

# Relativistic dissipative hydrodynamics for particles of arbitrary mass

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- 1 Motivation: Why study fluids and transport?
- 2 Goal: Dissipative Hydrodynamics
- 3 Tool: Kinetic theory
- 4 Results: Transport equations for hydrodynamics
  - Transport coefficients for arbitrary masses
  - Non-relativistic limit reproduces the Grad equations

# Phenomenology of fluid-dynamics: transport coefficients

Slides by Yuuka Kanakubo. *Dynamical evolution and collective dynamics.* Apr. 2025. URL:

<https://indico.cern.ch/event/1334113/contributions/6369376/>.

How does it flow?

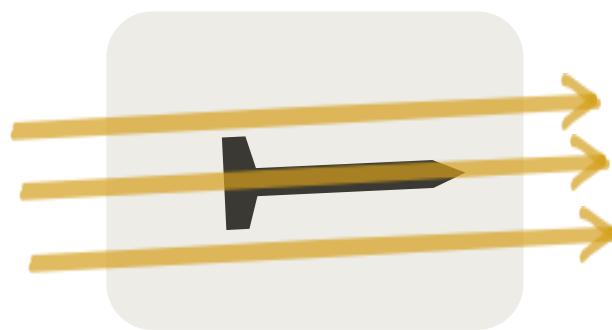


like water or honey?

Characterised by transport coefficient of Apfelwein

# Transport coefficients

How it responds when you "hit" it



Electric field



$$J = \sigma E$$

Electric current

## Electric conductivity

How easily electric current can flow through the nail

=> Property of the nail!

# Transport coefficients

How it responds when you "hit" it



Temperature gradient



Heat current

$$q^i = -\kappa \partial^i T$$

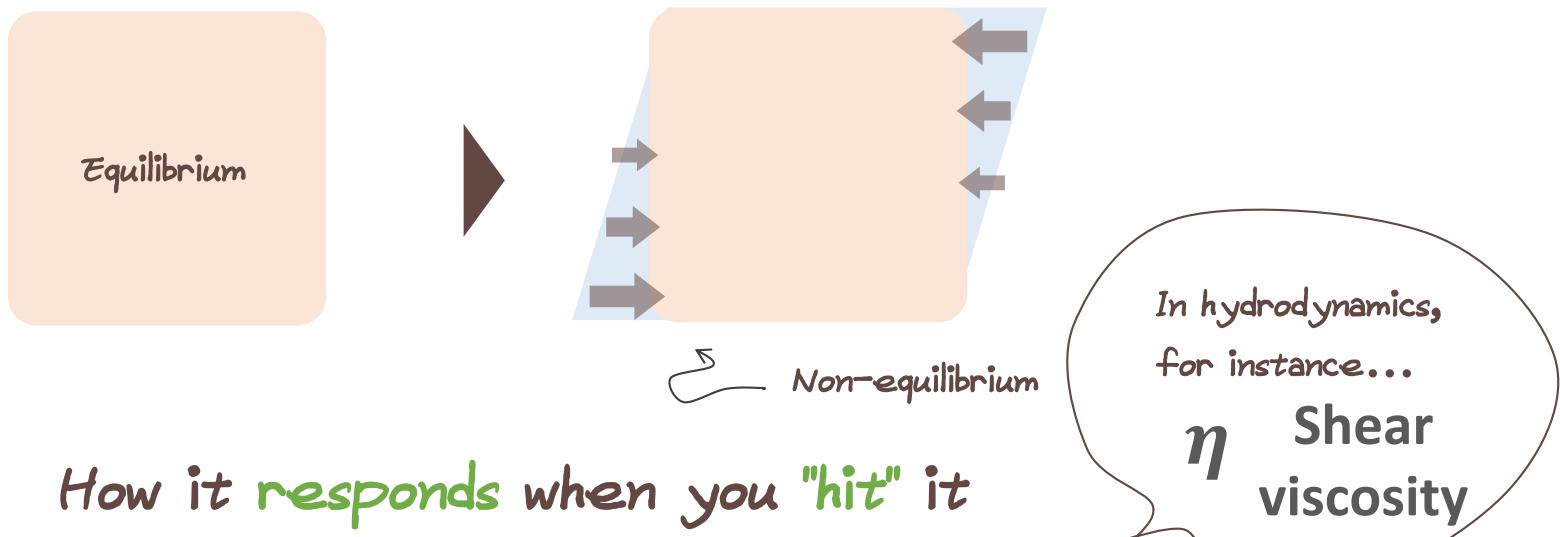
## Heat conductivity

How easily the cup can be heated up

=> Property of the cup!

# Transport coefficients

Drag the system slightly away from equilibrium



How it **responds** when you "hit" it  
=> **Dynamical property** of matter!

In hydrodynamics,  
for instance...

$\eta$  Shear  
viscosity

\*Linear response theory -> Transport coefficients can be derived microscopically

How does it flow?



like water or honey?

Characterised by transport coefficient of QGP

# Second-order kinetic equation for non-uniform gases



- ▶ Navier-Stokes (first-order) equations are universal for different gases **close to equilibrium**, but fail for more complicated flows.
- ▶ Going further out-of-equilibrium requires **higher-order** corrections to hydrodynamical quantities. 😢
- ▶ Grad's equations provide the evolution of **dissipative** quantities of the fluid at **second order**:

$$-\frac{2}{3} \frac{P_0}{\eta} \dot{\boldsymbol{q}}^i = \frac{d\boldsymbol{q}^i}{dt} + \frac{5}{2} P_0 \frac{\partial RT}{\partial x_i} + \frac{7}{5} \dot{\boldsymbol{q}}^i \frac{\partial u_k}{\partial x_k} + \dot{\boldsymbol{q}}_k \frac{\partial u^i}{\partial x_k} + \frac{2}{5} \dot{\boldsymbol{q}}_k \left( \frac{\partial u^i}{\partial x_k} + \frac{\partial u^k}{\partial x_i} \right) \quad (1)$$

$$- RT \pi^{ik} \frac{\partial \ln P_0}{\partial x^k} + RT \frac{\partial \pi^{ik}}{\partial x^k} + \frac{7}{2} \pi^{ik} \frac{\partial RT}{\partial x^k} - \frac{\pi^{ik}}{\rho_0} \frac{\partial \pi_{kl}}{\partial x_l},$$

$$-\frac{P_0}{\eta} \dot{\pi}^{ij} = \frac{d\pi^{ij}}{dt} + 2P_0 \frac{\partial u^{<i}}{\partial x_j} + \pi^{ij} \frac{\partial u_k}{\partial x_k} + 2\pi^{r<i} \frac{\partial u^j}{\partial x^r} + \frac{4}{5} \frac{\partial \dot{\boldsymbol{q}}^{<i}}{\partial x_j}, \quad (2)$$

where  $\boldsymbol{q}$  is the heat flow,  $\rho_0$  the mass density, and  $R \equiv k_B/m \equiv 1/m$  the gas constant per unit mass.

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Navier-Stokes fails to predict the correct flow profile for, e.g., force-driven Poiseuille flow.

# Brief reminder of deriving hydrodynamics from kinetic theory

Based on the talk by David Wagner. *Inverse Reynolds-Dominance approach to transient fluid dynamics.* June 2022. URL: <https://itp.uni-frankfurt.de/~hees/transport-meeting/ss22/talk-Wagner.pdf>.

## Hydrodynamics: Conservation equations

$$\partial_\mu T^{\mu\nu} = 0 , \quad \partial_\mu N^\mu = 0 \quad (3)$$

- ▶ Hydrodynamics from  $(4 + 1 = 5)$  conservation equations
  - Ideal case: Sufficient (if equation of state is supplied)
    - Variables:  $\epsilon, n, u^\mu$
  - Dissipative case: Underdetermined
    - Variables:  $\epsilon, n, u^\mu, \Pi, V^\mu, \pi^{\mu\nu}$
- ▶ Fundamental question: How to obtain the dissipative components of  $N^\mu$  and  $T^{\mu\nu}$ ? 😕

## Decomposition of conserved currents (Landau frame)

$$N^\mu = n u^\mu + V^\mu \quad (4)$$

$$T^{\mu\nu} = \epsilon u^\mu u^\nu - (P + \Pi) \Delta^{\mu\nu} + \pi^{\mu\nu} \quad (5)$$

Projectors:  $\Delta^{\mu\nu} \equiv g^{\mu\nu} - u^\mu u^\nu$ ,  $\overline{\Delta}_{\alpha\beta}^{\mu\nu} \equiv (\Delta_\alpha^\mu \Delta_\beta^\nu + \Delta_\beta^\mu \Delta_\alpha^\nu)/2 - \Delta^{\mu\nu} \Delta_{\alpha\beta}/3$

# Equations of hydrodynamics

Projecting the derivative,

$$\partial_\mu A = (u_\mu u_\nu \partial^\nu + \Delta_{\mu\nu} \partial^\nu) A \equiv u_\mu \dot{A} + \nabla_\mu A ,$$

we project the conservation equations:

$$\partial_\mu N^\mu = 0 \Leftrightarrow \dot{n}_0 + n_0 \theta + \partial_\mu V^\mu = 0 , \quad (6)$$

$$u_\alpha \partial_\beta T^{\alpha\beta} = 0 \Leftrightarrow \dot{\varepsilon}_0 + (\varepsilon_0 + P_0 + \Pi) \theta - \pi^{\alpha\beta} \sigma_{\alpha\beta} = 0 , \quad (7)$$

$$\Delta_\alpha^\mu \partial_\beta T^{\alpha\beta} = 0 \Leftrightarrow (\varepsilon_0 + P_0 + \Pi) \dot{u}^\mu - \nabla^\mu (P_0 + \Pi) + \Delta_\alpha^\mu \partial_\beta \pi^{\alpha\beta} = 0 , \quad (8)$$

which are generic equations for non-fluctuating **dissipative relativistic fluid**.

► To solve these fluid equations we need 

- an equation of state – ideal-gas  $P_0 = n_0 T$
- constitutive relations for  $\{\Pi, V, \pi\}$  – TBD

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Notation:  $A^{\langle\mu} B^{\nu\rangle} \equiv \Delta_{\alpha\beta}^{\mu\nu} A^\alpha B^\beta$ ;  $\theta \equiv \nabla^\mu u_\mu$ ,  $\sigma^{\mu\nu} \equiv \nabla^{\langle\mu} u^{\nu\rangle}$

# First- and second-order hydrodynamics

- ▶ First-order hydro: Relate **dissipative quantities** to fluid-dynamical gradients

$$\Pi = -\zeta \theta, \quad V^\mu = \kappa I^\mu, \quad \pi^{\mu\nu} = 2\eta \sigma^{\mu\nu} \quad (9)$$

- ▶ (In Eckart or Landau frame): **Acausal!**
- ▶ Second-order hydro: Treat dissipative quantities as dynamical, provide relaxation equations

## Relaxation equations

$$\tau_\Pi \dot{\Pi} + \Pi = -\zeta \theta + \text{h.o.t.} \quad (10a)$$

$$\tau_V \dot{V}^{\langle\mu\rangle} + V^\mu = \kappa I^\mu + \text{h.o.t.} \quad (10b)$$

$$\tau_\pi \dot{\pi}^{\langle\mu\nu\rangle} + \pi^{\mu\nu} = 2\eta \sigma^{\mu\nu} + \text{h.o.t.} \quad (10c)$$

- ▶ Needs input from **microscopic theory** 
- ▶ This talk: derived from **kinetic theory**

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$$I^\mu \equiv \nabla^\mu \alpha, \quad \alpha \equiv \mu/T$$

- ▶ Describe system in  $(x, k)$ -phase space through one-particle distribution function  $f_{\mathbf{k}}(x)$
- ▶ Connection to hydrodynamics through conserved currents

## Conserved quantities (moments of the distribution function)

$$N^\mu = \int dK f_{\mathbf{k}}(x) k^\mu , \quad T^{\mu\nu} = \int dK f_{\mathbf{k}}(x) k^\mu k^\nu \quad (11)$$

- ▶ Hydrodynamic quantities are distributed through  $f_{\mathbf{k}}(x)$ 
  - Governed by Boltzmann equation  $k^\mu \partial_\mu f_{\mathbf{k}}(x) = C[f]$
  - Only elastic  $2 \leftrightarrow 2$  scattering
- ▶ Separate into equilibrium part  $f_{0\mathbf{k}}(x)$  and deviation  $\delta f_{\mathbf{k}}(x)$ 
  - $f_{0\mathbf{k}}(x)$  determined by  $C[f_0] = 0$

---

$$dK \equiv d^3k / [(2\pi)^3 k^0]$$

Equilibrium distribution:  $f_{0\mathbf{k}}(x) = [\exp(-\alpha(x) + \beta(x)E_{\mathbf{k}}) + a]^{-1}$

- ▶  $a \in \{-1, 0, 1\}$  determined by statistics of particles
- ▶  $\alpha, \beta, u^\mu$ : Lagrange multipliers (fields)

Thermodynamic quantities can be determined from one integral:

## Thermodynamic integral

$$I_{nq} \equiv \frac{(-1)^q}{(2q+1)!!} \int dK f_{0\mathbf{k}} E_{\mathbf{k}}^{n-2q} (\Delta^{\alpha\beta} k_\alpha k_\beta)^q , \quad (12)$$

for example, density  $n_0 \equiv I_{10}$ , pressure  $P_0 \equiv I_{21}$ , etc.

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$$E_{\mathbf{k}} \equiv u^\mu(x) k_\mu$$

# Non-equilibrium: Moment expansion

- ▶ Question: Which parts of  $\delta f_{\mathbf{k}}(x)$  in momentum space are important for hydrodynamics?
- ▶ Expand in terms of complete and orthogonal basis  $\{1, k^{\langle \mu \rangle}, k^{\langle \mu_1 \mu_2 \rangle}, \dots\}$ 
  - Equivalent to spherical harmonics (angular part) and a radial part



## Expansion of $\delta f$

$$\delta f(x, k) = f_0 \tilde{f}_0 \sum_{\ell=0}^{\infty} \sum_{n=0}^{N_{\ell}} \mathcal{H}_{\mathbf{k}n}^{(\ell)} k^{\langle \mu_1 \dots \mu_{\ell} \rangle} \rho_{n, \mu_1 \dots \mu_{\ell}}(x) \quad (13)$$

- ▶ Irreducible moments  $\rho_n^{\mu_1 \dots \mu_{\ell}}$  carry all information

## Irreducible moments

$$\rho_r^{\mu_1 \dots \mu_{\ell}}(x) \equiv \int dK E_{\mathbf{k}}^r k^{\langle \mu_1 \dots \mu_{\ell} \rangle} \delta f_{\mathbf{k}}(x) \quad (14)$$

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$$\tilde{f}_0 \equiv 1 - a f_0$$

# Equations of motion

## Boltzmann equation

$$\delta \dot{f} = E_{\mathbf{k}}^{-1} C[f] - \dot{f}_0 - E_{\mathbf{k}}^{-1} k^\mu \nabla_\mu (f_0 + \delta f) \quad (15)$$

- ▶ Boltzmann equation determines evolution of all moments
  - Infinite set of coupled differential equations

## Moment equations

$$(\ell = 0) \quad \dot{\rho}_r + \sum_{n=0, \neq 1, 2}^{N_0} \mathcal{A}_{rn}^{(0)} \rho_n = \alpha_r^{(0)} \theta + \text{h.o.t.} \quad (16a)$$

$$(\ell = 1) \quad \dot{\rho}_r^{\langle \mu \rangle} + \sum_{n=0, \neq 1}^{N_1} \mathcal{A}_{rn}^{(1)} \rho_n^\mu = \alpha_r^{(1)} I^\mu + \text{h.o.t.} \quad (16b)$$

$$(\ell = 2) \quad \dot{\rho}_r^{\langle \mu \nu \rangle} + \sum_{n=0}^{N_2} \mathcal{A}_{rn}^{(2)} \rho_n^{\mu \nu} = 2 \alpha_r^{(2)} \sigma^{\mu \nu} + \text{h.o.t.} \quad (16c)$$

$$(\ell > 2) \quad \dot{\rho}_r^{\langle \mu_1 \cdots \mu_\ell \rangle} + \sum_{n=0}^{N_\ell} \mathcal{A}_{rn}^{(\ell)} \rho_n^{\mu_1 \cdots \mu_\ell} = \text{h.o.t.} \quad (16d)$$

## ▶ Problem?

Notation:  $\dot{f} \equiv u^\mu \partial_\mu f$ ; Matching conditions:  $\rho_1 = \rho_2 = \rho_1^\mu = 0$



# Irreducible moment equation of rank two



Just to give an impression:

$$\begin{aligned}
& \dot{\rho}_r^{\langle\mu\nu\rangle} + \sum_{n=0}^{N_2} \mathcal{A}_{rn}^{(2)} \rho_n^{\mu\nu} = 2 [I_{r+2,1} + (r-1)I_{r+2,2}] \sigma^{\mu\nu} \\
& - \frac{2}{7} \left[ (2r+5) \rho_r^{\lambda\langle\mu} - 2m^2(r-1) \rho_{r-2}^{\lambda\langle\mu} \right] \sigma_\lambda^{\nu\rangle} \\
& + 2 \rho_r^{\lambda\langle\mu} \omega_\lambda^{\nu\rangle} + \frac{2}{15} \left[ (r+4) \rho_{r+2} - (2r+3)m^2 \rho_r + (r-1)m^4 \rho_{r-2} \right] \sigma^{\mu\nu} \\
& + \frac{2}{5} \nabla^{\langle\mu} \left( \rho_{r+1}^{\nu\rangle} - m^2 \rho_{r-1}^{\nu\rangle} \right) - \frac{2}{5} \left[ (r+5) \rho_{r+1}^{\langle\mu} - m^2 r \rho_{r-1}^{\langle\mu} \right] \dot{u}^{\nu\rangle} \\
& - \frac{1}{3} \left[ (r+4) \rho_r^{\mu\nu} - m^2(r-1) \rho_{r-2}^{\mu\nu} \right] \theta \\
& + (r-1) \rho_{r-2}^{\mu\nu\lambda\kappa} \sigma_{\lambda\kappa} - \Delta_{\alpha\beta}^{\mu\nu} \nabla_\lambda \rho_r^{\lambda\alpha\beta} + r \rho_{r-1}^{\lambda\mu\nu} \dot{u}_\lambda ,
\end{aligned}$$

$$\begin{aligned}
\mathcal{A}_{rn}^{(\ell)} &= \frac{1}{\nu(2\ell+1)} \int dK dK' dP dP' W_{\mathbf{k}\mathbf{k}' \rightarrow \mathbf{p}\mathbf{p}'} f_{0\mathbf{k}} f_{0\mathbf{k}'} \tilde{f}_{0\mathbf{p}} \tilde{f}_{0\mathbf{p}'} E_{\mathbf{k}}^{r-1} k^{\langle\mu_1} \dots k^{\mu_\ell\rangle} \\
&\times \left( \mathcal{H}_{\mathbf{k}n}^{(\ell)} k_{\langle\mu_1} \dots k_{\mu_\ell\rangle} + \mathcal{H}_{\mathbf{k}'n}^{(\ell)} k'_{\langle\mu_1} \dots k'_{\mu_\ell\rangle} - \mathcal{H}_{\mathbf{p}n}^{(\ell)} p_{\langle\mu_1} \dots p_{\mu_\ell\rangle} - \mathcal{H}_{\mathbf{p}'n}^{(\ell)} p'_{\langle\mu_1} \dots p'_{\mu_\ell\rangle} \right) .
\end{aligned}$$



- ▶ Basic idea: Power-counting scheme to **second order** in two small quantities:
  1. Knudsen number  $\text{Kn} \sim \lambda_{\text{mfp}} / \lambda_{\text{hydro}}$ , and
  2. inverse Reynolds numbers  $\text{IRe} \sim \delta f / f_0$
- ▶ Interested in the evolution of  $T^{\mu\nu}$  and  $N^\mu$  as moments of  $\delta f$ 
  - Benchmark: Evolution equations for  $\Pi = -(m^2/3)\rho_0$ ,  $V^\mu = \rho_0^\mu$ ,  $\pi^{\mu\nu} = \rho_0^{\mu\nu}$
  - Only interested in moments with  $\ell \leq 2$
- ▶  $\rho_r^{\mu_1 \dots \mu_{\ell > 2}} = 0$ , corrections of order  $\mathcal{O}(\text{Kn}^2 \text{IRe}, \text{Kn}^3)$

## Moment equations

$$\sum_{n=0, \neq 1, 2}^{N_0} \tau_{rn}^{(0)} \dot{\rho}_n + \rho_r = -\zeta_r \theta + \text{h.o.t.} \quad (17a)$$

$$\sum_{n=0, \neq 1}^{N_1} \tau_{rn}^{(1)} \dot{\rho}_n^{\langle \mu \rangle} + \rho_r^\mu = \kappa_r I^\mu + \text{h.o.t.} \quad (17b)$$

$$\sum_{n=0}^{N_2} \tau_{rn}^{(2)} \dot{\rho}_n^{\langle \mu \nu \rangle} + \rho_r^{\mu \nu} = 2\eta_r \sigma^{\mu \nu} + \text{h.o.t.} \quad (17c)$$

- ▶ Still coupled system of  $N_0 + 3N_1 + 5N_2$  equations
- ▶ **How to decouple the remaining equations?**

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$$\tau^{(\ell)} \equiv (\mathcal{A}^{(\ell)})^{-1}$$

- ▶ Use asymptotic matching conditions to Navier-Stokes relations
- ▶ Express all irreducible moments through **dissipative quantities**
- ▶ Re-sum the relaxation modes into  $\tau$  and discard h.o.t

## Hydrodynamic transport equations (IReD)

$$\tau_{II} \dot{H} + H = -\zeta \theta + \mathcal{J} + \mathcal{R} \quad (18a)$$

$$\tau_V \dot{V}^{\langle\mu\rangle} + V^\mu = \kappa V^\mu + \mathcal{J}^\mu + \mathcal{R}^\mu \quad (18b)$$

$$\tau_\pi \dot{\pi}^{\langle\mu\nu\rangle} + \pi^{\mu\nu} = 2\eta \sigma^{\mu\nu} + \mathcal{J}^{\mu\nu} + \mathcal{R}^{\mu\nu} \quad (18c)$$

- ▶ Only terms  $\sim \mathcal{O}(IRe)$ ,  $\sim \mathcal{O}(Kn)$ ,  $\sim \mathcal{O}(Kn IRe)$ ,  $\sim \mathcal{O}(IRe^2)$  appear

# Example of couplings at second order: bulk viscosity

- ▶ In the bulk relaxation equation,

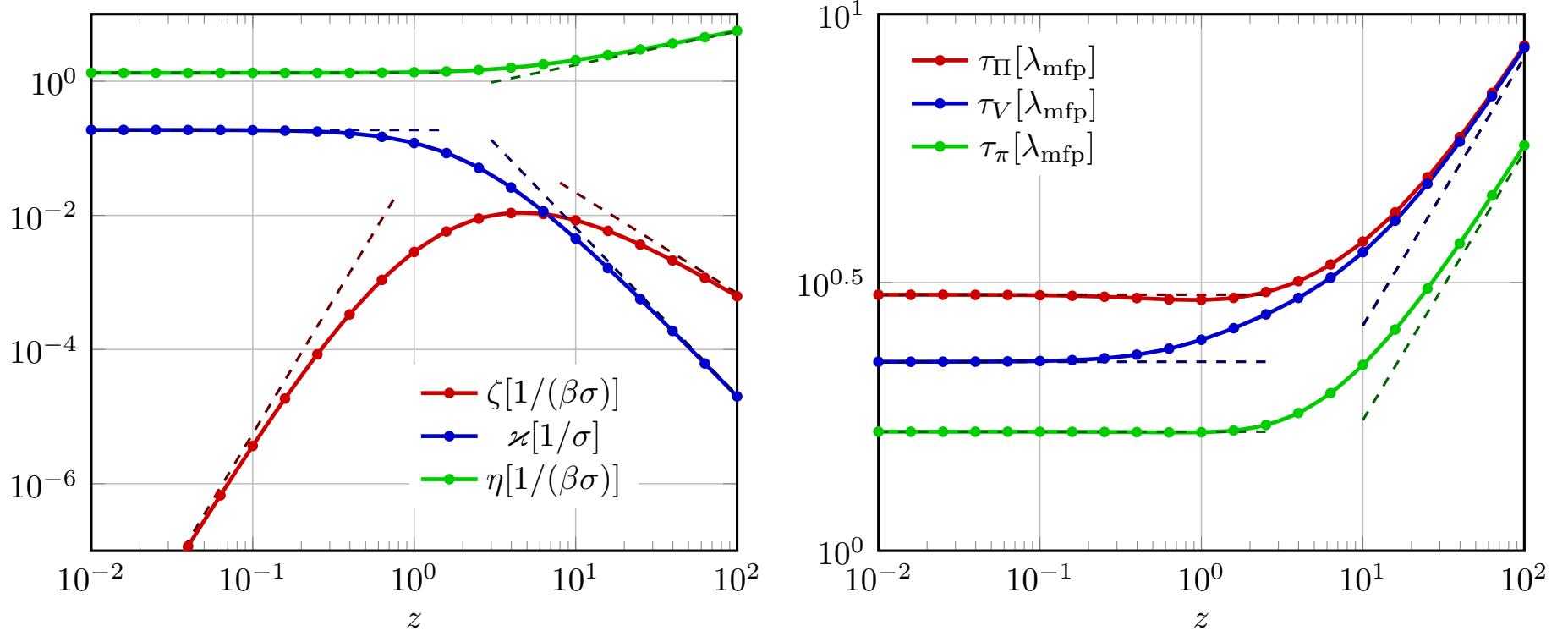
$$\tau_{\Pi} \dot{\Pi} + \Pi = -\zeta \theta + \mathcal{J} + \mathcal{R}$$

the second-order **dissipative** corrections to **bulk-viscous** pressure are

$$\begin{aligned}\mathcal{J} &\equiv -\ell_{\Pi V} \nabla_\mu V^\mu - \tau_{\Pi V} V^\mu F_\mu - \delta_{\Pi \Pi} \Pi \theta - \lambda_{\Pi V} V^\mu I_\mu + \lambda_{\Pi \pi} \pi^{\mu\nu} \sigma_{\mu\nu} , \\ \mathcal{R} &\equiv \varphi_1 \Pi^2 + \varphi_2 V^\mu V_\mu + \varphi_3 \pi^{\mu\nu} \pi_{\mu\nu} .\end{aligned}$$

The **transport coefficients** can be computed analytically as thermodynamic functions! 

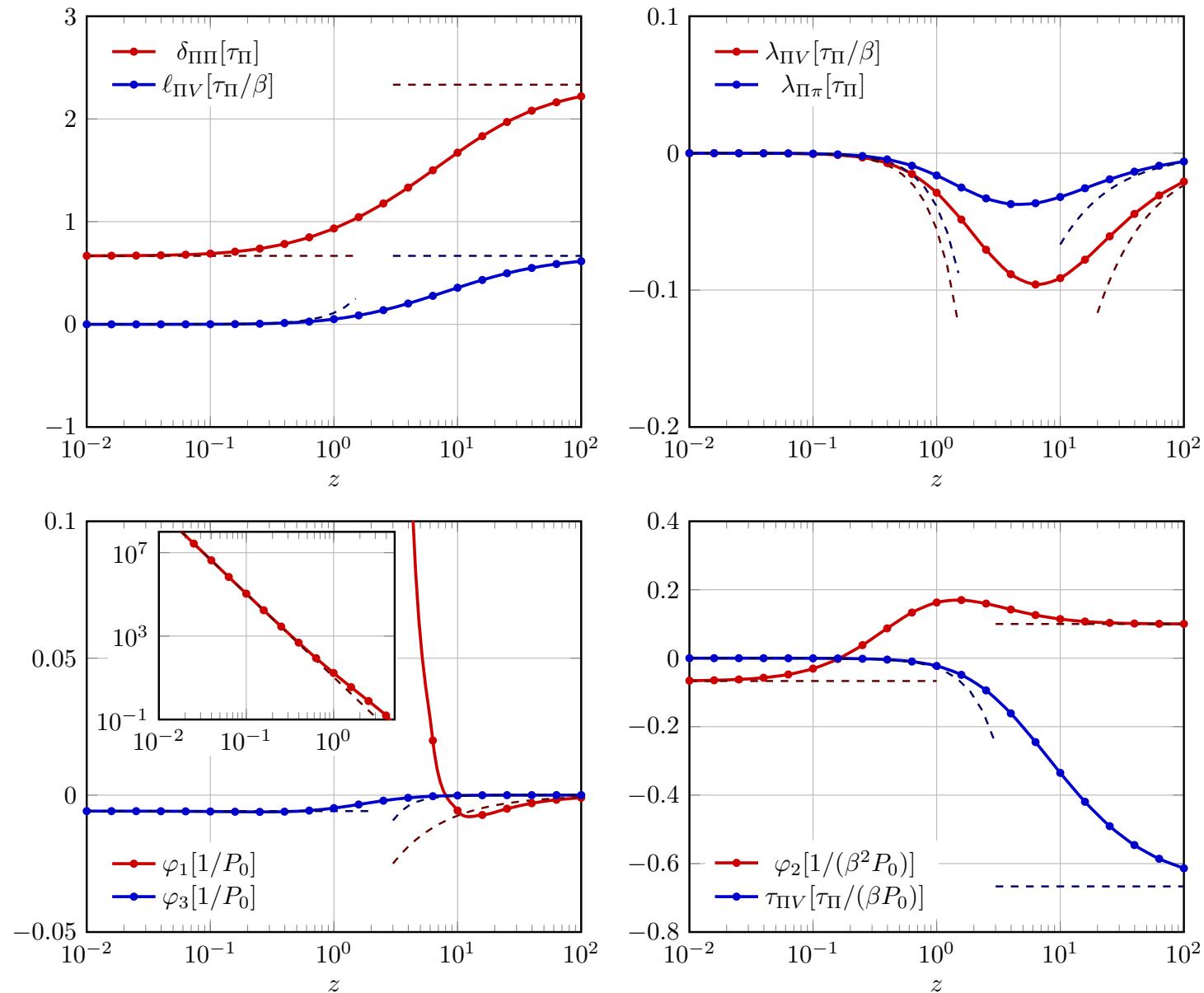
- ▶ We computed all transport coefficients of bulk-viscosity, particle-diffusion, and shear-viscosity.



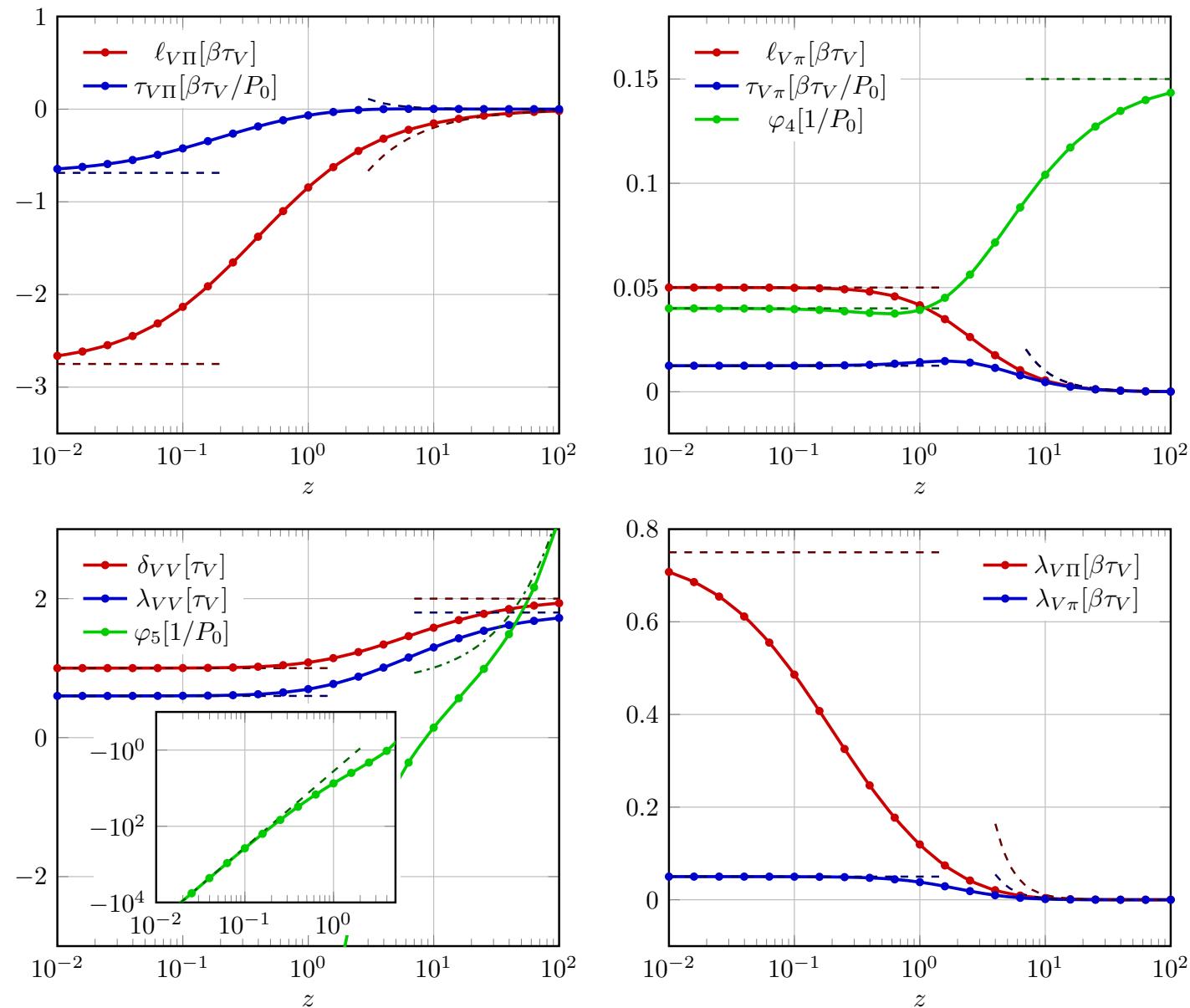
**Figure:** The first-order coefficients  $\zeta$ ,  $\kappa$ , and  $\eta$ , as well as the relaxation times  $\tau_{\Pi}$ ,  $\tau_V$ , and  $\tau_{\pi}$ .

$$z \equiv m/T$$

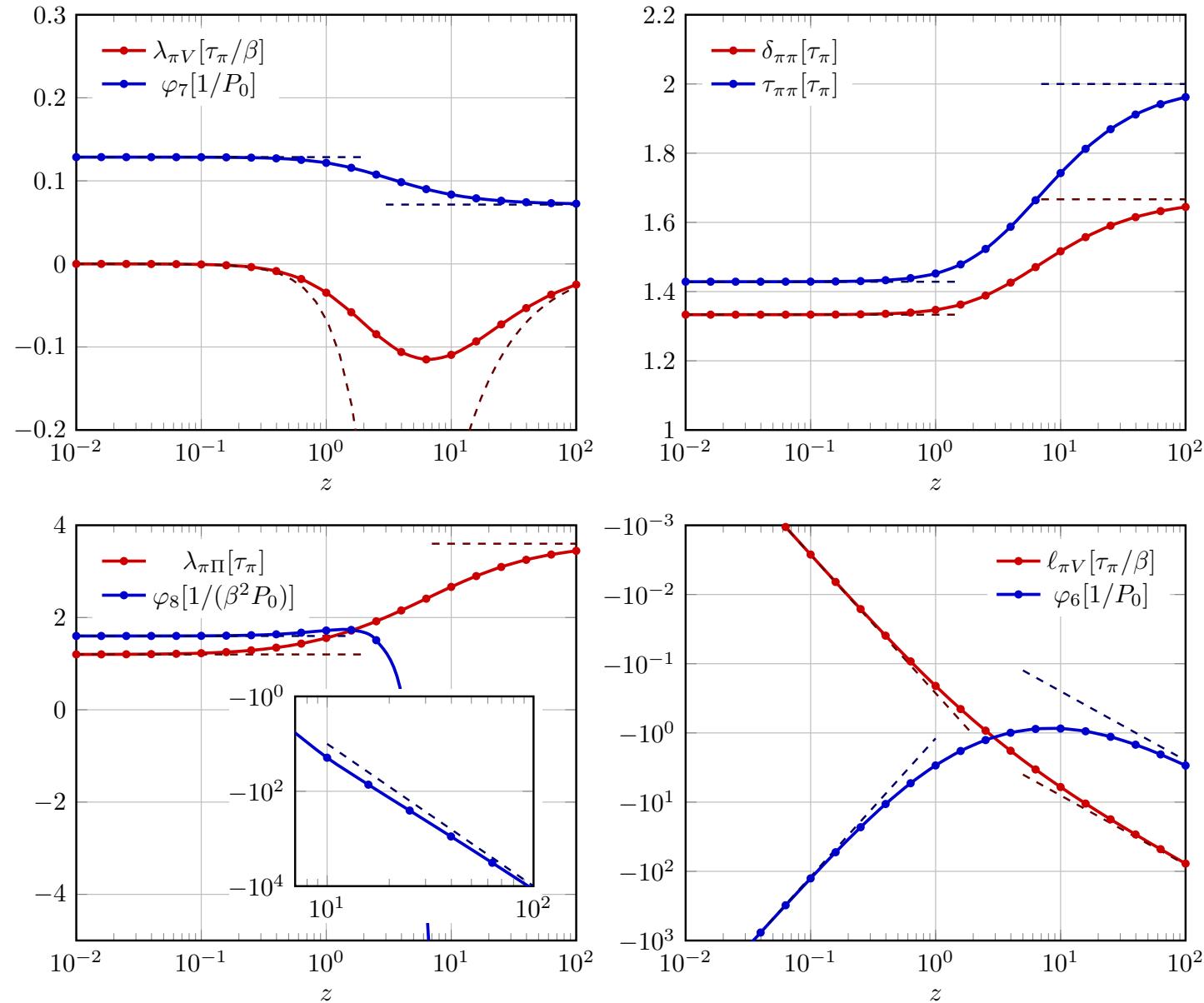
# Second-order coefficients: bulk



# Second-order coefficients: diffusion



# Second-order coefficients: shear



# Non-relativistic limit of relativistic transport equations

- ▶ Thermodynamic integral becomes

$$I_{nq} \rightsquigarrow \frac{e^{\alpha-z}}{(2\pi)^{3/2}\beta^{n+2}} z^{n-q+1/2}. \quad (19)$$

- For example, the pressure is

$$I_{21} \rightsquigarrow \frac{e^{\alpha-z}}{(2\pi)^{3/2}\beta^4} z^{3/2} \equiv P_0^{(\infty)}. \quad (20)$$

- ▶ Specific enthalpy reads

$$h_0 = mc^2 + m \left( e_0 + \frac{P_0}{\rho_0} \right) \rightsquigarrow mc^2 \equiv h_0^{(\infty)}, \quad (21)$$

where  $\rho_0 \equiv n_0 m$  the rest-mass density and  $e_0$  the kinetic energy density.

# Particle-heat correspondence

- ▶ Combining the Gibbs-Duhem relation and the Euler relation,

$$0 = s \, dT - dP + n_0 \, d\mu , \quad \epsilon = Ts - P + n_0\mu ,$$

respectively, yields

$$d\beta = h_0^{-1} \, d\alpha - \frac{\beta}{\epsilon_0 + P_0} \, dP_0 .$$

In terms of a covariant derivative, it holds

## Gibbs-Duhem relation for particle-heat correspondence

$$I_\mu - h_0 J_\mu = \nabla_\mu \ln P_0 .$$

This equation connects the **particle-change** to **heat-change** at variable pressure and temperature. !

- ▶ **Heat-flow** (in Landau frame),

$$q_\mu \equiv -h_0 V_\mu .$$

is a more natural variable than **particle-diffusion** in non-relativistic gases.

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$$h_0 \equiv (\epsilon_0 + P_0)/(\beta P_0) , \quad I_\mu \equiv \nabla_\mu \alpha , \quad J_\mu \equiv \nabla_\mu \beta$$

Grad's transport equations can be derived from the relativistic transport equations of IReD as follows :

- ▶ Rewrite the relativistic equations at  $\mathcal{O}(\text{Kn IRe}, \text{IRe}^2)$  in terms of heat variables
- ▶ Substitute the irreducible tensors  $\theta$ ,  $\sigma$ ,  $\omega$ , as well as other characteristic thermodynamic variables
- ▶ Insert the asymptotic values of the transport coefficients
- ▶ (Take the spatial part)

# Comparing asymptotic IReD to Grad: diffusion



- ▶ Comparing the transport equations for heat-flow:

- IReD (hard-spheres):

$$\begin{aligned}
 -\frac{2}{3} \frac{P_0^{(\infty)}}{\eta^{(\infty)}} q^\mu &= \Delta_\nu^\mu \frac{dq^\nu}{d\tau} - \frac{5}{2} P_0^{(\infty)} R \nabla^\mu T + q_\nu \nabla^\nu u^\mu + \frac{2}{5} q_\nu (\nabla^\nu u^\mu + \nabla^\mu u^\nu) \\
 &+ \frac{7}{5} q^\mu \nabla_\lambda u^\lambda - RT \pi^{\mu\nu} \nabla_\nu \ln P_0^{(\infty)} + RT \Delta^{\mu\nu} \nabla_\lambda \pi_\nu^\lambda \\
 &+ \frac{7}{2} R \pi^{\mu\nu} \nabla_\nu T - \underbrace{\frac{1}{10\eta^{(\infty)}} \pi^{\mu\nu} q_\nu}_{\text{interaction-specific}} ,
 \end{aligned} \tag{22}$$

- Grad (Maxwell-molecules):

$$\begin{aligned}
 -\frac{2}{3} \frac{P_0}{\eta} q^i &= \frac{dq^i}{dt} + \frac{5}{2} P_0 \frac{\partial RT}{\partial x_i} + \frac{7}{5} q^i \frac{\partial u_k}{\partial x_k} + q_k \frac{\partial u^i}{\partial x_k} + \frac{2}{5} q_k \left( \frac{\partial u^i}{\partial x_k} + \frac{\partial u^k}{\partial x_i} \right) \\
 &- RT \pi^{ik} \frac{\partial \ln P_0}{\partial x^k} + RT \frac{\partial \pi^{ik}}{\partial x^k} + \frac{7}{2} \pi^{ik} \frac{\partial RT}{\partial x^k} - \underbrace{\frac{\pi^{ik}}{\rho_0} \frac{\partial \pi_{kl}}{\partial x_l}}_{\sim \mathcal{O}(\text{KnIRe}^2)} .
 \end{aligned} \tag{23}$$

# Comparing asymptotic IReD to Grad: shear

- ▶ Comparing the transport equations for heat-flow:

- IReD (hard-spheres):

$$\begin{aligned}
 -\frac{P_0^{(\infty)}}{\eta^{(\infty)}}\pi^{\mu\nu} = & \Delta_{\alpha\beta}^{\mu\nu}\frac{d\pi^{\alpha\beta}}{d\tau} - 2P_0^{(\infty)}\nabla^{\langle\mu}u^{\nu\rangle} + \pi^{\mu\nu}\nabla_\lambda u^\lambda + 2\pi^{\lambda\langle\mu}\nabla_\lambda u^{\nu\rangle} - \frac{4}{5}\nabla^{\langle\mu}q^{\nu\rangle} \\
 & - \underbrace{\frac{1}{14\eta^{(\infty)}}\pi^{\lambda\langle\mu}\pi^{\nu\rangle}_\lambda}_{\text{interaction-specific}} + \underbrace{\frac{1}{100RT\eta^{(\infty)}}q^{\langle\mu}q^{\nu\rangle}}_{\text{interaction-specific}} .
 \end{aligned} \tag{24}$$

- Grad (Maxwell-molecules):

$$-\frac{P_0}{\eta}\pi^{ij} = \frac{d\pi^{ij}}{dt} + 2P_0\frac{\partial u^{\langle i}}{\partial x_j\rangle} + \pi^{ij}\frac{\partial u_k}{\partial x_k} + 2\pi^{r\langle i}\frac{\partial u^j\rangle}{\partial x^r} + \frac{4}{5}\frac{\partial q^{\langle i}}{\partial x_j\rangle} , \tag{25}$$

- ▶ Indeed, further away from equilibrium the geometry of gas particles becomes even more important. 

- ▶ Second-order transport coefficients were computed for arbitrary masses
  - Finally, all  $\text{IRe}^2$ -coefficients are available for arbitrary  $z$  ★
  - Many coefficients vanish for  $z \rightsquigarrow 0$  and/or  $z \rightsquigarrow \infty$ , but not in-between!
- ▶ We have shown how to derive Grad's by taking the non-relativistic limit of the transport equations
  - $\text{IRe}^2$ -terms depend on the geometry of the gas,
  - however,  $\tau_V^{(\infty)} / \tau_\pi^{(\infty)} = 3/2$  stays the same for Maxwell- and hard-sphere molecules

thank you. 

# Appendix

# DNMR vs. IReD

- ▶ Use asymptotic matching to express all irreducible moments through **dissipative quantities** and **fluid-dynamical gradients**
- ▶ Discard h.o.t

## Hydrodynamic transport equations (DNMR)

$$\tau_{\Pi} \dot{\Pi} + \Pi = -\zeta_0 \theta + \mathcal{J} + \mathcal{K} + \mathcal{R} \quad (26)$$

$$\tau_V \dot{V}^{\langle\mu\rangle} + V^\mu = \kappa_0 V^\mu + \mathcal{J}^\mu + \mathcal{K}^\mu + \mathcal{R}^\mu \quad (27)$$

$$\tau_\pi \dot{\pi}^{\langle\mu\nu\rangle} + \pi^{\mu\nu} = 2\eta_0 \sigma^{\mu\nu} + \mathcal{J}^{\mu\nu} + \mathcal{K}^{\mu\nu} + \mathcal{R}^{\mu\nu} \quad (28)$$

- ▶ First-order contributions  $\sim \mathcal{O}(\text{IRe})$  and  $\sim \mathcal{O}(\text{Kn})$
- ▶ Second-order contributions  $\sim \mathcal{O}(\text{Kn IRe})$  and  $\sim \mathcal{O}(\text{Kn}^2, \text{IRe}^2)$
- ▶ Contributions of order  $\mathcal{O}(\text{Kn}^2)$  result directly from asymptotic matching
  - Example:  $\theta \rho_r \rightarrow \theta \Pi, \theta^2$

- ▶ Consider the second-order terms of tensor-rank two:

$$\begin{aligned} \mathcal{K}^{\mu\nu} = & \tilde{\eta}_1 \omega^\lambda \langle \mu \omega^\nu \rangle_\lambda + \tilde{\eta}_2 \theta \sigma^{\mu\nu} + \tilde{\eta}_3 \sigma^\lambda \langle \mu \sigma^\nu_\lambda \rangle + \tilde{\eta}_4 \sigma^\lambda_\lambda \langle \mu \omega^\nu \rangle^\lambda + \tilde{\eta}_5 I^\lambda \langle \mu I^\nu \rangle \\ & + \tilde{\eta}_6 F^\lambda \langle \mu F^\nu \rangle + \tilde{\eta}_7 I^\lambda \langle \mu F^\nu \rangle + \tilde{\eta}_8 \nabla^\lambda \langle \mu I^\nu \rangle + \tilde{\eta}_9 \nabla^\lambda \langle \mu F^\nu \rangle \end{aligned} \quad (29)$$

- ▶ Second derivatives of fluid-dynamical quantities appear
  - Equations become **parabolic!**
  - Theory becomes acausal and unstable 🕯
- ▶ Conventional remedy: **Ignore** terms of order  $\mathcal{O}(Kn^2)$  (tDNMR)
  - Equations are hyperbolic again (but ad hoc)
- ▶ Is there a way to ensure  $\mathcal{K} = \mathcal{K}^\mu = \mathcal{K}^{\mu\nu} = 0$  from the beginning?

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$$F^\mu \equiv \nabla^\mu P_0, \quad \omega^{\mu\nu} \equiv (\nabla^\mu u^\nu - \nabla^\nu u^\mu)/2$$



DW, A. Palermo, V. E. Ambruş, arXiv:2203.12608

- ▶ General idea: Relate moments through their Navier-Stokes solutions

## IReD: Asymptotic matching

$$\rho_r = -\zeta_r \theta + \mathcal{O}(\text{Kn IRe}) \quad \mapsto \quad \rho_r = \frac{\zeta_r}{\zeta_n} \rho_n + \mathcal{O}(\text{Kn IRe}) \quad (30)$$

$$\rho_r^\mu = \kappa_r I^\mu + \mathcal{O}(\text{Kn IRe}) \quad \mapsto \quad \rho_r^\mu = \frac{\kappa_r}{\kappa_n} \rho_V^\mu + \mathcal{O}(\text{Kn IRe}) \quad (31)$$

$$\rho_r^{\mu\nu} = 2\eta_r \sigma^{\mu\nu} + \mathcal{O}(\text{Kn IRe}) \quad \mapsto \quad \rho_r^{\mu\nu} = \frac{\eta_r}{\eta_n} \rho_n^{\mu\nu} + \mathcal{O}(\text{Kn IRe}) \quad (32)$$

- ▶ ! No terms  $\sim \mathcal{O}(\text{Kn})$  appear in asymptotic matching (IRe dominance 😊)
- ▶ Equations of motion can be closed in terms of any set of moments  $\rho_n, \rho_n^\mu, \rho_n^{\mu\nu}$
- ▶ Choose  $n = 0$  to obtain closure in terms of hydrodynamic quantities

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Also known as “order-of-magnitude approximation” J. A. Fotakis, E. Molnár, H. Niemi, C. Greiner, D. H. Rischke

arXiv: 2203.11549

- ▶ Procedure analogous: use new asymptotic matching conditions to express all irreducible moments through **dissipative quantities** and **fluid-dynamical gradients**
- ▶ Diagonalize and discard **h.o.t**

## Hydrodynamic transport equations (IReD)

$$\tau_{\Pi} \dot{\Pi} + \Pi = -\zeta_0 \theta + \mathcal{J} + \mathcal{R} \quad (33a)$$

$$\tau_V \dot{V}^{\langle\mu\rangle} + V^\mu = \kappa_0 V^\mu + \mathcal{J}^\mu + \mathcal{R}^\mu \quad (33b)$$

$$\tau_\pi \dot{\pi}^{\langle\mu\nu\rangle} + \pi^{\mu\nu} = 2\eta_0 \sigma^{\mu\nu} + \mathcal{J}^{\mu\nu} + \mathcal{R}^{\mu\nu} \quad (33c)$$

- ▶ Structure is similar, but transport coefficients different for  $N_0 > 2$ ,  $N_1 > 1$ ,  $N_2 > 0$  ( $> 14$  moments)
- ▶ Only terms  $\sim \mathcal{O}(\text{IRe})$ ,  $\sim \mathcal{O}(\text{Kn})$ ,  $\sim \mathcal{O}(\text{Kn IRe})$ ,  $\sim \mathcal{O}(\text{IRe}^2)$  appear  
 → Equations stay **hyperbolic**, no need to discard terms
- ▶ Absence of parabolic terms due to modified asymptotic matching

# A different path to hyperbolicity



- ▶ Basic idea of IReD and DNMR: Relate quantities up to order  $\mathcal{O}(\text{Kn IRe})$
- ▶ 🤔 Ambiguities in second-order terms since to first order

$$\Pi \simeq -\zeta \theta, V^\mu \simeq \kappa_0 I^\mu, \pi^{\mu\nu} \simeq 2\eta_0 \sigma^{\mu\nu} \quad (34)$$

- ▶ For example:

$$\langle \mathcal{K} \rangle \ni \theta^2 = -\frac{\Pi \theta}{\zeta_0} \in \langle \mathcal{J} + \mathcal{O}(\text{Kn IRe}) \rangle ,$$

- ▶ 💡 IReD way: “Trade one power of Kn for one power of IRe”
- ▶ To eliminate the parabolic terms:
  1. Start with the DNMR approach
  2. Use prescription to absorb coefficients in  $\mathcal{K}, \mathcal{K}^\mu, \mathcal{K}^{\mu\nu}$  into  $\mathcal{J}, \mathcal{J}^\mu, \mathcal{J}^{\mu\nu}$
- ▶ Allows to relate transport coefficients in the two approaches
- ▶ **IReD and DNMR equivalent up to (and including) order  $\mathcal{O}(\text{Kn}^2, \text{Kn IRe}, \text{IRe}^2)$**

# Non-relativistic gas dynamics

The integral

$$I_{nq} = \frac{(-1)^q}{(2q+1)!!} \int dK f_{0\mathbf{k}} E_{\mathbf{k}}^{n-2q} (m^2 - E_{\mathbf{k}}^2)^q ,$$

can be rewritten for  $|\mathbf{k}| < m$  to

$$\begin{aligned} I_{nq} = & \frac{e^{\alpha-z}}{(2\pi)^{3/2} \beta^{n+2}} z^{n-q+1/2} \sum_{\ell=0}^{\infty} \binom{\frac{n-1}{2} - q}{\ell} z^{-\ell} \frac{(2\ell + 2q + 1)!!}{(2q+1)!!} \\ & \times \left\{ 1 + \sum_{j=1}^{\infty} \frac{z^{-j}}{4^j j!} \sum_{k=1}^j (-1)^{j+k} \frac{(2j + 2k + 2\ell + 2q + 1)!!}{2^k (2\ell + 2q + 1)!!} \right. \\ & \quad \left. \times B_{j,k} \left[ 1, \dots, \frac{(2j - 2k + 2)!}{(j - k + 2)!} \right] \right\} , \end{aligned} \quad (35)$$

where  $B_{j,k}$  denotes the incomplete exponential Bell polynomials .

# Brief detour on special functions: Bell theory

I will follow a nice exposition of Bell polynomials by [Nicholas Wheeler](#). *Bell Polynomials & Related Constructs*. 2020. URL:

<https://www.reed.edu/physics/faculty/wheeler/documents/Miscellaneous%20Math/Bell%20Polynomials,%20Pochhammer%20Symbols,%20Etc/Bell%20Polynomials.pdf>.

# Theory of Bell polynomials

For example, we want to expand the composite function

$$F(x) = f(g(x)) ,$$

for a generic  $g(x) \equiv \sum_k a_k x^k$ . Then,

$$f(x) \equiv \frac{1}{1-x} \implies F(x) = \sum_k D_k x^k ,$$

with the new coefficients determined by the Toeplitz matrix

$$D_n = \det \begin{pmatrix} a_1 & a_2 & \cdots & \cdots & a_{n-1} & a_n \\ -1 & a_1 & a_2 & \ddots & \ddots & a_{n-1} \\ 0 & -1 & a_1 & a_2 & \ddots & \vdots \\ 0 & 0 & -1 & a_1 & \ddots & \vdots \\ \vdots & \ddots & \ddots & \ddots & \ddots & a_2 \\ 0 & \dots & \dots & \dots & -1 & a_1 \end{pmatrix} = \sum_{k=1}^n a_k D_{n-k} .$$

Recursion formula for coefficients. 😎

# Theory of Bell polynomials

Suppose now  $F(x) = \exp(g(x))$ . Then,  $F(x) = \sum_k E_k x^k$ , where

$$E_n = \frac{1}{n!} \det \begin{pmatrix} a_1 & 2a_2 & \cdots & \cdots & (n-1)a_{n-1} & na_n \\ -1 & a_1 & 2a_2 & \ddots & \ddots & (n-1)a_{n-1} \\ 0 & -2 & a_1 & 2a_2 & \ddots & \vdots \\ 0 & 0 & -3 & a_1 & \ddots & \vdots \\ \vdots & \ddots & \ddots & \ddots & \ddots & 2a_2 \\ 0 & \cdots & \cdots & \cdots & -n+1 & a_1 \end{pmatrix}$$
$$= \sum_{k=1}^n \frac{(n-1)! k}{(n-k)!} a_k E_{n-k} .$$

Arising connection to binomial coefficient? 🤔

# Theory of Bell polynomials

Finally, suppose the generating function is  $F(x) = \exp\left(\sum_k \frac{b^k}{k!} x_k\right)$ . Then,

$$B_n(b_1, \dots, b_n) = \det \begin{pmatrix} b_1 & \frac{1}{1!} b_2 & \cdots & \cdots & \frac{1}{(n-2)!} b_{n-1} & \frac{1}{(n-1)!} b_n \\ -1 & b_1 & \frac{1}{1!} b_2 & \ddots & \ddots & (n-2)! b_{n-1} \\ 0 & -2 & b_1 & \frac{1}{1!} b_2 & \ddots & \ddots \\ 0 & 0 & -3 & b_1 & \ddots & \ddots \\ \vdots & \ddots & \ddots & \ddots & \ddots & \frac{1}{1!} b_2 \\ 0 & \cdots & \cdots & \cdots & -n+1 & b_1 \end{pmatrix}.$$

Indeed, we get a binomial coefficient:

$$B_{n+1}(b_1, \dots, b_{n+1}) = \sum_{m=0}^n \binom{n}{m} b_{m+1} B_{n-m}(b_1, \dots, b_{n-m}).$$

where  $B_n(b_1, \dots, b_n)$  are the so-called Bell polynomials. 😊

To recap, the Bell polynomials, generated by

$$\exp \left( \sum_{k=1}^{\infty} x_k \frac{t^k}{k!} \right) = \sum_{n=0}^{\infty} B_n (x_1, \dots, x_n) \frac{t^n}{n!} ,$$

are actually multinomials. The Bell numbers,

$$B_n \equiv B_n (1, \dots, 1) \leftarrow B_n (x_1, \dots, x_n) ,$$

count set partitions. For example,  $\{a, b, c\}$  can be partitioned ( $B_3 = 5$ )-many ways:

$$\begin{array}{ll} \{a\}, \{b\}, \{c\} & \{a, b, c\}, \{\}, \{\} \\ \{\}, \{b\}, \{a, c\} & \{a, b\}, \{\}, \{c\} \\ \{a\}, \{b, c\}, \{\} . \end{array}$$

$$\underbrace{B_j(x_1, \dots, x_j)}_{\text{Complete}} = \sum_{k=1}^j \underbrace{B_{j,k}(x_1, \dots, x_{j-k+1})}_{\text{Incomplete}} . \quad (36)$$

Incomplete Bell polynomials are homogeneous functions,

$$B_{j,k}(\gamma\delta x_1, \dots, \gamma\delta^{j-k+1} x_{j-k+1}) = \gamma^k \delta^j B_{j,k}(x_1, \dots, x_{j-k+1}) . \quad (37)$$

Bell polynomials are connected to Hermite polynomials by

$$\sum_{n=0}^{\infty} \textcolor{red}{B_n(x, -1, 0, \dots, 0)} \frac{t^n}{n!} = \exp\left(xt - \frac{1}{2}t^2\right) = \sum_{n=0}^{\infty} \textcolor{red}{\text{He}_n(x)} \frac{t^n}{n!} .$$

Notably, H. Grad used Hermite polynomials to construct moment equations.

Combinatorial nature of polynomials gives rise to recursive structures like

1					
1	2				
2	3	5			
5	7	10	15		
15	20	27	37	52	

