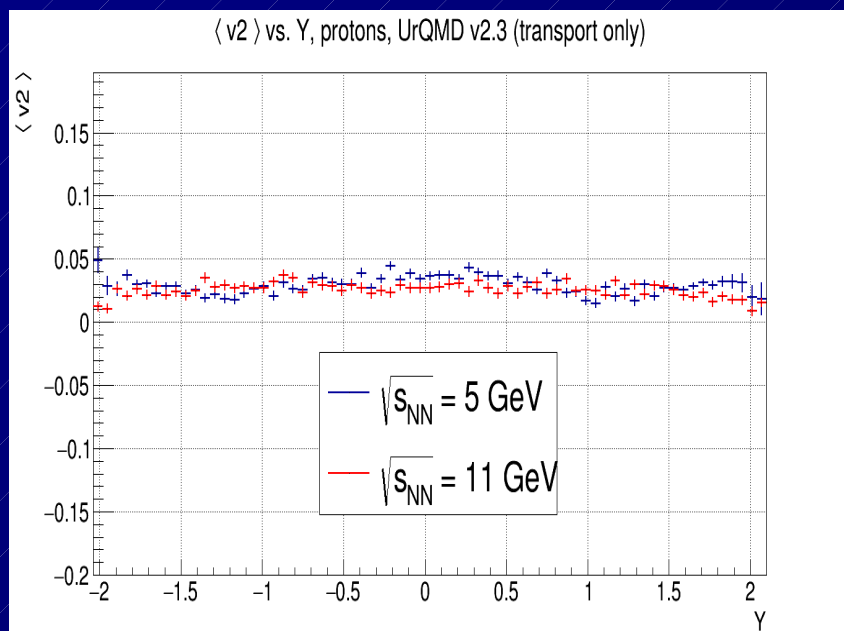
$\langle v_2 \rangle$

UrQMD + hydro (1<sup>st</sup> order PT)

Pions



Protons

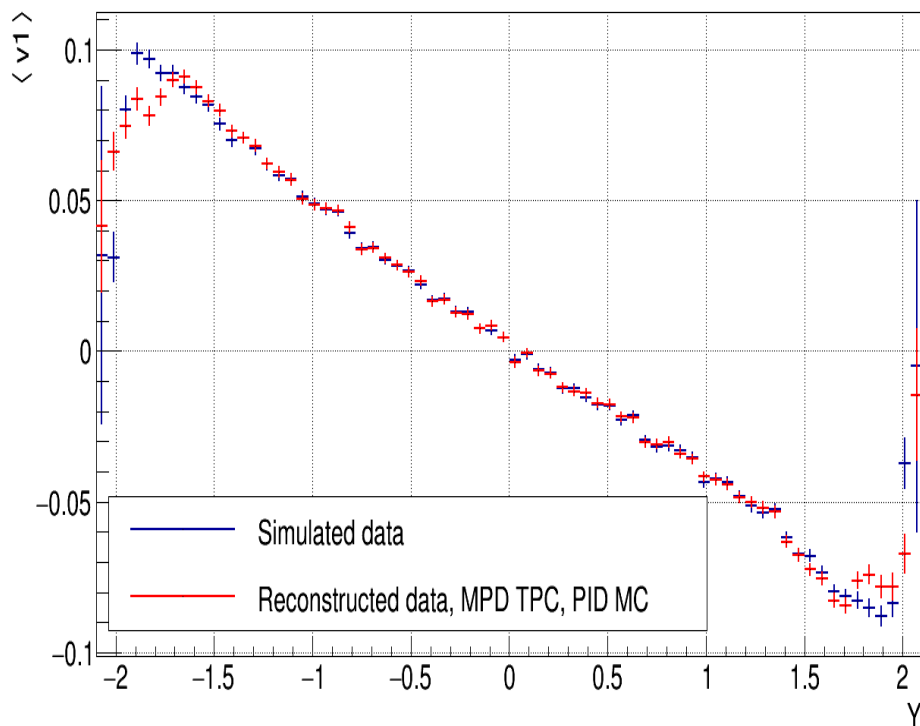




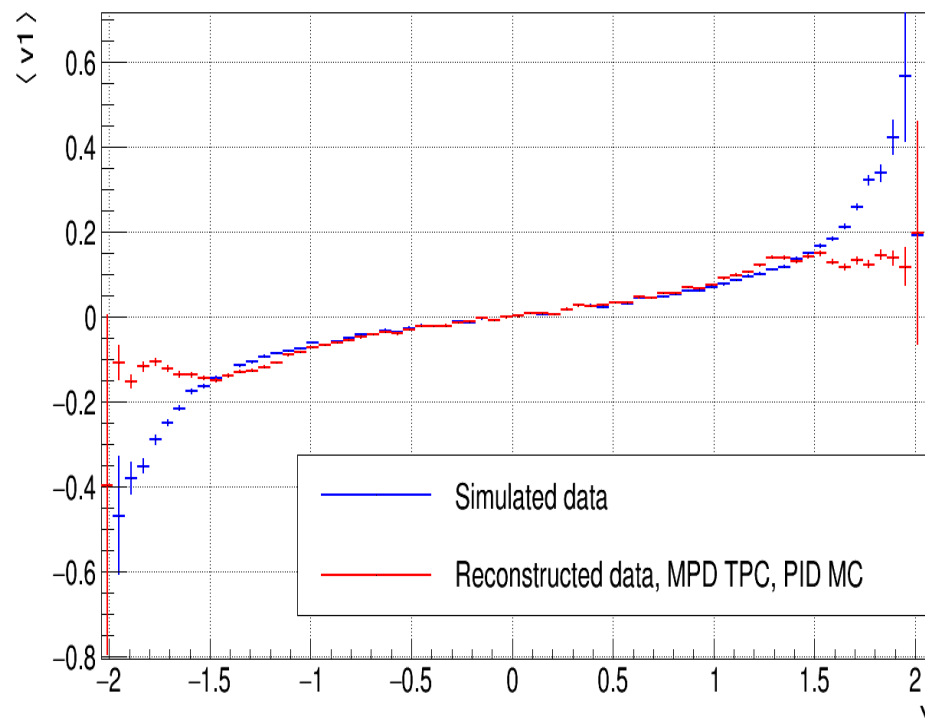
# Further results of test simulations – flow observables

Detector Simulation with GEANT : Excellent reproduction of simulation results !

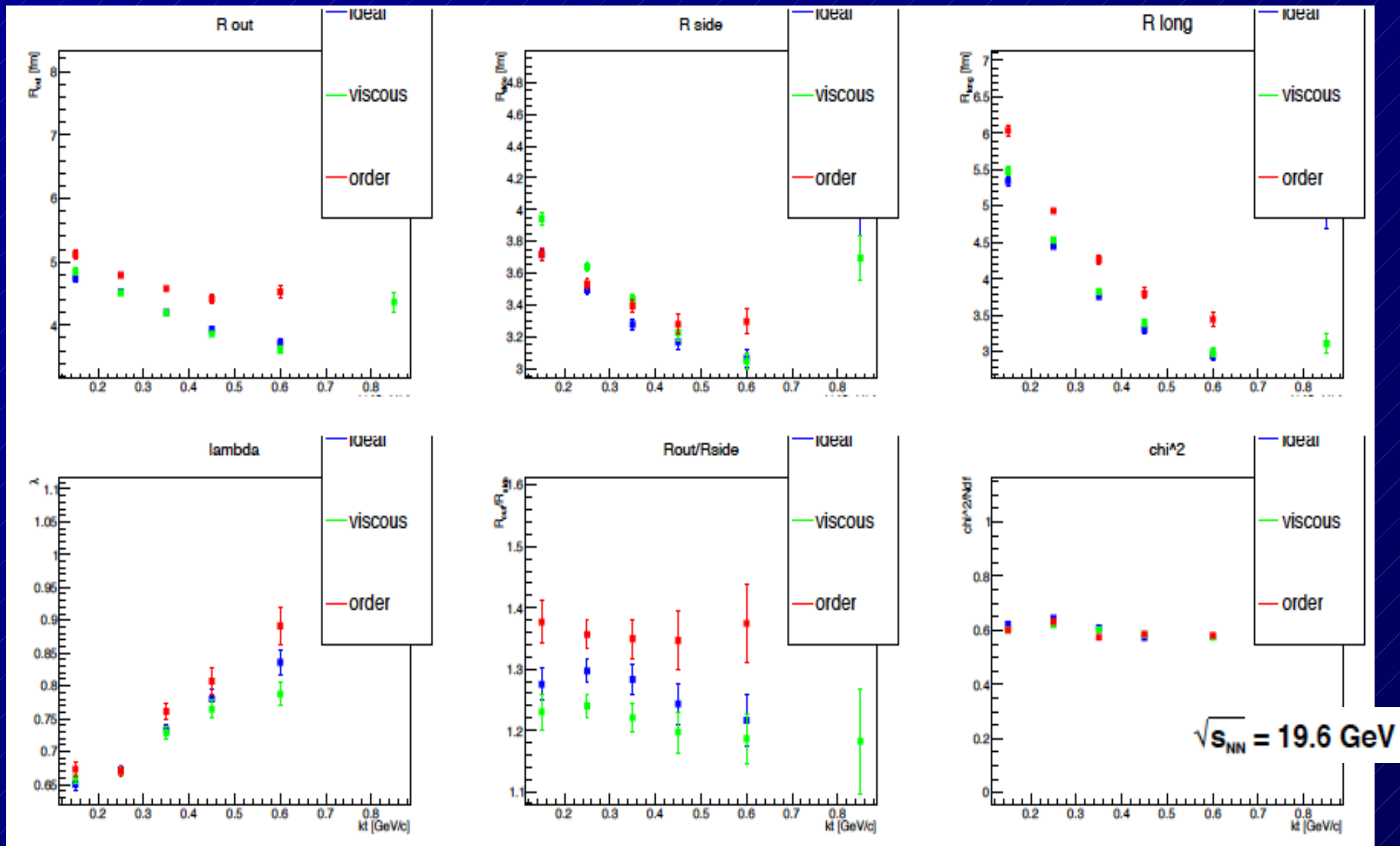
$\langle v_1 \rangle$  vs.  $Y$ ,  $\pi$ -mesons,  $\sqrt{s_{NN}} = 9$  GeV

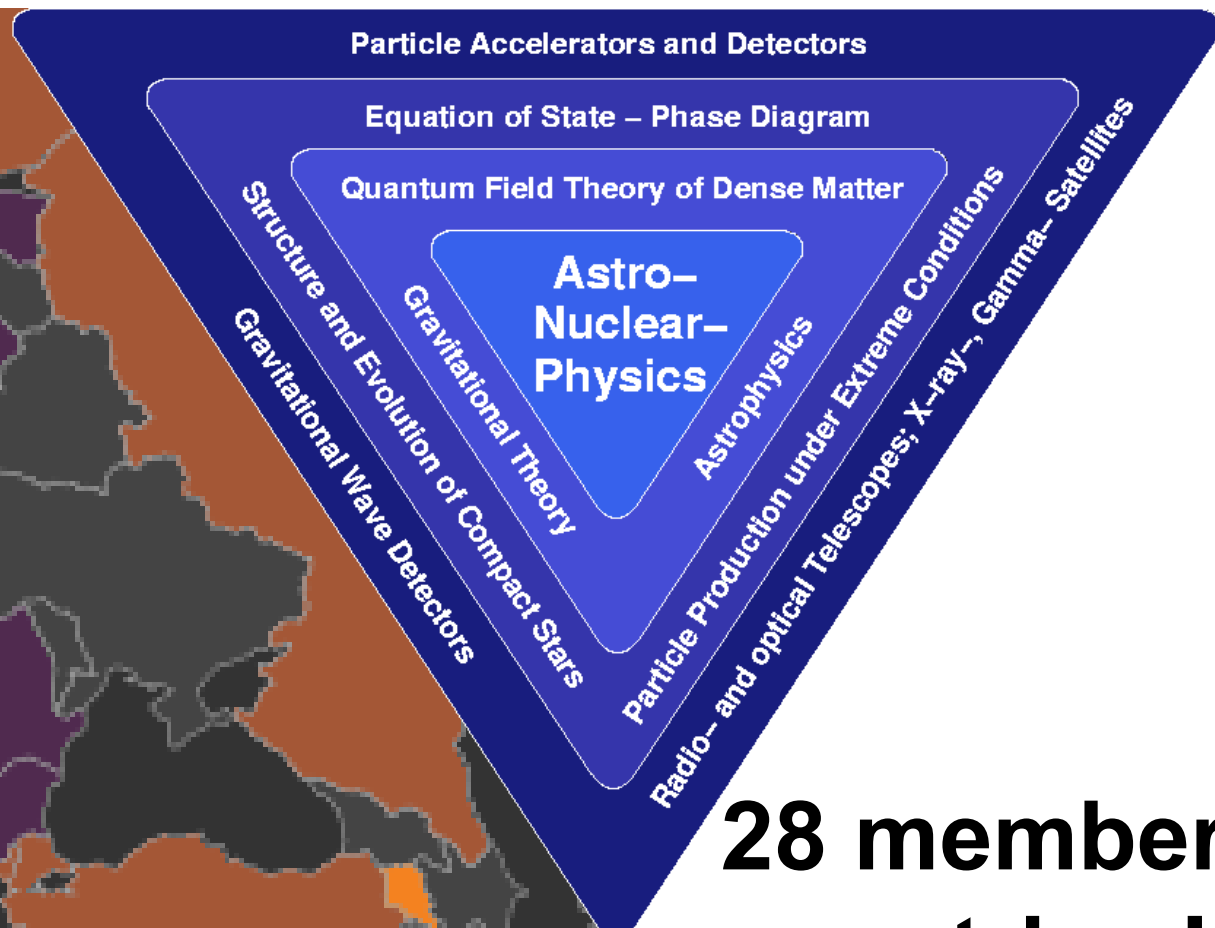
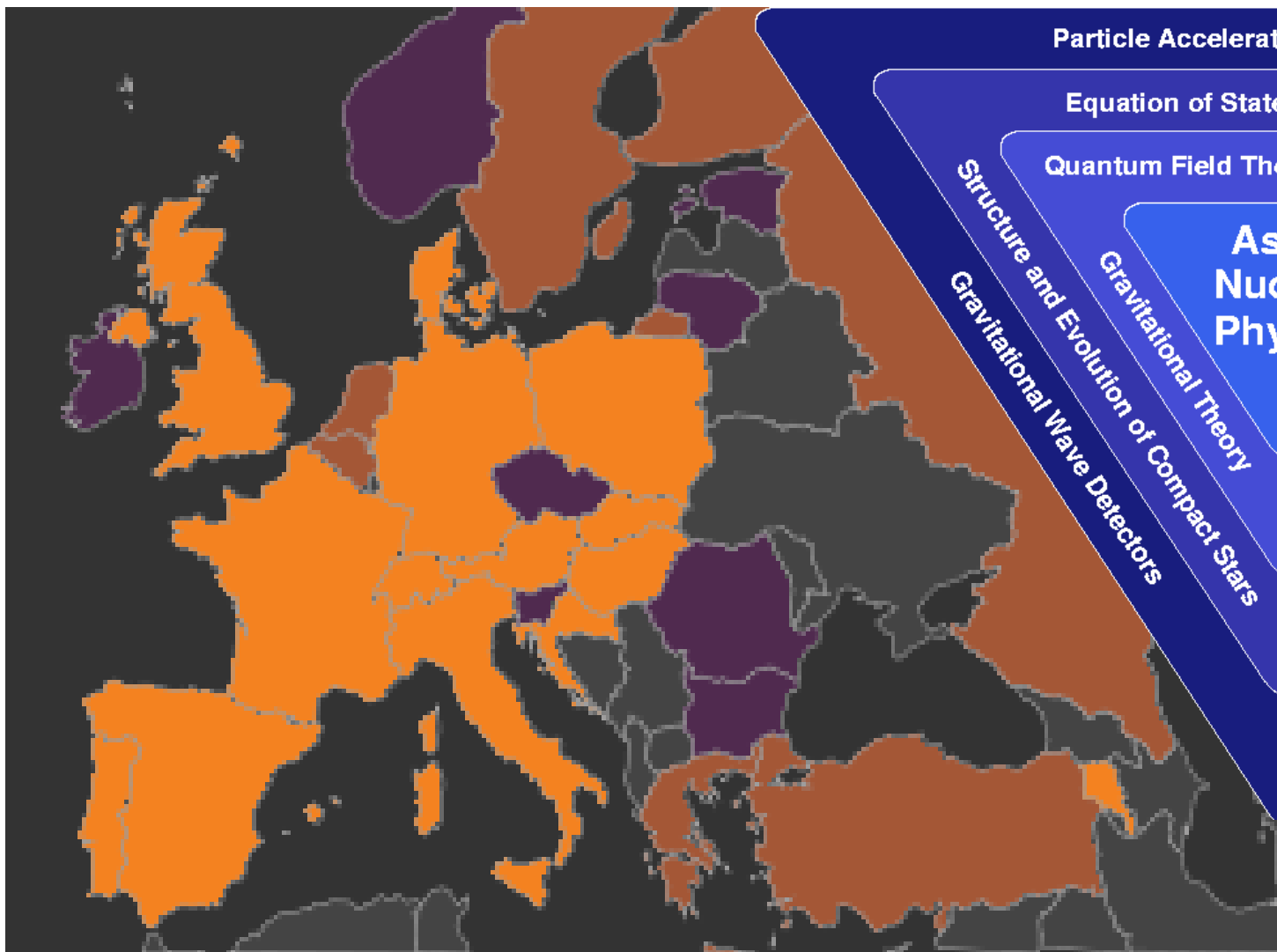


$\langle v_1 \rangle$  vs.  $Y$ , protons,  $\sqrt{s_{NN}} = 9$  GeV



# Further results of test simulations – HBT radii





**28 member  
countries !!  
(MP1304)**

**New**



**Kick-off: Brussels, November 25, 2013**

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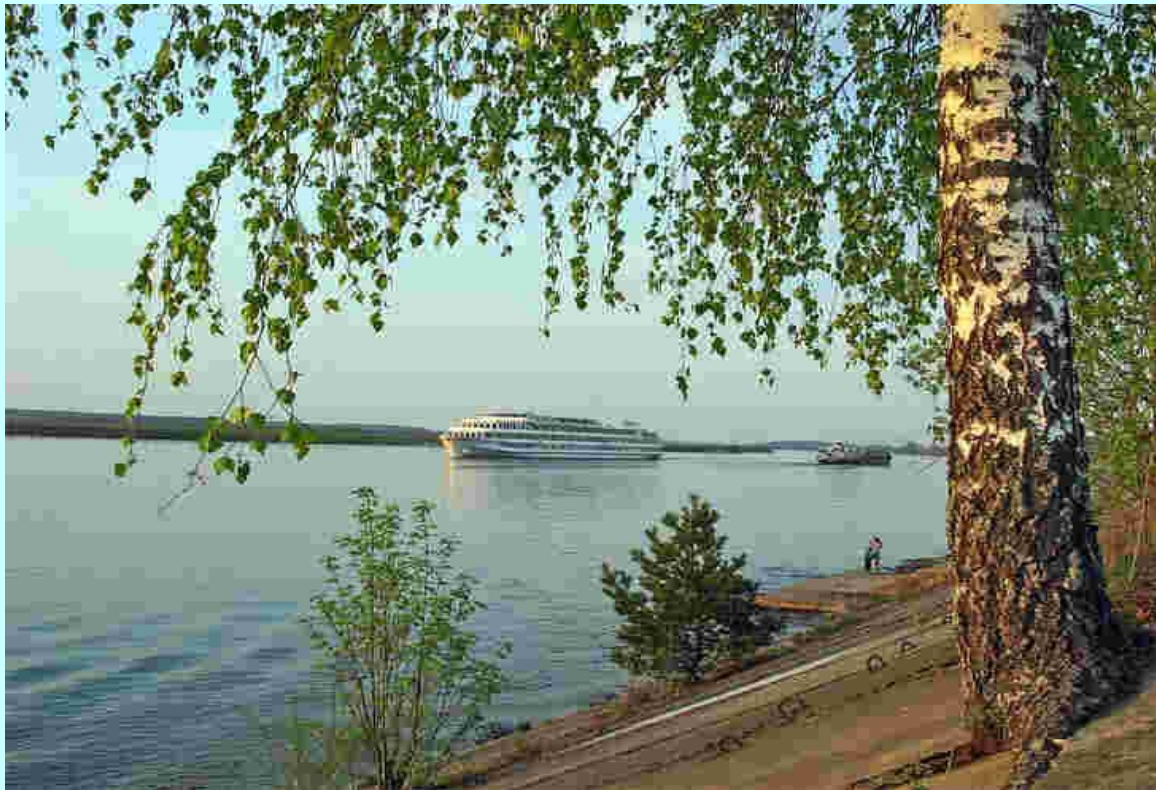
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# Strangeness in Quark Matter 2015

Dubna, 6.-11. July 2015



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Website: <http://sqm.jinr.ru>

Satellite Meetings:

Summer School “Dense Matter”, Dubna, June 29 – July 11, 2015

Roundtable “Physics at NICA”, Dubna, 5. July 2015