

# Canonical Quantization of Brownian Motion and Quantum Thermodynamics

Tomoi Koide (IF UFRJ)

(collaboration with Fernando Nicacio (IF UFRJ))

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# Table of contents

1. Introduction (Stochastic and Quantum thermodynamics)
2. Generalized Brownian Motion
3. Canonical quantization and new quantum master equation
4. Other topics
5. Concluding remarks

# Thermodynamics

1. Systems involve macroscopic degrees of freedom ( thermodynamical limit, small fluctuation )
2. Initial and final states of processes are in equilibrium.

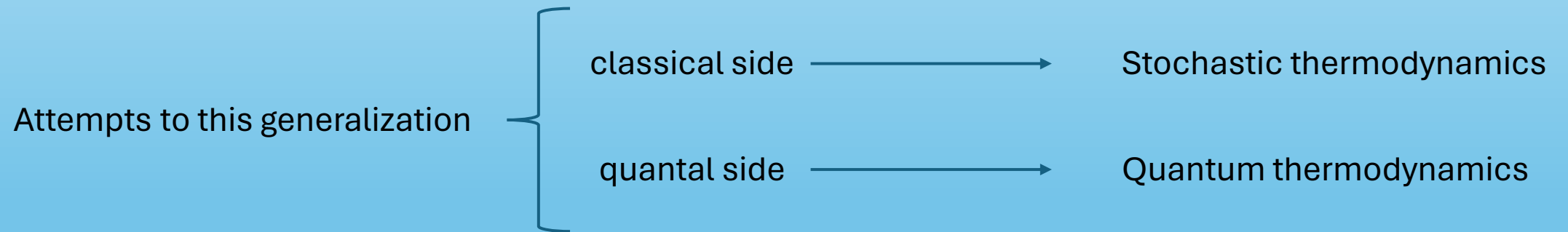
# Thermodynamics

1. Systems involve macroscopic degrees of freedom ( thermodynamical limit, small fluctuation )

 Small systems

2. Initial and final states of processes are in equilibrium.

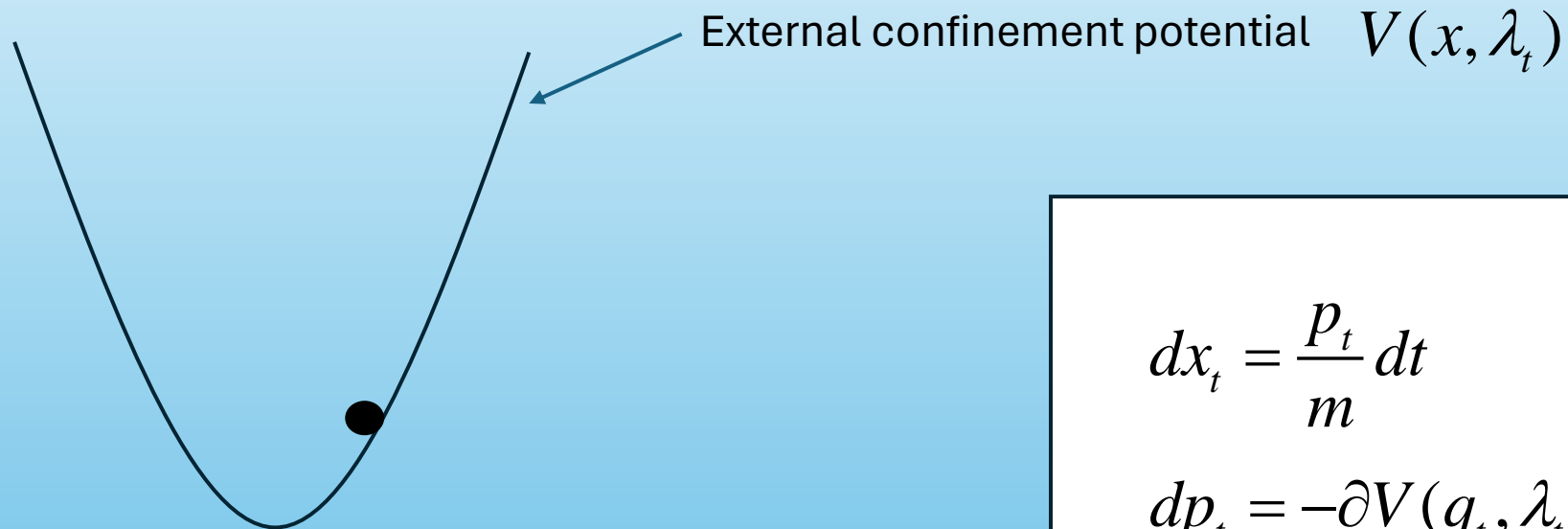
 Time evolutions from arbitrary states



These formulations are **yet under construction**.

# Stochastic Thermodynamics

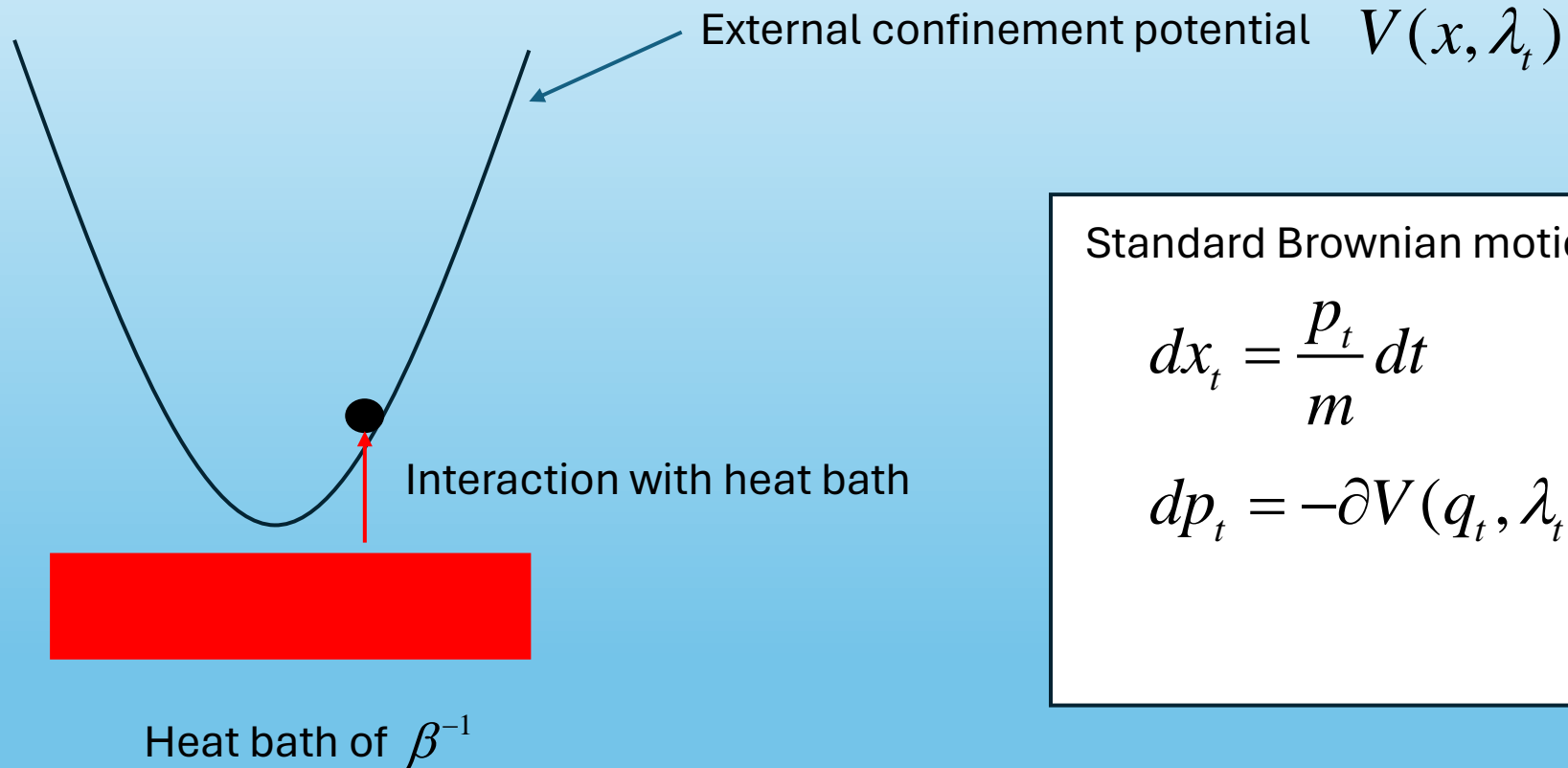
# Stochastic thermodynamics (stochastic energetics)



$$dx_t = \frac{p_t}{m} dt$$

$$dp_t = -\partial V(q_t, \lambda_t) dt$$

# Stochastic thermodynamics (stochastic energetics)

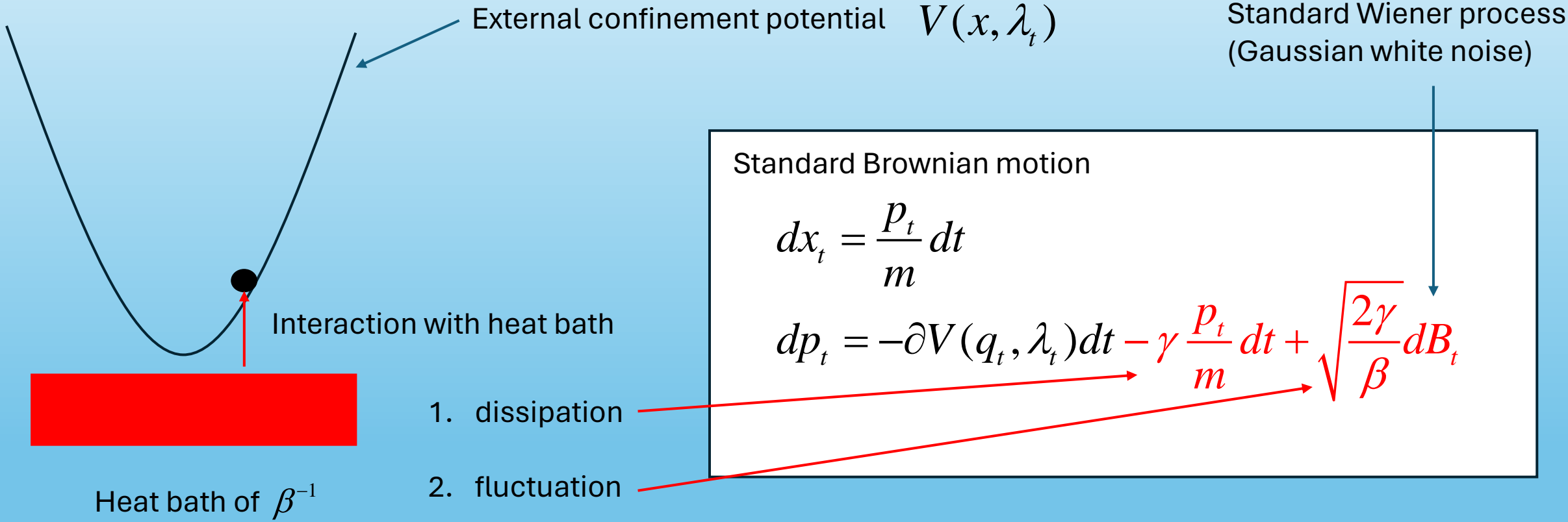


Standard Brownian motion

$$dx_t = \frac{p_t}{m} dt$$

$$dp_t = -\partial V(q_t, \lambda_t) dt$$

# Stochastic thermodynamics (stochastic energetics)





# Stochastic thermodynamics (stochastic energetics)

Sekimoto, “Stochastic Energetics” (Springer, 2010)

**Heat** absorbed by the system is interpreted as a **work** done by the heat bath on the system.

$$dQ_t^c = \int d\Gamma_0 f_0(\Gamma_0) \mathbf{E} \left[ \left( -\gamma \frac{p_t}{m} dt + \sqrt{\frac{2\gamma}{\beta}} \right) \circ dx_t \right]$$

$\mathbf{E}[\dots]$  : ensemble average for thermal fluctuation (Wiener)

$$d\Gamma_0 = dq_0 dp_0$$

$f_0(\Gamma_0)$  : initial phase space distribution

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Energy

$$U_t^c = \int d\Gamma_0 f_0(\Gamma_0) \mathbf{E} \left[ \frac{p_t^2}{2m} + V(x_t, \lambda_t) \right]$$

Work by heat bath

$$dW_t^c = \int d\Gamma_0 f_0(\Gamma_0) \mathbf{E} \left[ d\tilde{W}_t^c \right] \quad d\tilde{W}_t^c := \frac{\partial V(q_t, \lambda_t)}{\partial \lambda_t} d\lambda_t$$

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First law

$$U_{t_f}^c - U_{t_i}^c = dQ_t^c + dW_t^c$$

Energy

$$U_t^c = \int d\Gamma_0 f_0(\Gamma_0) \mathbf{E} \left[ \frac{p_t^2}{2m} + V(x_t, \lambda_t) \right]$$

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# Stochastic thermodynamics (stochastic energetics)

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Information entropy

$$S_c = -k_B \int d\Gamma f(\Gamma, t) \ln f(\Gamma, t)$$

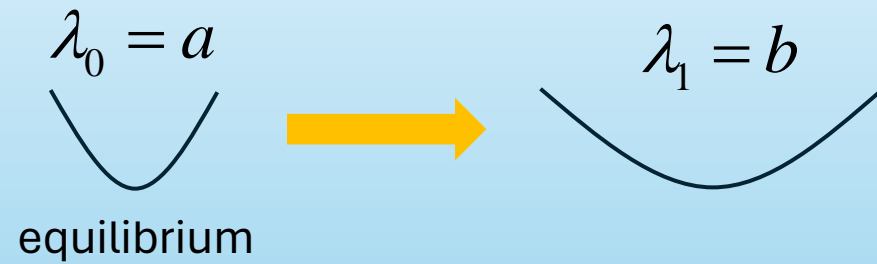
Phase space distribution

Second law

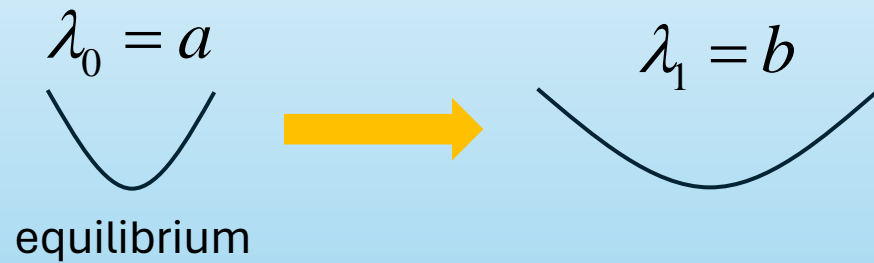
$$\frac{dS_c}{dt} - k_B \beta \frac{dQ_t^c}{dt} \geq 0$$

The equality is satisfied for the equilibrium distribution,  $f_{eq} = \frac{1}{Z} e^{-\beta \left( \frac{p^2}{2m} + V(x, \lambda_\infty) \right)}$

# Jarzynski equality



# Jarzynski equality

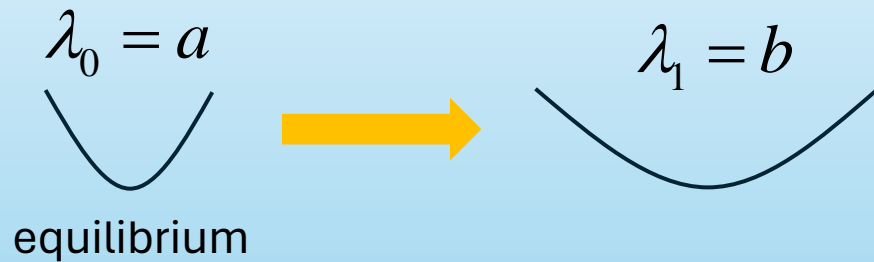


Let us consider the external perturbation characterized by  $\lambda_0 = a \longrightarrow \lambda_1 = b$ . Then the averaged work in this process is

$$W_{0 \rightarrow 1}^c = \int d\Gamma_0 f_{eq}(\Gamma_0) \mathbb{E} \left[ \int_0^1 d\tilde{W}_t^c \right]$$

Then, for the above process, we can show

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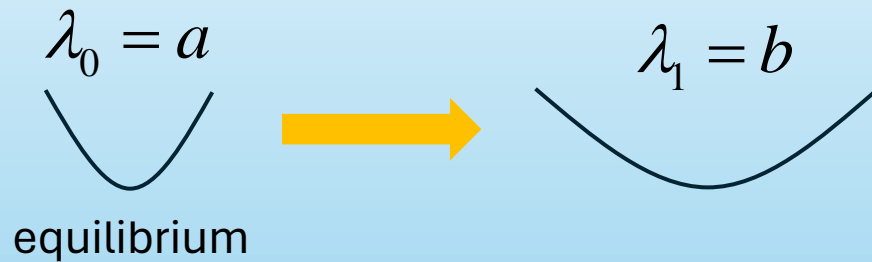
Jarzynski equality

$$\int d\Gamma_0 f_{eq}(\Gamma_0) \mathbb{E} \left[ \exp \left( -\beta \left( \int_0^1 d\tilde{W}_t^c - \Delta F \right) \right) \right] = 1$$

$$\Delta F = \beta^{-1} (-\ln Z_1 + \ln Z_0) \quad Z_t = \int d\Gamma e^{-\beta \left( \frac{p^2}{2m} + V(x, \lambda_t) \right)}$$

Jarzynski equality is one of fluctuation theorems.

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second law

$$-W_{0 \rightarrow 1}^c \leq -\Delta F$$

$$\Delta F = \beta^{-1} (-\ln Z_1 + \ln Z_0) \quad Z_t = \int d\Gamma e^{-\beta \left( \frac{p^2}{2m} + V(x, \lambda_t) \right)}$$

Jarzynski equality is one of fluctuation theorems.

There always exist **stochastic events** which change in an opposite direction to the mean behavior of entropy.



Of course, stochastic thermodynamics is **not** applicable to extremely small systems where quantum fluctuation should be considered.



How do we introduce a quantum dissipative model which is thermodynamically consistent?

# Quantum Thermodynamics

# CPTP map

What are the requirements for the density matrix in open quantum systems (system + environment)?

Dynamical map (time evolution)

$$\hat{\rho} \rightarrow M[\hat{\rho}]$$



1. Linear time evolution

$$M[a\hat{\rho}_1 + b\hat{\rho}_2] = aM[\hat{\rho}_1] + bM[\hat{\rho}_2]$$

2. Completely positive

$$\hat{\rho}_{AB} \geq 0 \longrightarrow M_A \otimes I_B[\hat{\rho}_{AB}] \geq 0$$

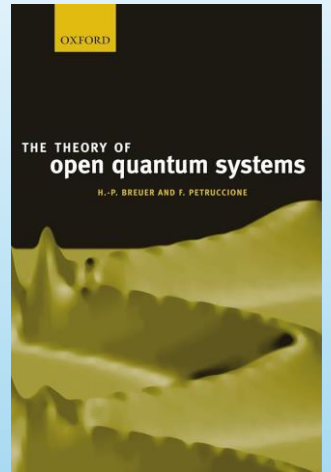
3. Trace conservation

$$\text{Tr}[\hat{\rho}] = \text{Tr}[M[\hat{\rho}]]$$

The time evolutions satisfying these conditions are called completely positive and trace-preserving (CPTP) maps. We require that open quantum dynamics is described by the CPTP evolution.

# How do we obtain non-eq. dynamics?

**There is no systematic method to obtain non-equilibrium dynamics consistent with the CPTP map for arbitrary Hamiltonian.**

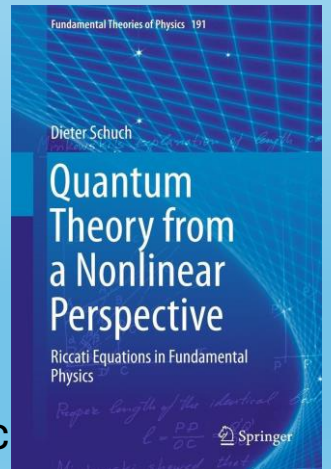


Systematic coarse-graining of environment degrees of freedom

1. Projection operator method (Nakajima-Zwanzig, Mori, Kawasaki-Gunton, Shibata-Hashitsume, etc)
2. Coarse-graining based on path integrals (influence functional, closed time path, etc)

Quantization of classical dissipative system

1. Canonical quantization (Caldirola, Kanai, Bateman, etc )
2. Non-linear Schrödinger equation through the Ehrenfest theorem (Kostin, Hasse, Schuch, etc)



Non-Hermitian (PT-symmetric) quantum mechanics

# GKSL equation

In the standard discussions of quantum thermodynamics, we often employ the equation proposed by Lindblad (1976) and Gorini-Kossakowski-Sudarshan (1976).

$$\frac{d}{dt} \hat{\rho} = -\frac{i}{\hbar} [\hat{H}, \hat{\rho}] + D[\hat{\rho}]$$

irreversible current

$$D[\hat{\rho}] = \sum_i \gamma_i \left\{ -\frac{1}{2} [\hat{L}_i^\dagger \hat{L}_i, \hat{\rho}] + \hat{L}_i \hat{\rho} \hat{L}_i^\dagger \right\}$$

reversible current

$\gamma_i \geq 0$  : Dissipative coefficients  
 $\hat{L}_i$  : jump (Lindblad) operator

1. The Gorini-Kossakowski-Sudarshan-Lindblad (GKSL) equation is **an example of the CPTP evolution**.
2. There is **no systematic method** to define the Lindblad operator.
3. It is **not clear** whether this is applicable to describe thermal relaxation processes.

# Application to harmonic oscillator

For the harmonic oscillator Hamiltonian, we can find the Lindblad operator which is consistent with thermodynamics

$$H = \hbar\omega \left( \hat{a}^\dagger \hat{a} + \frac{1}{2} \right)$$

$$\hat{L}_+ = \hat{a} = \hat{L}_-^\dagger$$

$$\hat{L}_- = \hat{a}^\dagger = \hat{L}_+^\dagger$$

$$\frac{\gamma_-}{\gamma_+} = e^{-\beta\hbar\omega}$$

Detailed balance condition

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Quantum heat

$$dQ_t^q := \text{Tr} \left[ \hat{H} d\hat{\rho}_t \right]$$

von Neumann entropy

$$S_t^q := -k_B \text{Tr} \left[ \hat{\rho}_t \ln \hat{\rho}_t \right]$$

$$\frac{dS_t^q}{dt} - k_B \beta^{-1} \frac{dQ_t^q}{dt} \geq 0$$

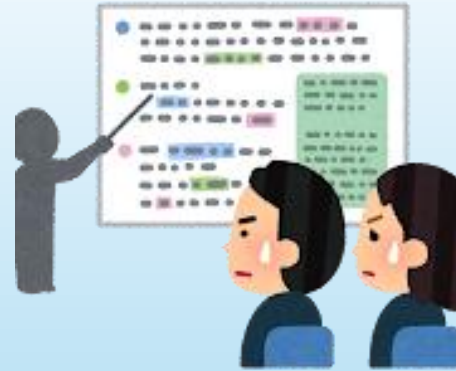
The GKSL equation is consistent with thermodynamics at least in the application to the harmonic oscillator.



We want to find a systematic procedure to obtain open quantum dynamics which is consistent with CPTP and describe thermal relaxation processes.



Our strategy is.....



Brownian motion



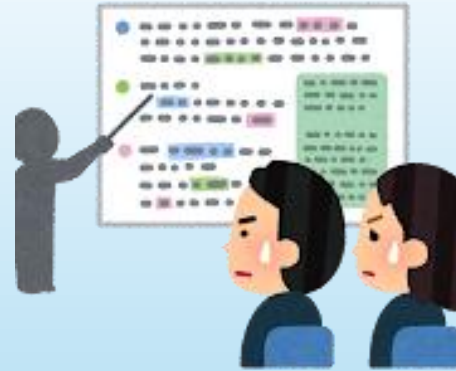
Stochastic thermodynamics

CPTP evolution



Quantum thermodynamics

Our strategy is.....



More general theory of Brownian motion (but in flat spacetime)

1) propose



Brownian motion



Stochastic thermodynamics

CPTP evolution

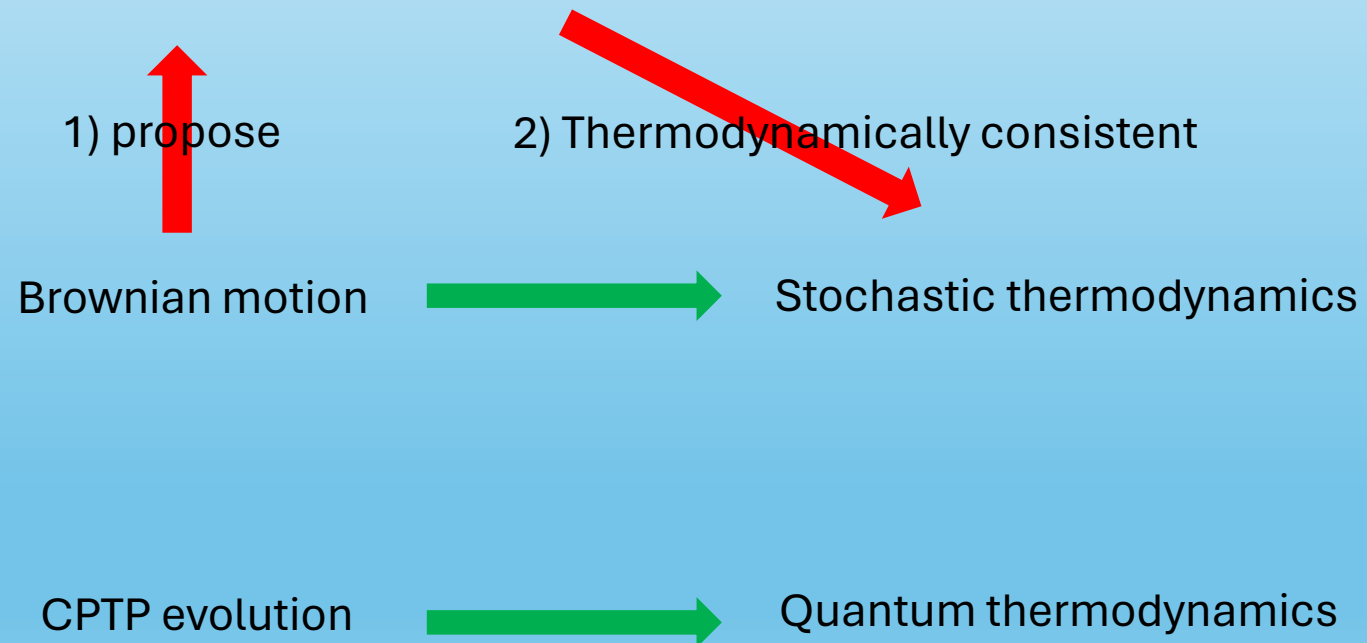


Quantum thermodynamics

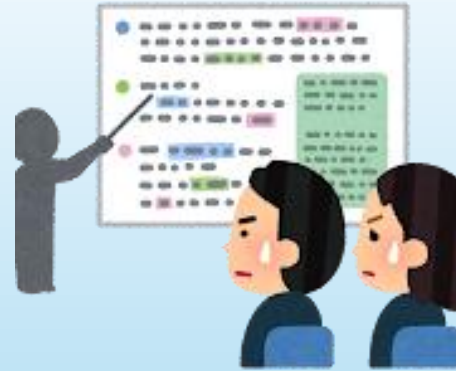
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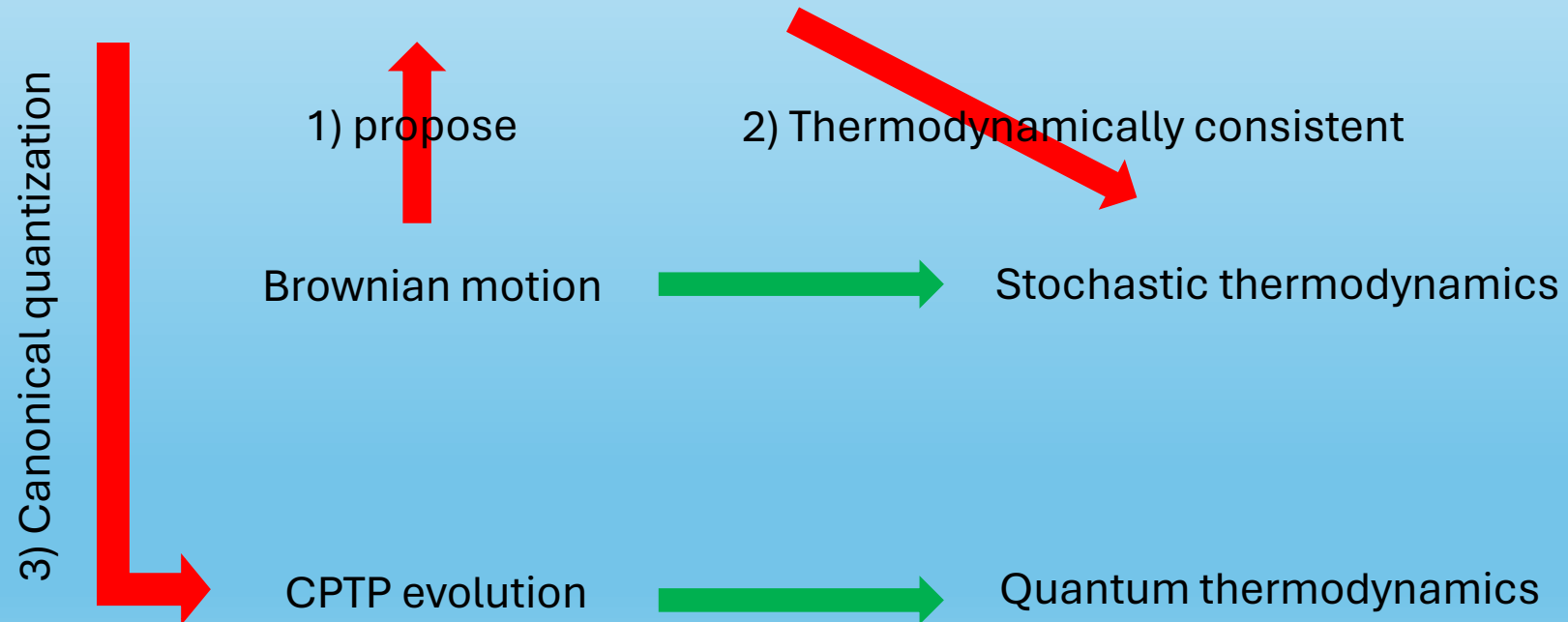
More general theory of Brownian motion (but in flat spacetime)



Our strategy is.....



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# Generalized Brownian Motion

# Generalized Brownian motion

Koide&Nicacio, Phys. Lett. A494, 129277 (2024).

Let us consider a thermal relaxation process with a general Hamiltonian  $H$ .

Our new model of Brownian motion for the  $i$ -th particle

$$d\vec{q}_{(i)t} = \frac{\partial H}{\partial \vec{p}_{(i)t}} dt$$

$$d\vec{p}_{(i)t} = -\frac{\partial H}{\partial \vec{q}_{(i)t}} dt$$

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$$d\vec{p}_{(i)t} = -\frac{\partial H}{\partial \vec{q}_{(i)t}} dt - \gamma_{q_i} \frac{\partial H}{\partial \vec{p}_{(i)t}} dt + \sqrt{\frac{2\gamma_{p_i}}{\beta_i}} d\vec{B}_{p(i)t}$$

$$\mathbb{E} \left[ dB_{q(i)t}^\alpha \right] = \mathbb{E} \left[ dB_{p(i)t}^\alpha \right] = 0$$

$$\mathbb{E} \left[ dB_{q(i)t}^\alpha dB_{q(j)t'}^\beta \right] = \mathbb{E} \left[ dB_{p(i)t}^\alpha dB_{p(j)t'}^\beta \right] = dt \delta_{ij} \delta_{\alpha\beta} \delta_{tt'}$$

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# Thermodynamical quantities

Koide&Nicacio, Phys. Lett. A494, 129277 (2024).

Energy 
$$U_t^c = \int d\Gamma_0 f_0(\Gamma_0) \mathbb{E} [H(\Gamma_t, \lambda_t)]$$

Work 
$$dW_t^c = \int d\Gamma_0 f_0(\Gamma_0) \mathbb{E} [d\tilde{W}_t^c] \quad d\tilde{W}_t^c := \frac{\partial H(\Gamma_t, \lambda_t)}{\partial \lambda_t} d\lambda_t$$

Heat 
$$dQ_{(i)t}^c = \int d\Gamma_0 f_0(\Gamma_0) \sum_{\alpha} \mathbb{E} \left[ \left( -\gamma_{p_i} \frac{\partial H}{\partial p_{(i)t}^{\alpha}} dt + \sqrt{\frac{2\gamma_{p_i}}{\beta_i}} dB_{p(i)t}^{\alpha} \right) \circ dq_{(i)t}^{\alpha} \right]$$

$$- \int d\Gamma_0 f_0(\Gamma_0) \sum_{\alpha} \mathbb{E} \left[ \left( -\gamma_{q_i} \frac{\partial H}{\partial q_{(i)t}^{\alpha}} dt + \sqrt{\frac{2\gamma_{q_i}}{\beta_i}} dB_{q(i)t}^{\alpha} \right) \circ dp_{(i)t}^{\alpha} \right]$$

1. These are reduced to standard quantities by using  $H = \frac{p^2}{2m} + V$  and  $\gamma_{q_i} = 0$ .

2. We can choose even interacting and relativistic Hamiltonians.

Phys. Rev. E83,061111 (2011), J. Phys. Commun. 2, 021001 (2018)

# Stochastic energetics

Koide&Nicacio, Phys. Lett. A494, 129277 (2024).

First law 
$$U_{t+dt}^c - U_t^c = \sum_i dQ_{(i)t}^c + dW_t^c$$

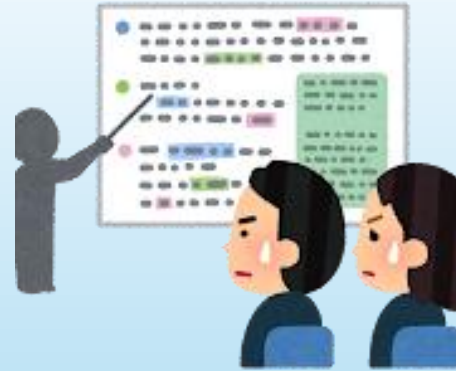
Second law 
$$\frac{dS_t^c}{dt} - \sum_{i=1}^N k_B \beta_i^{-1} \frac{dQ_{(i)t}^c}{dt} \geq 0 \quad S_t^c = -k_B \int d\Gamma f(\Gamma, t) \ln f(\Gamma, t)$$

We can still apply thermodynamical interpretations to this generalized BM.

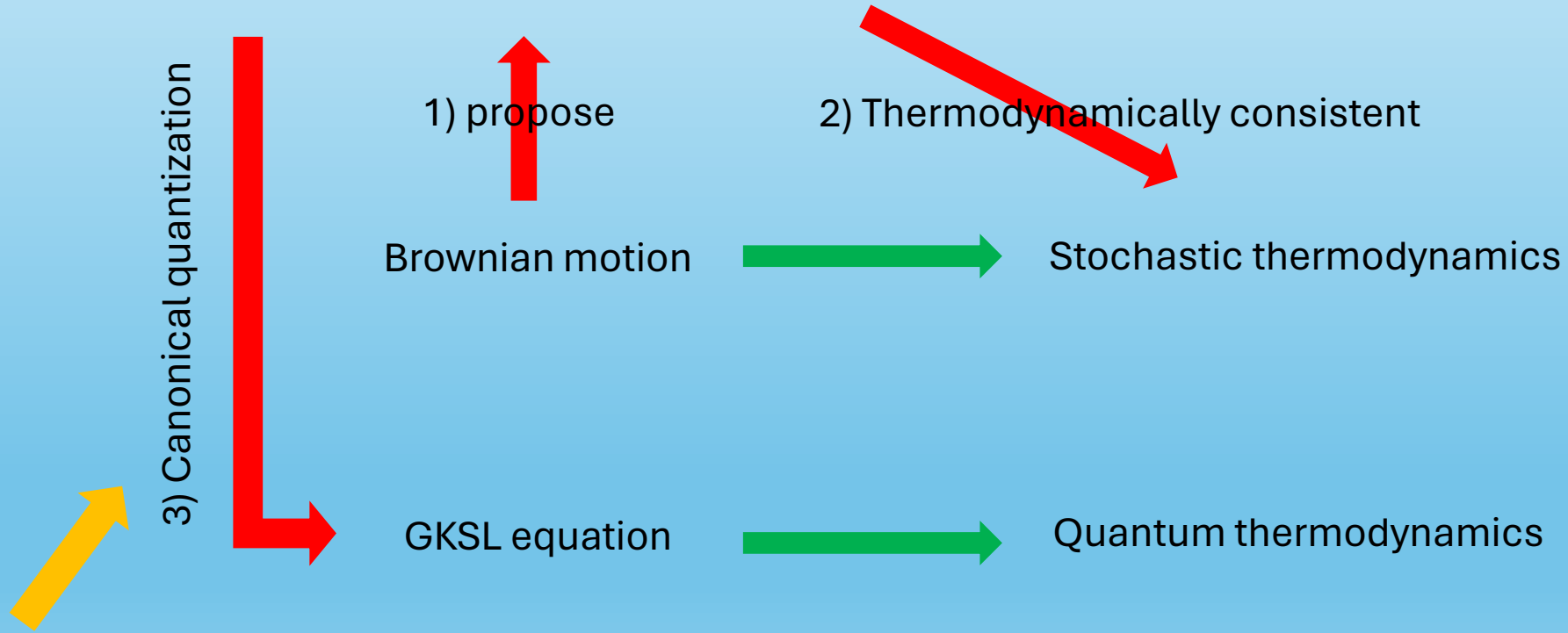


# Canonical quantization and new quantum master equation

Our strategy is.....



Most general theory of Brownian motion (but in flat spacetime)



Apply quantization to phase space distribution.

# Generalized Kramers equation

Koide&Nicacio, Phys. Lett. A494, 129277 (2024).

The phase space distribution

$$f(\Gamma, t) := \int d\Gamma_0 f_0(\Gamma_0) \prod_{\alpha, i} \mathbb{E} \left[ \delta(q_{(i)}^\alpha - q_{(i)t}^\alpha) \delta(p_{(i)}^\alpha - p_{(i)t}^\alpha) \right] \quad \alpha: \text{Components of vectors}$$

The differential equation of the phase space distribution (generalized Kramers equation) is

$$\begin{aligned} \partial_t f = & -\{f, H\}_{PB} + \sum_{i=1}^N \sum_{\alpha=1}^D \frac{\gamma_{p_i}}{\beta_i} \left\{ e^{-\beta_i H} \left\{ e^{\beta_i H} f, q_{(i)}^\alpha \right\}_{PB}, q_{(i)}^\alpha \right\}_{PB} \\ & + \sum_{i=1}^N \sum_{\alpha=1}^D \frac{\gamma_{q_i}}{\beta_i} \left\{ e^{-\beta_i H} \left\{ e^{\beta_i H} f, p_{(i)}^\alpha \right\}_{PB}, p_{(i)}^\alpha \right\}_{PB} \end{aligned}$$

$$\{g, h\}_{PB} = \sum_{i=1}^N \sum_{\alpha=1}^D \left( \frac{\partial g}{\partial q_i^\alpha} \frac{\partial h}{\partial p_i^\alpha} - \frac{\partial g}{\partial p_i^\alpha} \frac{\partial h}{\partial q_i^\alpha} \right)$$

# Canonical quantization

Koide&Nicacio, Phys. Lett. A494, 129277 (2024).

$$\{g, h\}_{PB}$$

$$f(\Gamma, t)$$

$$e^{\pm\beta_i H} f(\Gamma, t)$$



$$-\frac{i}{\hbar} [\hat{g}, \hat{h}] \quad -\frac{i}{\hbar} [\hat{q}, \hat{p}] = 1$$

$$\hat{\rho}(t)$$

$$e^{\pm\beta_i \hat{H}/2} \hat{\rho}(t) e^{\mp\beta_i \hat{H}/2}$$

# New quantum master equation

Koide&Nicacio, Phys. Lett. A494, 129277 (2024).

$$\frac{d\hat{\rho}}{dt} = \frac{i}{\hbar} [\hat{\rho}, \hat{H}] + D[\hat{\rho}]$$

$$D[\hat{\rho}] = -\sum_{i=1}^N \sum_{\alpha=1}^D \frac{\gamma_{p_i}}{\beta_i \hbar^2} \left[ e^{-\beta_i \hat{H}/2} \left[ e^{\beta_i \hat{H}/2} \hat{\rho} e^{\beta_i \hat{H}/2}, \hat{q}_{(i)}^\alpha \right] e^{-\beta_i \hat{H}/2}, \hat{q}_{(i)}^\alpha \right] \\ - \sum_{i=1}^N \sum_{\alpha=1}^D \frac{\gamma_{q_i}}{\beta_i \hbar^2} \left[ e^{-\beta_i \hat{H}/2} \left[ e^{\beta_i \hat{H}/2} \hat{\rho} e^{\beta_i \hat{H}/2}, \hat{p}_{(i)}^\alpha \right] e^{-\beta_i \hat{H}/2}, \hat{p}_{(i)}^\alpha \right]$$

When all particles interact with the same heat bath ( $\beta_1 = \beta_2 = \dots = \beta_N = \beta$ ), the stationary solution is given by thermal equilibrium state,

$$\frac{d\hat{\rho}_{eq}}{dt} = 0 \quad \hat{\rho}_{eq} = \frac{1}{Z} e^{-\beta \hat{H}}$$



# Is this time evolution CPTP?

Koide&Nicacio, Phys. Lett. A494, 129277 (2024).

Let us consider a harmonic oscillator,  $\hat{H} = \frac{\hat{p}^2}{2m} + \frac{m}{2} \omega^2 \hat{q}^2$

violates CPTP

$$D[\hat{\rho}] \longrightarrow -\frac{1}{2\hbar} \sum_{\mu\nu=1}^4 \eta_{\mu\nu} \left\{ \left[ \hat{L}_\mu^\dagger \hat{L}_\nu, \hat{\rho} \right]_+ - 2\hat{L}_\mu \hat{\rho} \hat{L}_\nu^\dagger \right\} \quad \eta_{\mu\nu} = \text{Diag}(1, 1, 1, -1)$$

$$\hat{L}_1 = \Gamma_1 \sqrt{\frac{m\omega}{2\hbar}} \left( \hat{q} + \frac{i}{m\omega} \hat{p} \right)$$

$$\hat{L}_2 = \Gamma_2 \sqrt{\frac{m\omega}{2\hbar}} \left( \hat{q} - \frac{i}{m\omega} \hat{p} \right)$$

$$\hat{L}_3 = \sqrt{\delta} \hat{q}$$

$$\hat{L}_4 = \frac{\sqrt{\delta}}{m\omega} \hat{p}$$

$$\Gamma_i = \sqrt{\frac{(\delta+2)\gamma_p}{2\beta m\omega}} e^{(-1)^{i+1} \beta \hbar \omega / 2}$$

$$\delta = \frac{\gamma_q}{\gamma_p} (m\omega)^2 - 1$$



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$$D[\hat{\rho}] \longrightarrow -\frac{1}{2\hbar} \sum_{\mu\nu=1}^4 \eta_{\mu\nu} \left\{ \left[ \hat{L}_\mu^\dagger \hat{L}_\nu, \hat{\rho} \right]_+ - 2\hat{L}_\mu \hat{\rho} \hat{L}_\nu^\dagger \right\} \quad \eta_{\mu\nu} = \text{Diag}(1, 1, 1, -1)$$

$$\hat{L}_1 = \Gamma_1 \sqrt{\frac{m\omega}{2\hbar}} \left( \hat{q} + \frac{i}{m\omega} \hat{p} \right)$$

$$\hat{L}_2 = \Gamma_2 \sqrt{\frac{m\omega}{2\hbar}} \left( \hat{q} - \frac{i}{m\omega} \hat{p} \right)$$

$$\hat{L}_3 = \sqrt{\delta} \hat{q}$$

$$\hat{L}_4 = \frac{\sqrt{\delta}}{m\omega} \hat{p}$$

$$\Gamma_i = \sqrt{\frac{(\delta+2)\gamma_p}{2\beta m\omega}} e^{(-1)^{i+1} \beta \hbar \omega / 2}$$

$$\delta = \frac{\gamma_q}{\gamma_p} (m\omega)^2 - 1$$

For our master equation to be CPTP, we need to set

$$\delta = 0 \rightarrow \gamma_p = \gamma_q (m\omega)^2$$

# Reproduction of GKSL equation

Koide&Nicacio, Phys. Lett. A494, 129277 (2024).

Our master equation is finally reduced to the GKSL equation with the detailed balance condition

$$\frac{d\hat{\rho}_t}{dt} = \frac{i}{\hbar} [\hat{\rho}_t, \hat{H}] - \frac{1}{2\hbar} \sum_{\mu=\pm} \gamma_{\mu} \left\{ [L_{\mu}^{\dagger} L_{\mu}, \hat{\rho}_t]_{+} - 2L_{\mu} \hat{\rho}_t L_{\mu}^{\dagger} \right\}$$

$$\hat{L}_{+} = \left( \hat{q} + \frac{i}{m\omega} \hat{p} \right) \quad \hat{L}_{-} = \left( \hat{q} - \frac{i}{m\omega} \hat{p} \right) \quad \gamma_{\pm} = e^{\pm\beta\hbar\omega/2} \frac{\gamma_p}{2\hbar\beta} \quad \longrightarrow \quad \frac{\gamma_{-}}{\gamma_{+}} = e^{-\beta\hbar\omega}$$

We can show the laws analogous to the first and second law.



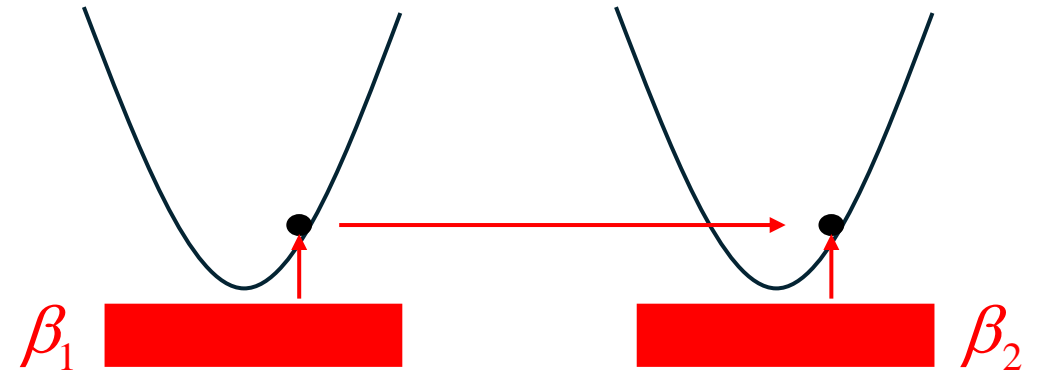
GKSL equation  $\longrightarrow$  CPTP, but **not** always converges toward equilibrium

Our master equation  $\longrightarrow$  **Not** always CPTP, but converges toward equilibrium

# CPTP even in other interactions?

Nicacio&Koide in preparation.

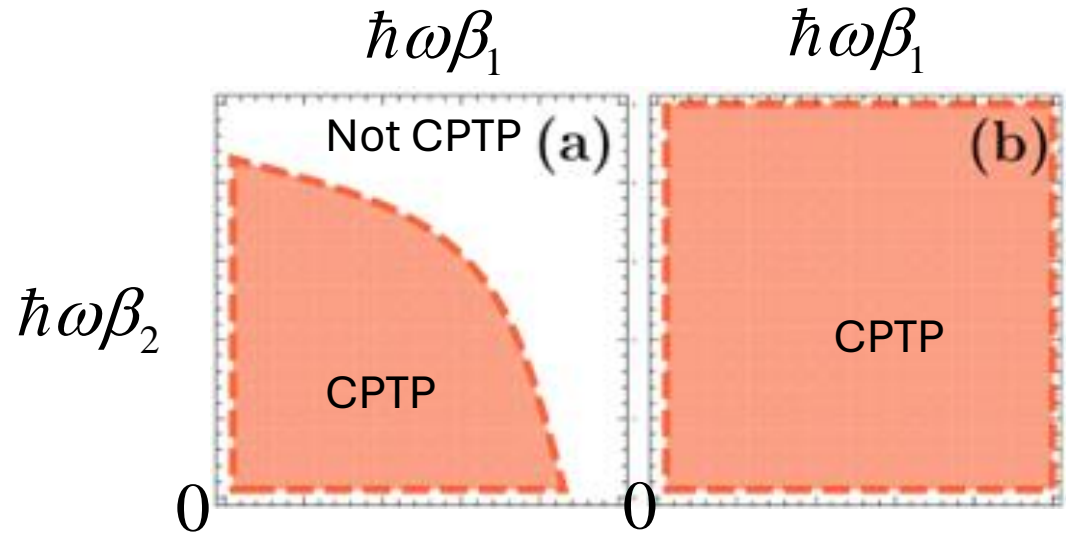
Model for heat conduction (network model)



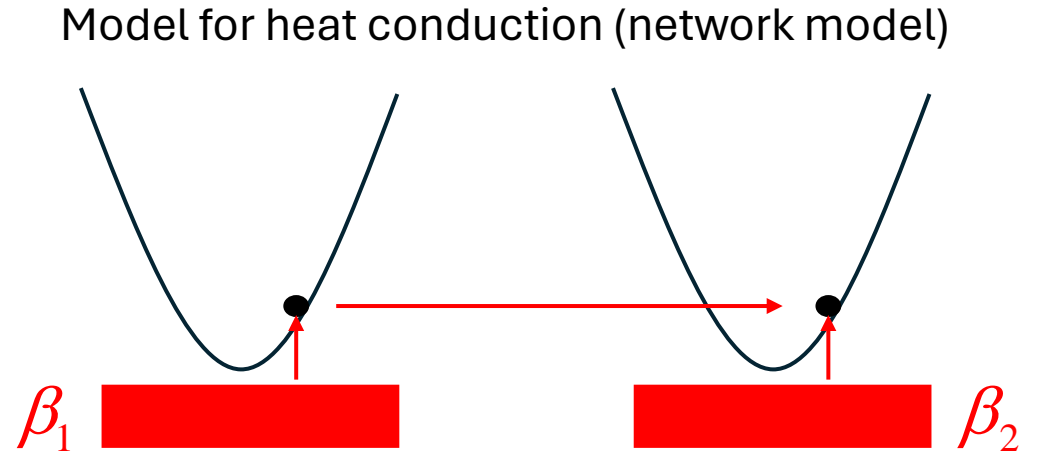
$$\hat{H} = \sum_{i=1}^2 \left( \frac{\hat{p}_i^2}{2m} + \frac{m}{2} \omega^2 \hat{q}_i^2 \right) + \frac{K}{2} (\hat{q}_1 - \hat{q}_2)^2 \quad \begin{array}{l} m = \omega^{-1} \\ K = \omega \end{array}$$

# CPTP even in other interactions?

Nicacio&Koide in preparation.



$$(\gamma_q, \gamma_p) = \left( \frac{1}{4\omega}, \frac{1}{4\omega} \right) \quad \left( \frac{1}{4\omega}, \frac{1}{2\omega} \right)$$



$$\hat{H} = \sum_{i=1}^2 \left( \frac{\hat{p}_i^2}{2m} + \frac{m}{2} \omega^2 \hat{q}_i^2 \right) + \frac{K}{2} (\hat{q}_1 - \hat{q}_2)^2 \quad \begin{array}{l} m = \omega^{-1} \\ K = \omega \end{array}$$

We can find appropriate parameters where our quantum master equation conforms to a CPTP evolution even in the network model.

Other topics

# Stochastic energetics in Field theory

Brownian motion for the scalar field

$$d\phi(x_i, t) = (dt) \frac{\delta H}{\delta \Pi(x_i, t)} - (dt) \gamma_\phi \frac{\delta H}{\delta \phi(x_i, t)} + \sqrt{\frac{2\gamma_\phi}{(dx)\beta}} dB^\phi(x_i, t)$$
$$d\Pi(x_i, t) = -(dt) \frac{\delta H}{\delta \phi(x_i, t)} - (dt) \gamma_\Pi \frac{\delta H}{\delta \Pi(x_i, t)} + \sqrt{\frac{2\gamma_\Pi}{(dx)\beta}} dB^\Pi(x_i, t)$$

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Heat

$$dQ_t^{sf} = \int dx \left( -\gamma_\Pi \frac{\delta H}{\delta \Pi} + \sqrt{\frac{2\gamma_\Pi}{dx\beta}} \frac{dB^\Pi}{dt} \right) \circ d\phi - \int dx \left( -\gamma_\phi \frac{\delta H}{\delta \phi} + \sqrt{\frac{2\gamma_\phi}{dx\beta}} \frac{dB^\phi}{dt} \right) \circ d\Pi$$

Law analogous to the second law

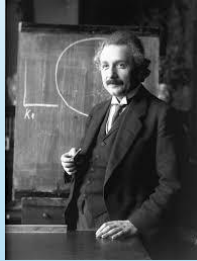
$$\frac{dS_t^{sf}}{dt} - k_B \beta^{-1} \frac{dQ_t^{sf}}{dt} \geq 0$$

Information entropy

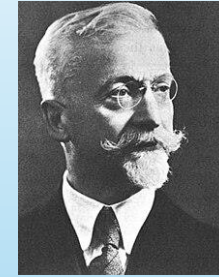
Formulation of quantum thermodynamics is under investigation.....



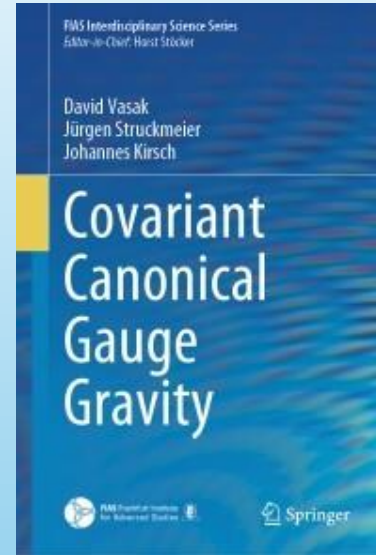
# Thermodynamics in (modified) gravity



Gravity = Curvature + Torsion



Dark matter,  
Dark energy, ...



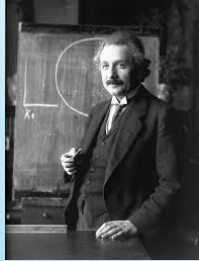
Free diffusion (Brownian motion)

Flat space without torsion

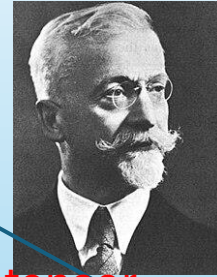
$$\partial_t \rho = v \Delta \rho$$

Curved space with torsion  $\partial_t \rho = -\nabla_i \left( \rho (v g^{ij} K^k_{jk}) \right) + v g^{ij} \nabla_i \nabla_j \rho$   $\nabla_i u^j = \partial_i u^j + \left\{ \begin{matrix} j \\ ki \end{matrix} \right\} u^k + K^j_{ki} u^k$

# Thermodynamics in (modified) gravity

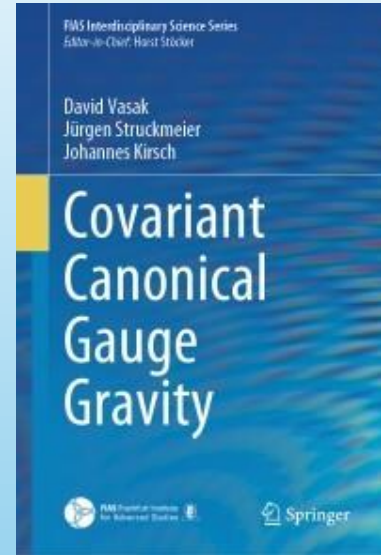


Gravity = Curvature + Torsion



Dark matter,  
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Contorsion tensor



Free diffusion (Brownian motion)

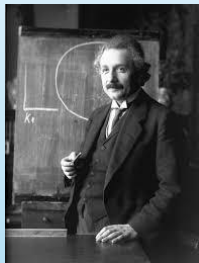
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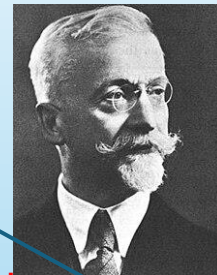
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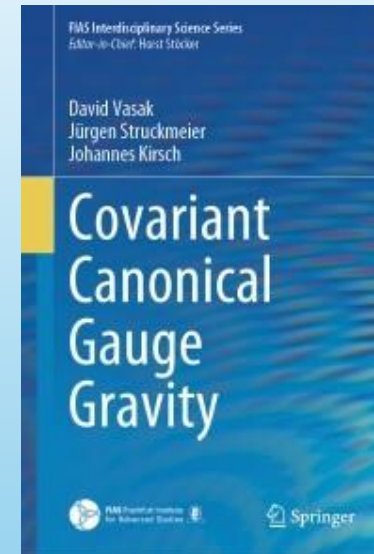


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1. Torsion (vierbein) is considered exclusively associated with the spin degrees of freedom. However, it is also noteworthy that we **cannot avoid introducing vierbein** to construct Brownian motion in generalized coordinates.
2. One way to study thermodynamical behavior in curved (spacetime) geometry with or without torsion, is to consider Brownian motion.
3. Is it possible to construct stochastic and quantum thermodynamics in this case?

No torsion

# Thermodynamics and Universe

1) Expansion of universe

$$T_{rad} \propto R^{-1}(t) \quad s_{rad} R^3 \sim R^3 T^3 = const$$

$$T_{mat} \propto R^{-2}(t) \quad S_{mat} \sim \ln(R^3 T^{3/2}) = const$$

Different from the expansion  
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What is the non-equilibrium effect?

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2) Self gravitating system



From the virial theorem,

$$\Delta T \sim \Delta E_{kin} = -\Delta E_{tot}$$

Can this be modeled using Brownian motion?  
Fluctuation effect in self-gravitating systems?

Sakagami&Taruya,Contium Mech. Thermodyn. 16, 279 (2004)

Chavanis et al.,Phys.Rev.E66, 036105 (2002)

- Heat capacity is negative
- Thermo. inhomogeneity is enhanced.
- gravitational contraction (collapse)

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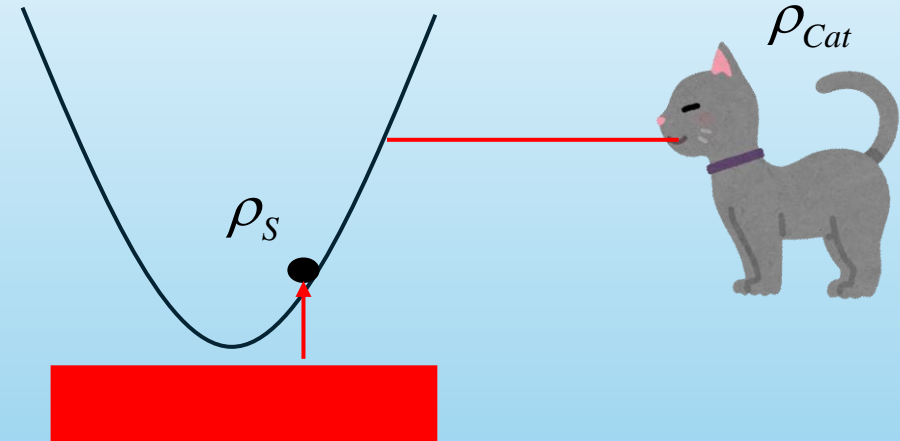
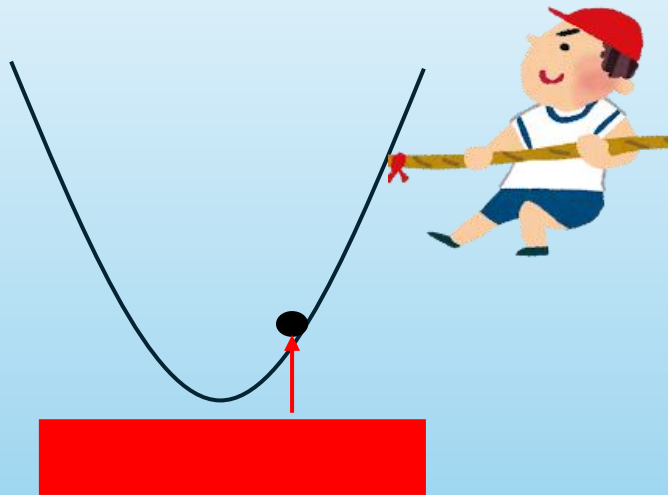
3) Brownian motion and Black Hole populations in globular clusters

Roupas, A&A 646 A20 (2021).

Chavanis&Mannella, Eur. Phys. J. B 78, 139 (2010).

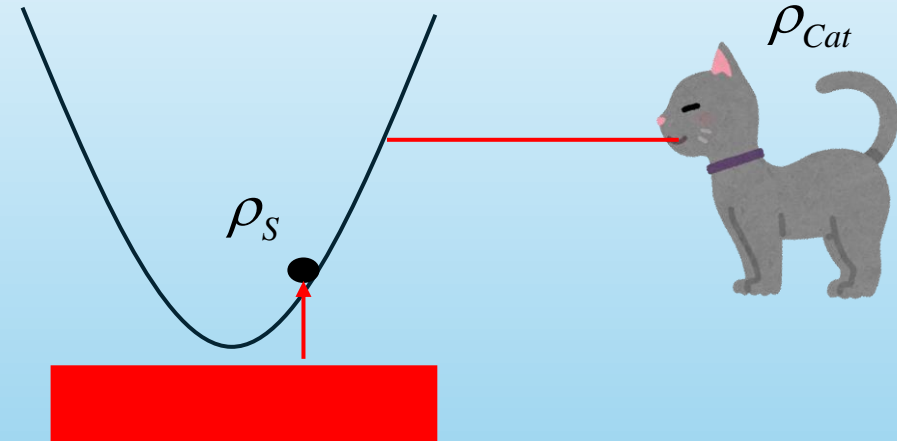
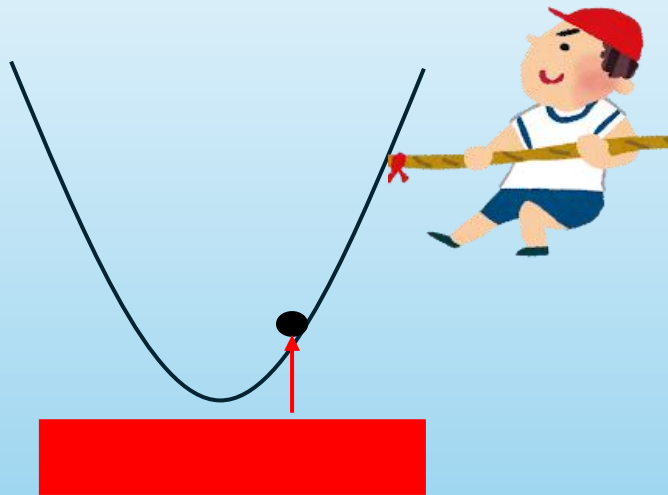
4) Black hole and entropy

# Thermo. work and Q. measurement



$$\rho_S \otimes \rho_{Cat} \Rightarrow U \rho_S \otimes \rho_{Cat} U^\dagger$$

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$$\rho_S \otimes \rho_{Cat} \Rightarrow U \rho_S \otimes \rho_{Cat} U^\dagger$$

1. What is the relation between classical and quantum optimal controls for external perturbation?

Sekimoto&Sasa, J. Phys. Soc. Japan, 66, 3658 (2001)

Schmiedl&Seifert, Phys. Rev. Lett. 98, 108301 (2007)

Koide, J. Phys. A50, 325001 (2017, **Berry's phase**)

?????

2. What is the role of quantum measurement (Maxwell Deamon) and what is its classical limit?

Classical limit?



Scully et al., Science 299, 862 (2003).

Kammerlander&Anders, Scientific reports 6, 22174 (2016).

Elouard et al., Phys. Rev. Lett. 118, 260603 (2017).



# TUR (thermodynamical uncertainty relations)

Barato&Seifert, PRL114, 158101 (2015) Hasagawa&Van Vu, PRL123,20001 (2017)

1. A particle trajectory in a nonequilibrium steady state  $\Gamma$
2. An observable  $\phi$  satisfying  $\phi(\Gamma^{TR}) = -\phi(\Gamma)$ 

Time reversed trajectory
3. The probability distribution of the trajectories  $P(s, \phi)$
4. The fluctuation theorem  $\frac{P(s, \phi)}{P(-s, -\phi)} = e^s$ 

Entropy production

TUR

$$\frac{\langle (\phi - \langle \phi \rangle)^2 \rangle}{\langle \phi \rangle^2} \geq \frac{2}{e^{\sigma\tau} - 1}$$

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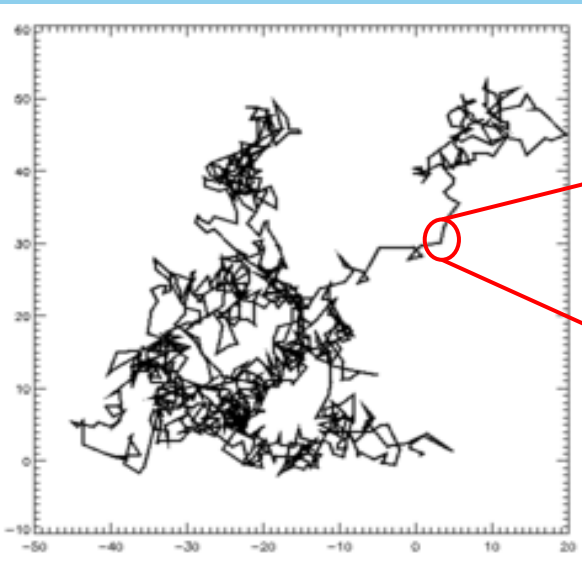
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Is there TUR induced by non-differentiability?

de Matos et al., WATER12,3263 (2020)

# Furthermore.....

$$\eta_{Car} = 1 - \frac{T_c}{T_h} \quad \eta_{CA} = 1 - \sqrt{\frac{T_c}{T_h}}$$

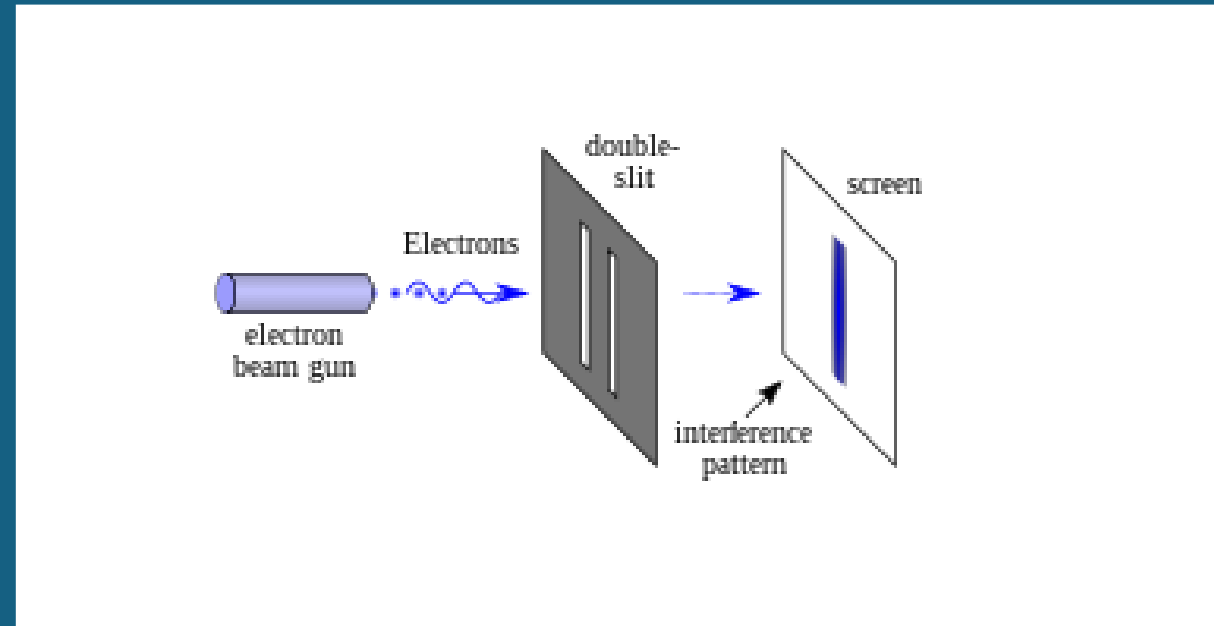
1. Efficiency of heat engine in finite-time operations (Cruzon-Ahlborn efficiency, Am. J. Phys. 43, 22, 1975)
2. How can we take the thermodynamical limit in stochastic and quantum thermodynamics?
3. Finite chemical potential? (Neidig, et al., arXiv:2308.07659)
4. Relation between the GKSL equation and the quantization of damped HO
5. Entanglement in classical stochastic mechanics (Reciprocal process)?  
(Schrödinger, Akad.Phys.Math.Klasse 1, 144 (1931), Koide, J. Phys. Commun. 2, 021001 (2018))

# Concluding remarks

1. We develop a general model of Brownian motion in flat spacetime.
2. We can define heat and entropy so that the behaviors of the model are consistent with thermodynamics.
3. A quantum master equation is derived from the model by applying the canonical quantization.
4. Given a system Hamiltonian, the form of our quantum master equation is determined except for a few parameters. (**Advantage 1**)
5. Regardless of the choice of the system Hamiltonian, the classical limit of the quantum master equation always describes a thermal relaxation process. (**Advantage 2**)
6. The derived master equation does not always satisfy the CPTP condition but, in several applications, we can find the parameters where the quantum master equation becomes a GKSL equation.
7. Our approach enables us formulate **a unified framework** of stochastic and quantum thermodynamics.



# Let us consider the double-slit experiment.

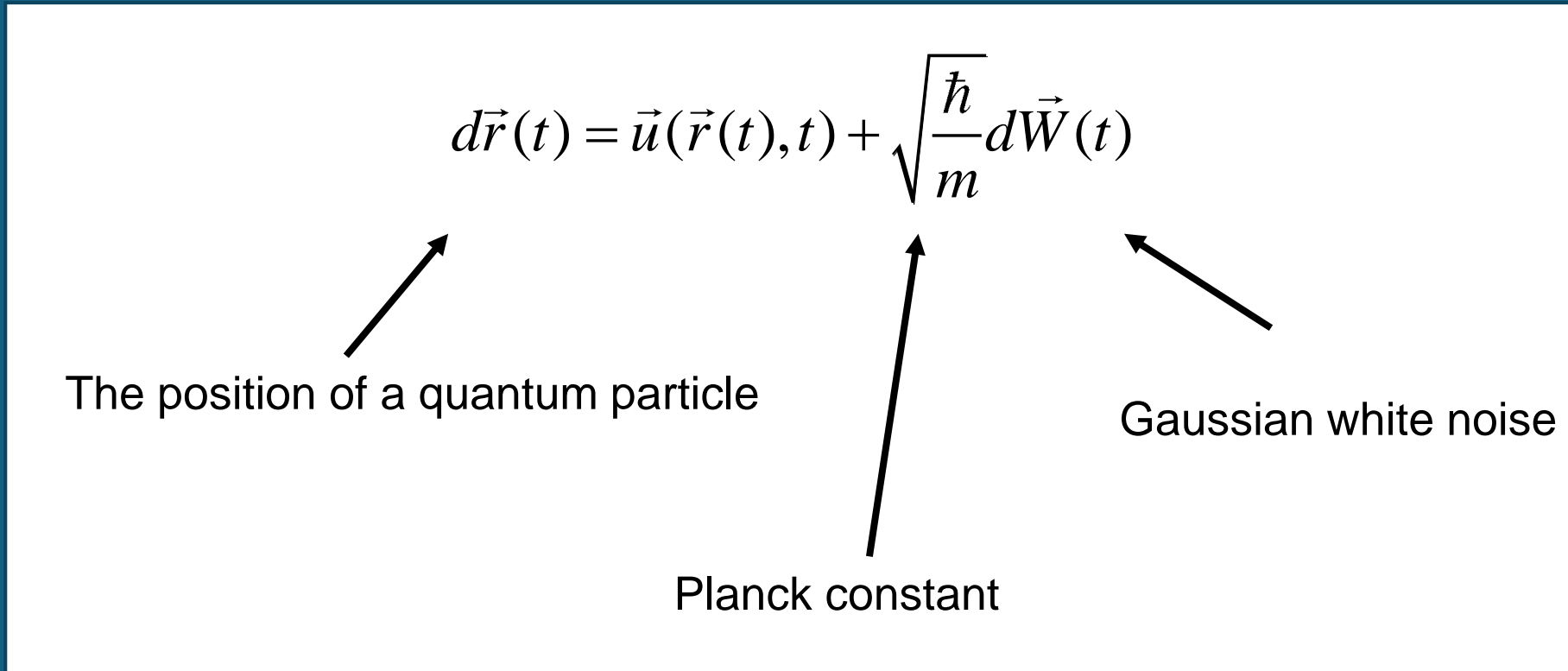


1. The wave function of this system  $\phi(\vec{x}, t)$
2. Then we define a vector field by

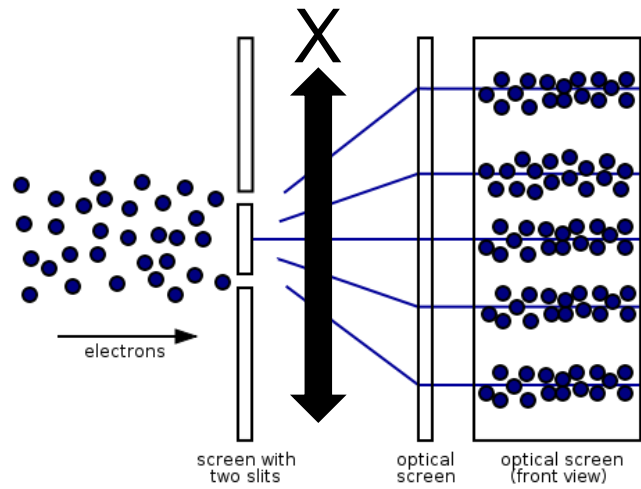
$$\vec{u}(\vec{x}, t) = \frac{\hbar}{m} \nabla \left\{ \text{Re} [\ln \phi(\vec{x}, t)] + \text{Im} [\ln \phi(\vec{x}, t)] \right\}$$

The probability distribution is reproduced by the frequency distribution of Brownian motion,

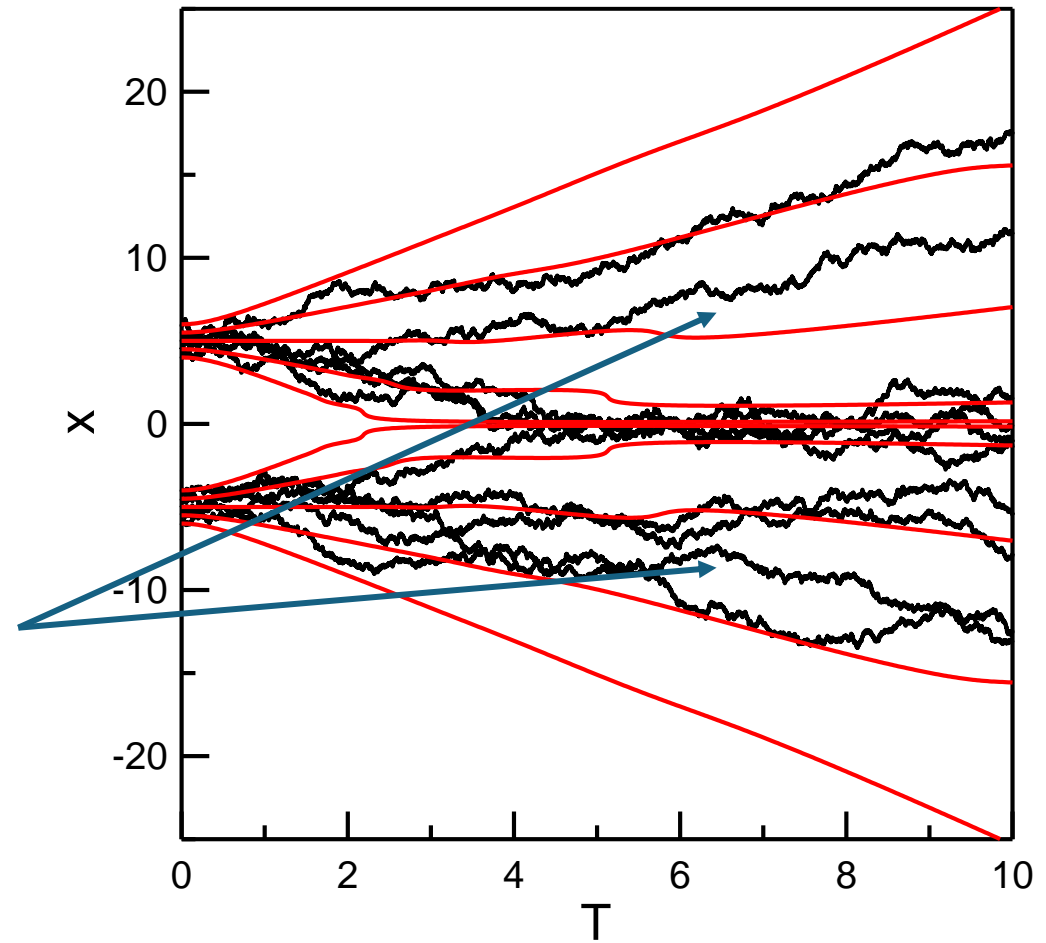
Nelson, PR150,1079('66)



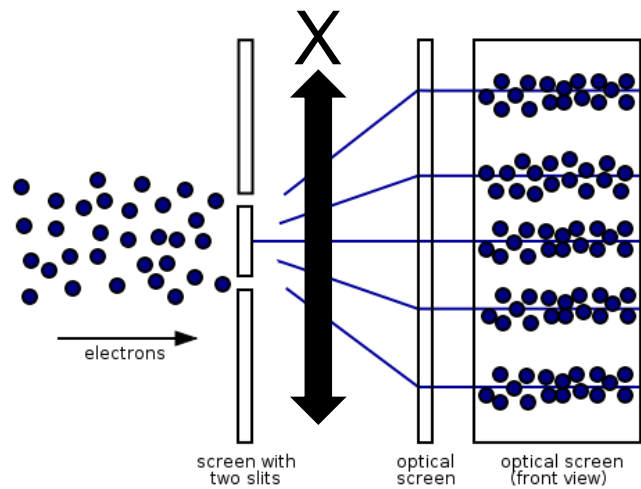
$$\vec{u}(\vec{x}, t) = \frac{\hbar}{m} \nabla \left\{ \text{Re}[\ln \phi(\vec{x}, t)] + \text{Im}[\ln \phi(\vec{x}, t)] \right\}$$



Brownian motion

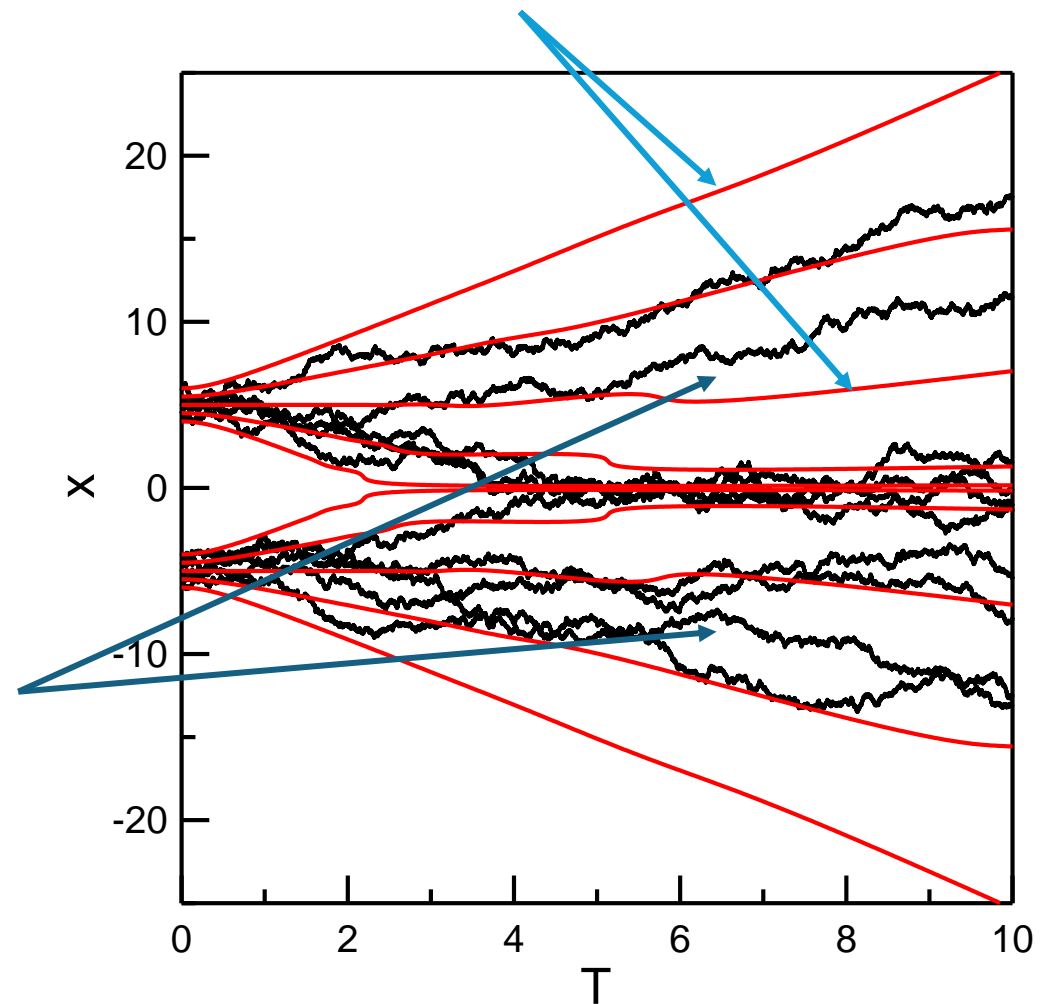






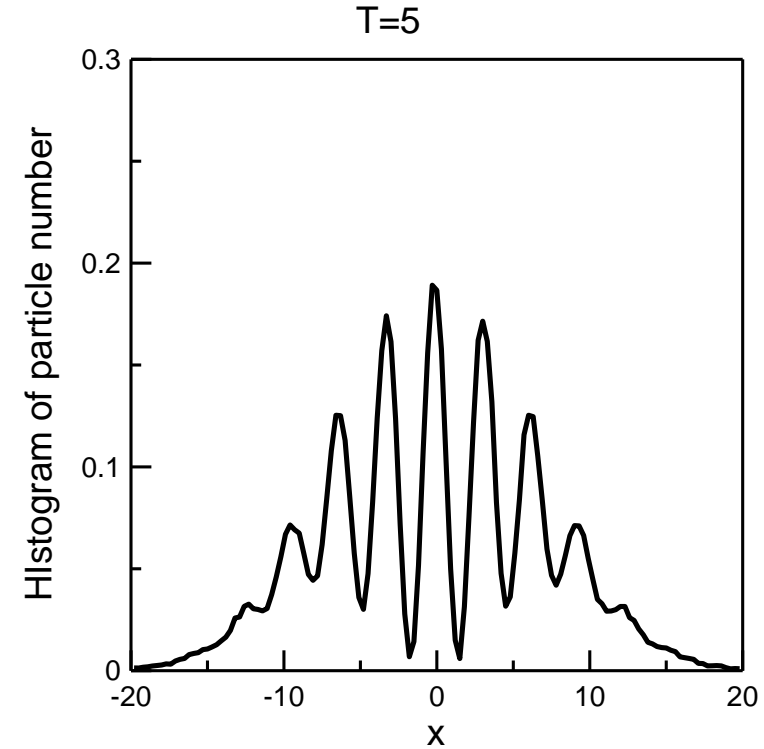
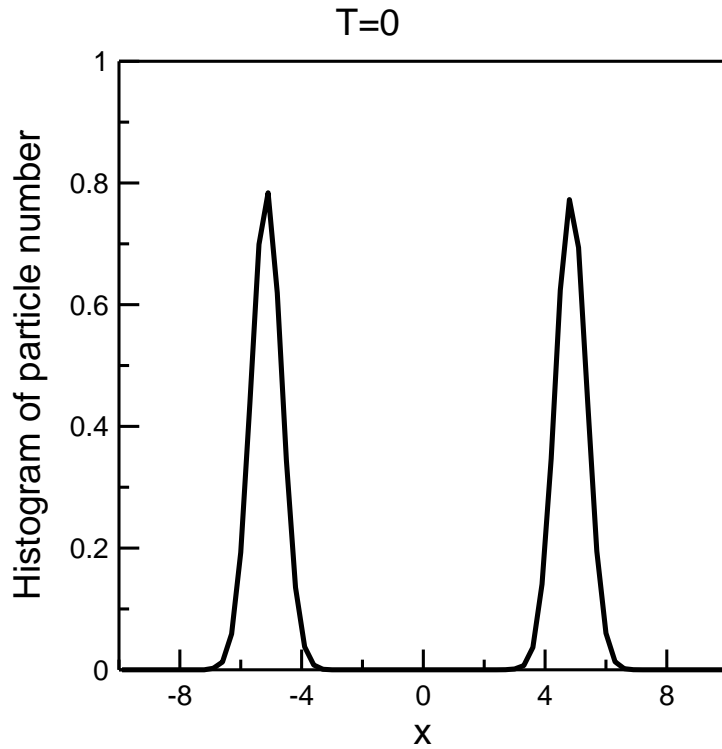
Brownian motion

Bohmian trajectory



# 200000 Brownian particles

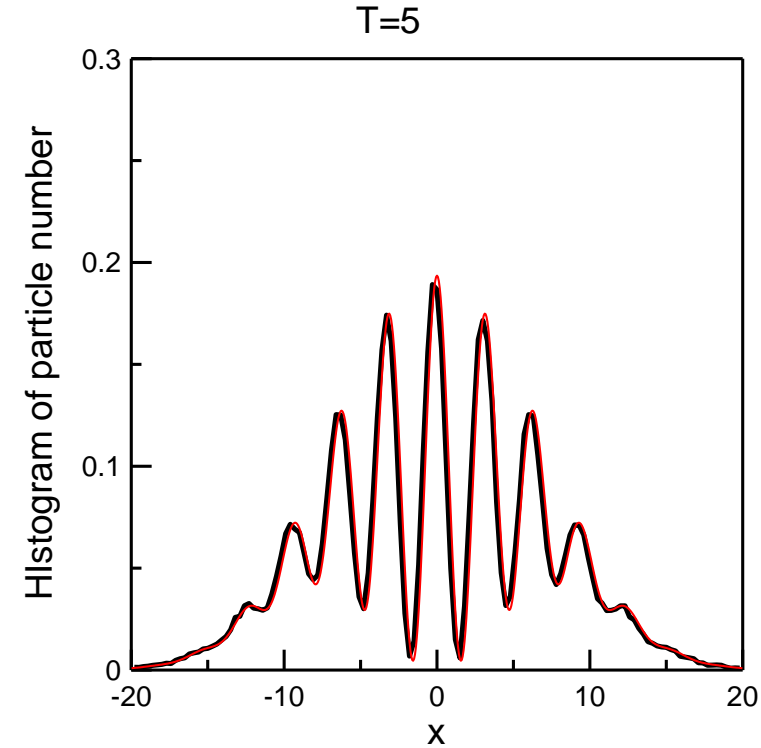
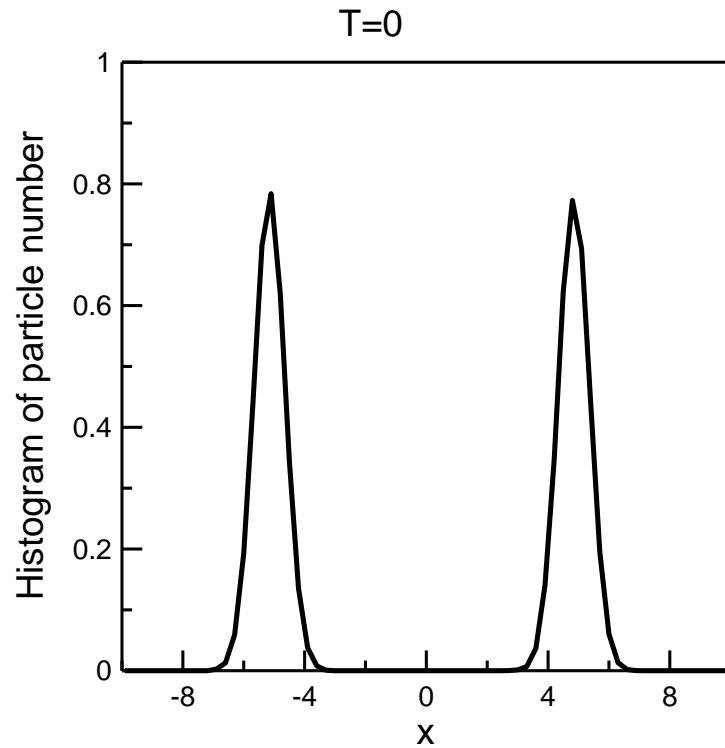
5000 time steps



Simulation of  $d\vec{r}(t) = \vec{u}(\vec{r}(t), t)dt + \sqrt{\frac{\hbar}{m}}d\vec{W}(t)$

# 200000 Brownian particles

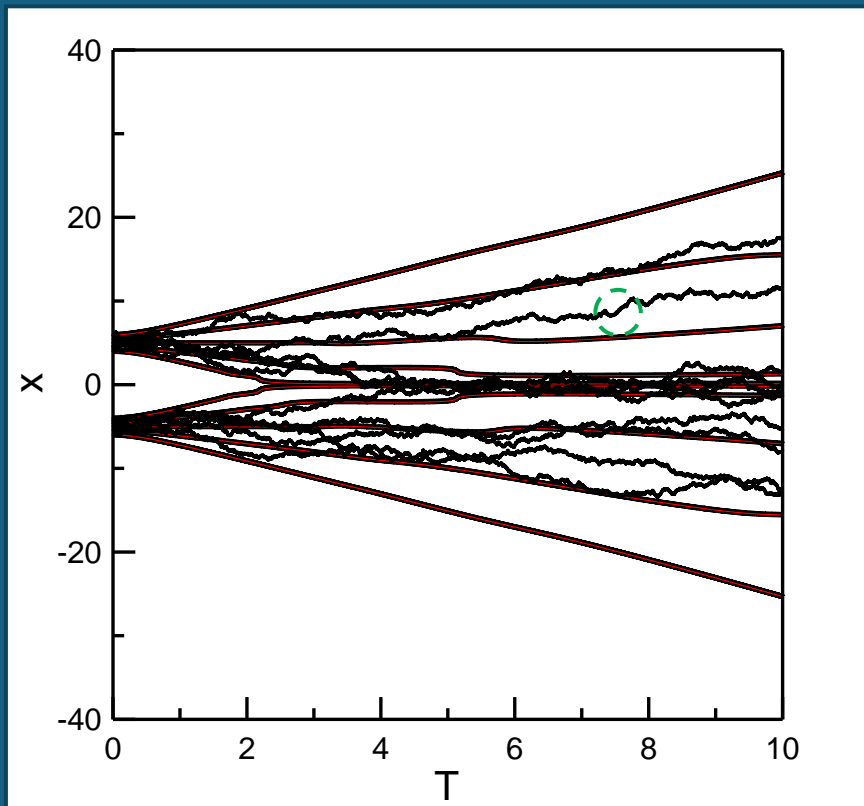
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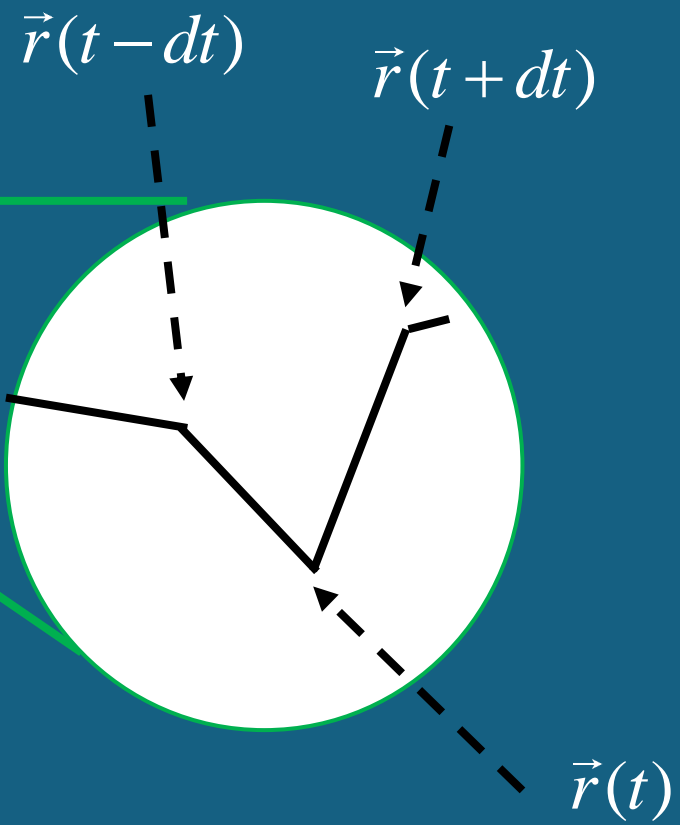
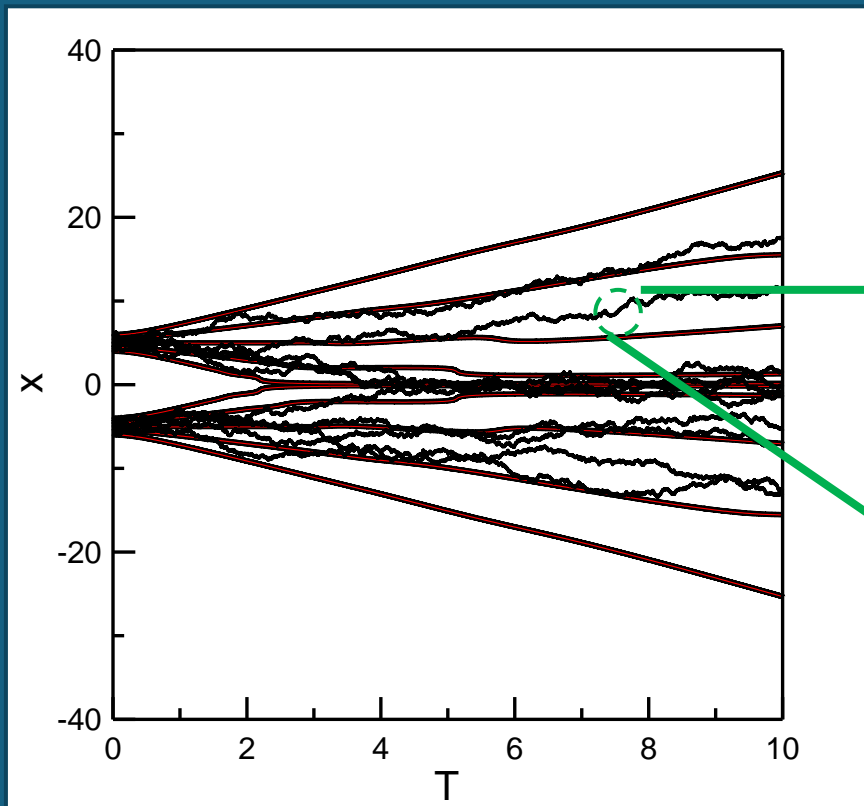


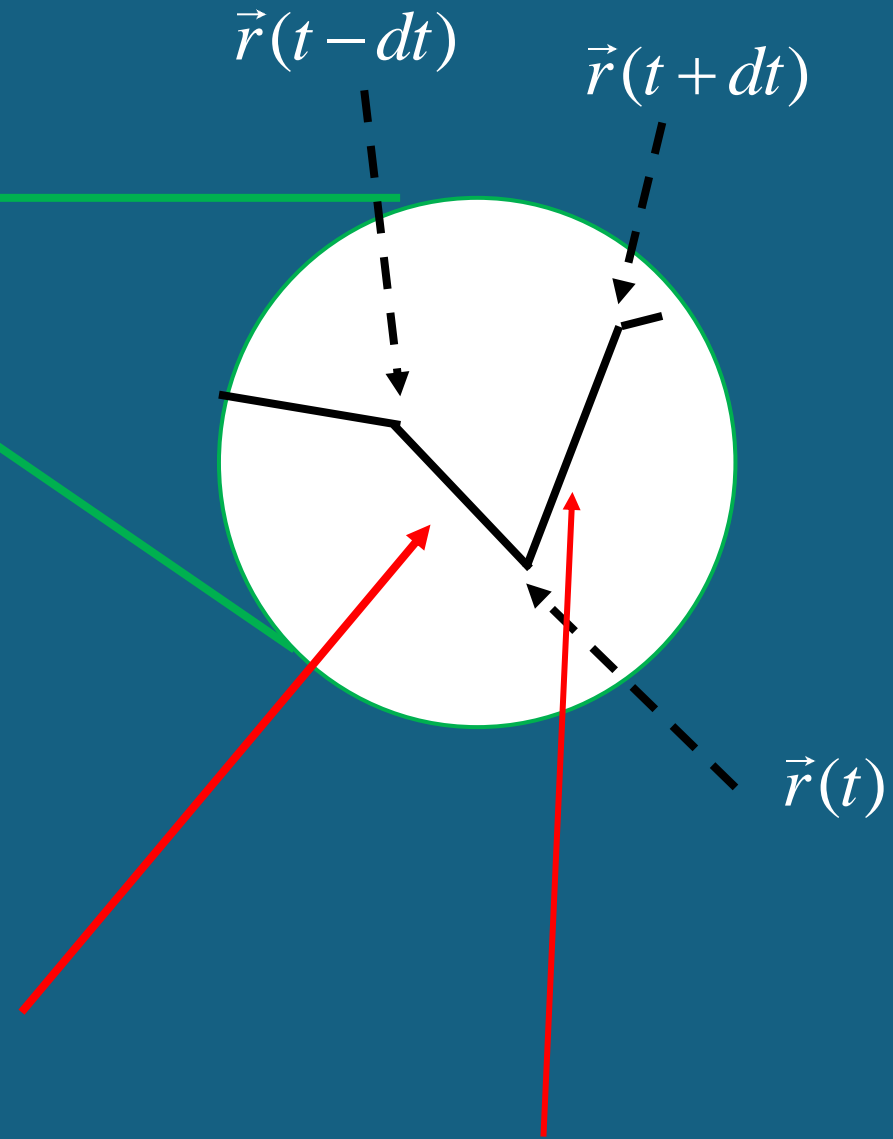
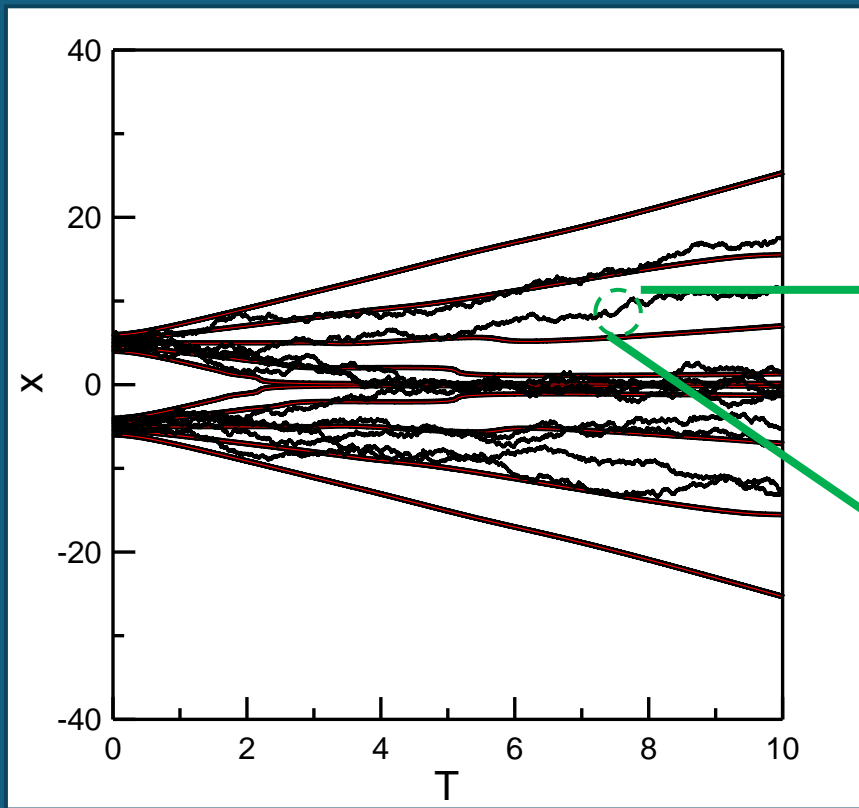
Simulation of  $d\vec{r}(t) = \vec{u}(\vec{r}(t), t)dt + \sqrt{\frac{\hbar}{m}}d\vec{W}(t)$



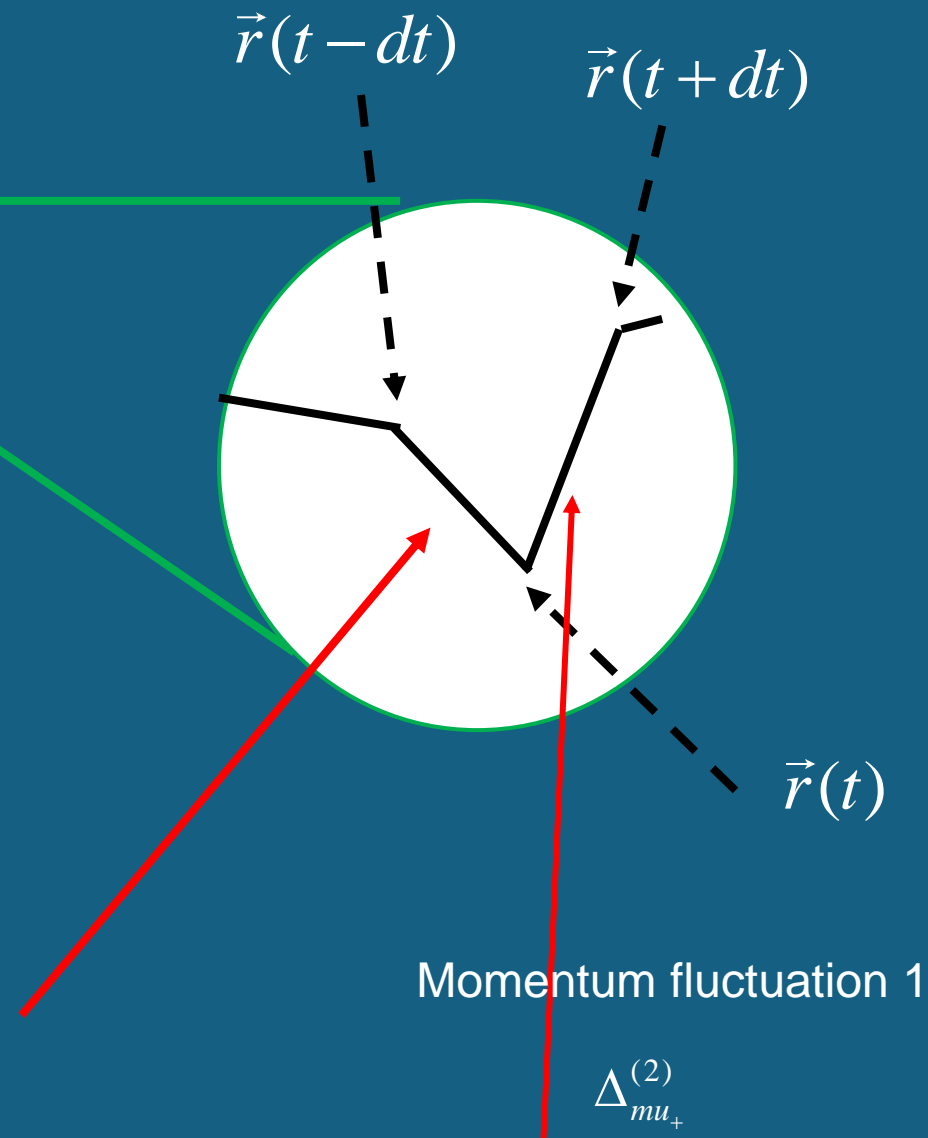
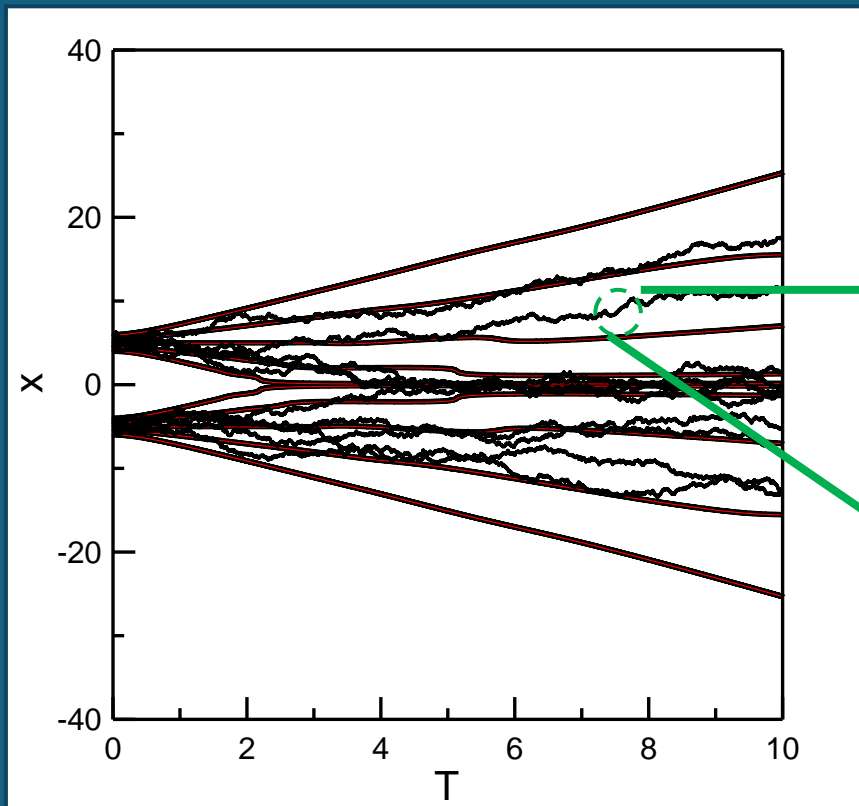
Exact result



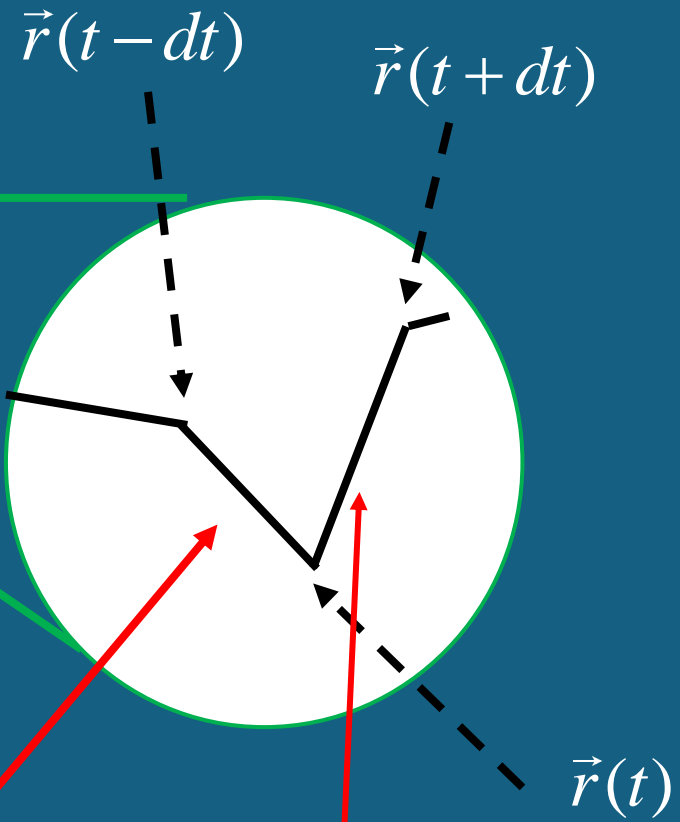
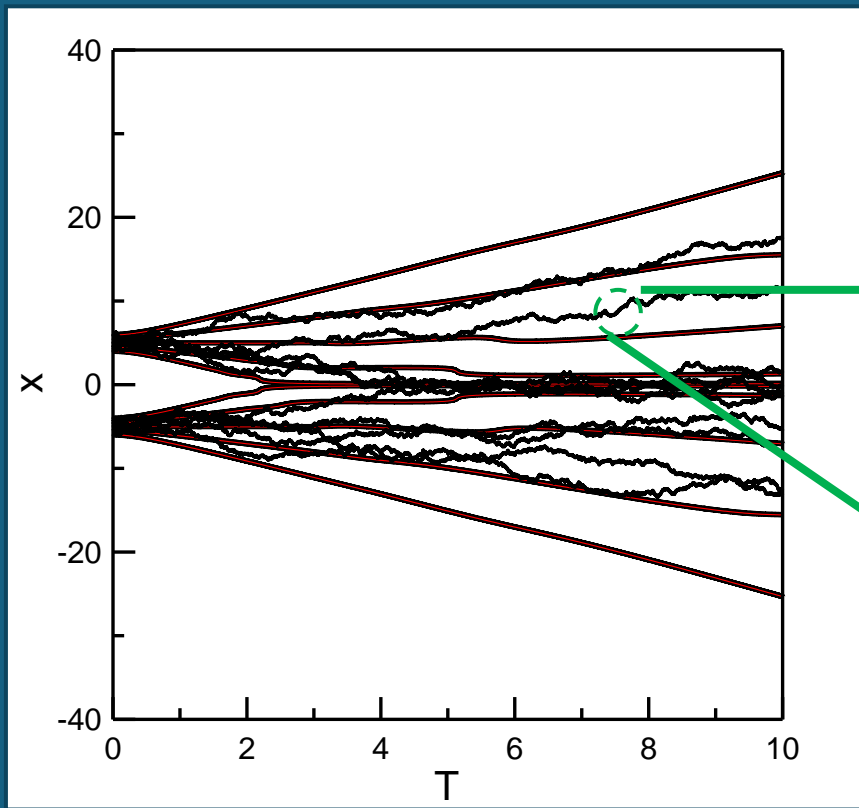




We can find, at least, two velocities.



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Momentum fluctuation 2

$$\Delta_{mu_-}^{(2)}$$

Momentum fluctuation 1

$$\Delta_{mu_+}^{(2)}$$

We can find, at least, two velocities.



Suppose that the fluctuation of the observed momentum is given by the average of the two fluctuations.

$$\Delta_p^{(2)} = \frac{\Delta_{mu_+}^{(2)} + \Delta_{mu_-}^{(2)}}{2}$$



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The Kennard inequality in quantum mechanics



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The Kennard inequality in quantum mechanics

The uncertainty relation can be induced from the non-differentiability of observables.

